

## MATHRMATICS OR SBISMDLOGY

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## 1. GREEN'S FUNCTIONS

In this chapter we discuss those properties of scalar Green's functions which make them nseful as sources by themselves and as adjunct elemental sources in finding field distributions for more complicated problems involving surfaces and volumes. We begin with the Green's functions for the wave equation in $n=1,2,3$ spatial dimensions and 1 temporal dimension. Osing different boundary conditions we show that in ( $n, 1$ ) dimensions there are five possible Green's functions and, using their interrelationships, only three independent ones. Each one has various uses depending on the problem at hand. Integration over time yields the corresponding Green's functions for the Helmholtz equation again in $n=1,2,3$ spatial dimensions. We also treat separately, and briefly, the causal Green's fanction for the parabolic wave equation.

Fourier transform representations of the se Helmholtz Green's functions are of ten usefal, as are additional integral representations formed by integration over one or two of the Fourier transform variables. This 1 eads to integral representations associated with the names Weyl, Sommerfeld, and Meyrich. These are all spectral representations of some kind, the Weyl representation being a two-dimensional integral over the transverse wavenumber components, the Sommerfeld representation a one-dimensional representation over the (radial) transverse wavenumber, and the Weyrich representation a one-dimensional representation over the vertical wavenumber component. Both the latter use cylindrical symmetry properties. Plane wave spectral decompositions are also treated as are their interrelations with the above representations. An example is discussed of the use of the representations in the half-plane.

For complicated geometries the most straightforward approach to solving
boundary value problems is to use integral equations. To develop surface integral equations using Green's functions as elemental sources, or their derivatives as dipole sources, it is necessary to know their analytic properties. In particular spatial singularities must be treated in such a way that the resulting integral equations may be solved using classical techniques. This is called regularization, and we demonstrate the regularization of the first and second vector derivatives of the Helmholtz Green's function.

Finally we discuss the Green's function in one-dimension using conventional methods here generalized to inhomogeneous media. We treat a general method of finding profiles for which the one-dimensional Helmholtz equation is solvable in terms of known classical functions.

## 1. 1 GREEN'S FUNCTION FOR THR DAVE EQDATION

### 1.1.1 (3,1)-DIMENSIONS

The Green's function is defined as

$$
\begin{equation*}
G^{(3,1)}\left(\underset{\sim}{x}, x_{\sim}^{\prime} ; t, t^{\prime}\right) \tag{1.1}
\end{equation*}
$$

in three spatial and one temporal dimensions. It satisfies the wave equation given by (c is the wave speed)

$$
\begin{equation*}
\left[\nabla_{x}^{2}-c^{-2} \partial_{t}^{2}\right] G^{(3,1)}\left(\underset{\sim}{x}, x^{\prime} ; t, t^{\prime}\right)=-\delta\left(\underset{\sim}{x}-x_{\sim}^{\prime}\right) \delta\left(t-t^{\prime}\right) \tag{1.2}
\end{equation*}
$$

where the Laplacian is defined by

$$
\begin{equation*}
\nabla_{x}^{2}=\partial_{x}^{2}+\partial_{y}^{2}+\partial_{z}^{2} \tag{1.3}
\end{equation*}
$$

It is convenient to do many of the manipulations in four-vector notation

$$
\begin{equation*}
x=\left(\underset{\sim}{x}, x_{0}\right) \quad x_{0}=c t, \tag{1.4}
\end{equation*}
$$

with the scalar product defined by

$$
x \cdot x=\underset{\sim}{x} \cdot x-x_{0}^{2}\left[\begin{array}{ll}
>0 & \text { space-1ike }  \tag{1.5}\\
=0 & \text { 1ight-1ike } \\
<0 & \text { time-1ike }
\end{array}\right.
$$

Osing this notation the Green's function satisfies

$$
\begin{equation*}
\square G^{(3,1)}\left(x, x^{\prime}\right)=-\delta\left(x-x^{\prime}\right)=-\delta\left(\underset{\sim}{x}-x^{\prime}\right) \delta\left(x_{0}-x_{0}^{\prime}\right), \tag{1.6}
\end{equation*}
$$

where we have defined the $d^{\prime}$ Alembertian operator

$$
\begin{equation*}
\square=\nabla_{x}^{2}-c^{-2} \partial_{t}^{2}=\nabla_{x}^{2}-\partial_{0}^{2} . \tag{1.7}
\end{equation*}
$$

Since we have $\delta\left(x_{0}-x_{0}^{\prime}\right)=c^{-1} \delta\left(t-t^{\prime}\right)$ we see from (1.2) and (1.6) that our Green's functions are related by

$$
\begin{equation*}
G^{(3,1)}\left(\underset{\sim}{x}, x^{\prime}, t, t^{\prime}\right)=c G^{(3,1)}\left(x, x^{\prime}\right) \tag{1.8}
\end{equation*}
$$

Since the delta function source term is a function of the difference between the space-time source point ( $x^{\prime}$ ) and receiver point ( $x$ ), and the coefficients of the differential operators in the diembertian are constants (homogeneous medium), the Green's function only depends functionally on the difference $x-x^{\prime}$. For convenience we wite it as a
function of this difference, and introduce the Fourier transform in the difference argument

$$
\begin{equation*}
G^{(3,1)}(x)=(2 \pi)^{-4} \iiint \int \exp (i k \cdot x) \tilde{G}^{(3,1)}(k) d^{4} k, \tag{1.9}
\end{equation*}
$$

with notation ( $\omega$ is circular frequency)

$$
\begin{align*}
& k \cdot x=\underset{\sim}{k} \cdot \underset{\sim}{x}-k_{0} x_{0}, k_{0}=\omega / c \\
& d^{4} k=d \underset{\sim}{k} d k_{0} . \tag{1.10}
\end{align*}
$$

Applying the d'Alembertian operator to (1.9) we see that (1.6) is satisfied provided

$$
\begin{equation*}
\tilde{G}^{(3,1)}(k)=\left(k^{2}-k_{0}^{2}\right)^{-1}=k^{-2} \tag{1.11}
\end{equation*}
$$

He note that the four-dimensional delta function is written as

$$
\begin{equation*}
\delta(x)=(2 \pi)^{-4} \iiint \int e x p(i k \cdot x) d^{4} k \tag{1.12}
\end{equation*}
$$

Dsing (1.11) and defining $\omega_{k}=|k|$ we can write (1.9) as

$$
\begin{equation*}
G^{(3,1)}(x)=-\frac{1}{(2 \pi)^{4}} \iiint \int \frac{e x p(i k \cdot x)}{\left(k_{0}-\omega_{k}\right)\left(k_{0}+\omega_{k}\right)} d^{4} k \tag{1.13}
\end{equation*}
$$

To evaluate (1.13) we must first evaluate the $k_{0}$ integral

$$
\begin{equation*}
I=\int_{-\infty}^{\infty} \frac{\exp \left(-i k_{0} x_{0}\right)}{\left(k_{0}-\omega_{k}\right)\left(k_{0}+\omega_{k}\right)} d k_{0} \tag{1.14}
\end{equation*}
$$

To do this we must define how to treat the pole terms at $k_{0}={ }^{ \pm} \omega_{k}$ in the integrand. There are two equivalent ways to do this. The first is
(a) Fix the poles - offset the contour

There are five ways to do this illustrated below:
$1 . \rightarrow$ R $\rightarrow$ R: retarded
Here the integration contour is written as semicircles above the two poles at $\pm_{\omega_{k}}$. It is called the retarded contour for reasons which will be clear later.
$2 . \rightarrow \underset{\rightarrow}{ } \rightarrow$ a: advanced

Here the integration contour is written as semicirlces below the two poles at $\pm_{\omega_{k}}$.
3. $\rightarrow \times \longmapsto \longrightarrow \mid \times P$ principal value

Here the poles are evaluated using the Cauchy principal value definition of the integral.
4.


Here the integration contour is written using semicircles, one below the pole at $-\omega_{k}$ and one above the pole at $+\omega_{k}$. The name arises from $F$. Dyson in his work on quantum field the ory.


Here the semicircles are reversed from the Dyson contour.

The second method is to:
(b) Fix the contour - offset the poles

For this method we fix the contour along the real $k_{0}$ aris from $-\infty$ to $\infty$, and shift the poles. The above five become

1. $\rightarrow X$ R: retarded
2. $\mathrm{x} \quad \mathrm{x}$ : advanced
3. $\longrightarrow x \longrightarrow x \longrightarrow p:$ principal value
4. 



D: Dyson
5.


C: causal

Each method makes clear that we will treat (1.14) as an integral in the complex plane. We use method (b), i.e. we shift the poles by an amount ie and consider the results in the 1 imit as $\varepsilon \rightarrow 0$. That is we shift them into
the imaginary part of the complex $k_{0}$ plane. We thus have

$$
\begin{equation*}
\left(k_{0}-\omega_{k}\right)\left(k_{0}+\omega_{k}\right) \rightarrow\left[k_{0}-\left(\omega_{k}+i \alpha \varepsilon\right)\right]\left[k_{0}+\omega_{k}-i \beta \varepsilon\right] \tag{1.15}
\end{equation*}
$$

where $\alpha, \beta= \pm 1,0$ depending on the shift. For the five cases we have that

R: $\quad \alpha=\beta=-1$
$A: \quad \alpha=\beta=+1$
P: $\quad \alpha=\beta=0$
D: $\quad \alpha=-1, \quad \beta=1$
C: $\quad \alpha=1, \beta=-1$.

To evaluate the poles, use the Dirac-Plemelj relations for distributions (we assume the 1 imit $\varepsilon \rightarrow 0$ )

$$
\begin{equation*}
\frac{1}{y \pm i \varepsilon}=P \frac{1}{y} \mp \pi i \delta(y) \tag{1.17}
\end{equation*}
$$

That is, we express the poles at $y \pm$ ie in terms of principal value (P) distributions and half-residue terms from the semicircles. (Ref. 1.4, p. 476.) We thas have for one pole

$$
\begin{equation*}
\frac{1}{k_{0}-\left[\omega_{k}+i a \varepsilon\right]}=P \frac{1}{k_{0}-\omega_{k}}+a \pi i \delta\left(k_{0}-\omega_{k}\right) \tag{1.18}
\end{equation*}
$$

and for the product of two poles

$$
\begin{align*}
& \frac{1}{\left(k_{0}-\omega_{k}-i \alpha \varepsilon\right)\left(k_{0}+\omega_{k}-i \beta_{\varepsilon}\right)} \\
& \quad=P \frac{1}{k_{0}^{2}-\omega_{k}^{2}}+\frac{\pi i}{2 \omega_{k}}\left[\alpha \delta\left(k_{0}-\omega_{k}\right)-\beta \delta\left(k_{0}+\omega_{k}\right)\right] \tag{1.19}
\end{align*}
$$

Substituting (1.19) into (1.13) we can thoswrite all of our examples in terms of two integrals as

$$
\begin{equation*}
G^{(3,1)}(x)=-(2 \pi)^{-4}\left[I_{1}(x)+(\pi i / 2)\left[\alpha I_{2}\left(\underset{\sim}{x},-x_{0}\right)-\beta I_{2}\left(\underset{\sim}{x}, x_{0}\right)\right]\right] \tag{1.20}
\end{equation*}
$$

where the integrals are defined by

$$
\begin{equation*}
I_{1}(x)=p \iiint \int \frac{e x p(i k \cdot x)}{k_{0}^{2}-\omega_{k}^{2}} d_{k}^{4} \tag{1.21}
\end{equation*}
$$

and

$$
\begin{equation*}
I_{2}\left(\underset{\sim}{x}, x_{0}\right)=\iiint \frac{\mathrm{exp}\left[i\left(\underset{\sim}{k} \cdot \underset{\sim}{x}+\omega_{k} x_{0}\right)\right]}{\omega_{k}} d \underset{\sim}{k} \tag{1.22}
\end{equation*}
$$

To evaluate $I_{1}$ use the following distributional relation

$$
\begin{equation*}
P \frac{1}{\tau}=-\frac{i}{2} \int_{-\infty}^{\infty} e x p(i a \tau) \frac{a}{|a|} d a \tag{1.23}
\end{equation*}
$$

which can be easily proved as follows

$$
\begin{aligned}
& \int_{-\infty}^{\infty} \exp (i a \tau) \frac{a}{|a|} d a=\int_{0}^{\infty} e \operatorname{xp}(i a \tau) d a-\int_{-\infty}^{0} \exp (i a \tau) d a \\
&=\lim _{\varepsilon \rightarrow 0}\left[\int_{0}^{\infty} \exp [i a[\tau+i \varepsilon]] d a-\int_{-\infty}^{0} \exp [i a[\tau-i \varepsilon]] d a\right] \\
& \quad= \lim _{\varepsilon \rightarrow 0}\left[\frac{i}{\tau+i \varepsilon}+\frac{i}{\tau-i \varepsilon}\right] \\
& \quad=2 i P \frac{1}{\tau}
\end{aligned}
$$

where the latter step follows from the relations (1.17). $I_{1}$ can thus be
written as

$$
\begin{equation*}
I_{1}(x)=-\frac{i}{2} \int d a \frac{a}{|a|} \iiint \int e x p(-i a k \cdot k+i k \cdot x) d^{4} k \tag{1.24}
\end{equation*}
$$

Completing the square in the four-dimensional integral and introducing the change of variables

$$
\begin{equation*}
k^{\prime}=k-(2 a)^{-1} x \tag{1.25}
\end{equation*}
$$

and the integrals

$$
\begin{equation*}
\int_{-\infty}^{\infty} d p e x p\left(\mp i a p^{2}\right)=[\pi /( \pm i a)]^{1 / 2} \tag{1.26}
\end{equation*}
$$

we can evaluate the four-dimensional integral to get

$$
\begin{equation*}
I_{1}(x)=\left(\pi^{2} / 2\right) \int d a a^{-2} \exp \left[i x \cdot x(4 a)^{-1}\right] \tag{1.27}
\end{equation*}
$$

The further change of variables $a=(4 a)^{-1}$ then yields

$$
\begin{equation*}
I_{1}(x)=-4 \pi^{3} \delta\left(x^{2}\right) \tag{1.28}
\end{equation*}
$$

which can also be written as

$$
\begin{equation*}
I_{1}(x)=-4 \pi^{2}(2 r)^{-1}\left[\delta\left(r+x_{0}\right)+\delta\left(r-x_{0}\right)\right] \tag{1.29}
\end{equation*}
$$

The latter follows from the general distributional result

$$
\begin{equation*}
\delta(f(x))=\sum_{i} \delta\left(x-x_{i}\right) /\left|f^{\prime}\left(x_{i}\right)\right| \tag{1.30}
\end{equation*}
$$

where $f\left(x_{i}\right)=0$.
We evaluate $I_{2}$ using spherical coordinates with

$$
\begin{align*}
& d \underset{\sim}{k}=\omega_{k}^{2} d \omega_{k} d(\cos \theta) d \varphi \\
& \underset{\sim}{k} \cdot \underset{\sim}{X}=\omega_{k} r \cos \theta \quad ; \quad r=|x| \tag{1.31}
\end{align*}
$$

The angular integrals are straightforward, and we get

$$
\begin{equation*}
\left.I_{2}\left(\underset{\sim}{x}, x_{0}\right)=(2 \pi / i r)\right\}_{0}^{\infty}\left[\exp \left[i \omega_{k}\left(x_{0}+r\right)\right]-\exp \left[i \omega_{k}\left(x_{0}-r\right)\right]\right] d \omega_{k} \tag{1.32}
\end{equation*}
$$

Introduce a small convergence factor

$$
\begin{align*}
& I_{2}\left(x_{\sim}, x_{0}\right)=(2 \pi / i r) \\
& \lim _{\varepsilon \rightarrow 0} \int_{0}^{\infty}\left[\exp \left[i \omega_{k}\left(x_{0}+r+i \varepsilon\right)\right]\right. \\
&\left.\quad-\exp \left[i \omega_{k}\left(x_{0}-r+i \varepsilon\right)\right]\right] d \omega_{k}  \tag{1.33}\\
&=(2 \pi / r) \\
& \lim _{\varepsilon \rightarrow 0}\left[\frac{1}{x_{0}+r+i \varepsilon}-\frac{1}{x_{0}-r+i \varepsilon}\right]
\end{align*}
$$

If we use (1.17) and combine the terms we get

$$
\begin{equation*}
I_{2}\left(x, x_{0}\right)=4 \pi P \frac{1}{x^{2}}+\frac{2 \pi^{2} i}{r}\left[\delta\left(x_{0}-r\right)-\delta\left(x_{0}+r\right)\right] \tag{1.34}
\end{equation*}
$$

By substituting $-x_{0}$ for $x_{0}$ and rewriting we get that

$$
\begin{align*}
I_{2}\left(\underset{\sim}{x},-x_{0}\right) & =\left[I_{2}\left(\underset{\sim}{x}, x_{0}\right)\right]^{*} \\
& =4 \pi P \frac{1}{x^{2}}-\frac{2 \pi^{2} i}{r}\left[\delta\left(x_{0}-r\right)-\delta\left(x_{0}+r\right)\right] \tag{1.35}
\end{align*}
$$

We thus have all the necessary integrals from (1.29), (1.34) and (1.35) and can evaluate our Green's functions from (1.20). The results are:

$$
\begin{equation*}
G_{R}^{(3,1)}(x)=(4 \pi r)^{-1} \delta\left(x_{0}-r\right) . \tag{1.36}
\end{equation*}
$$

This result means that after a time $t$, a pulse in three-dimensions is concentrated on the surface of a sphere of radius $r=x_{0}=c t$ (i.e. it is an out going spherical wave). Also we note that since

$$
\delta\left(x_{0}-r\right)=\delta(c t-r)=c^{-1} \delta(t-r / c),
$$

we have that following (1.8)

$$
G_{R}^{(3,1)}(x)=c^{-1} G_{R}^{(3, x)}(\underset{\sim}{x}, t)
$$

In addition this is called a retarded Green's function since any field a can be expressed as an integral over a source function $f$ as

$$
u(x)=u\left(x, x_{0}\right)=\iiint d x^{\prime} \int d x_{0}^{\prime} G_{R}^{(3,1)}\left(x-x^{\prime}\right) f\left(x_{\sim}^{\prime}, x_{0}\right) .
$$

Substituting the Green's function from (1.36) and evaluating the $x_{0}^{\prime}$ integration yields

$$
u(x)=\iiint d x_{\sim}^{\prime}(4 \pi x)^{-1} f\left(x^{\prime}, x_{0}-x\right),
$$

where $r=|\underset{\sim}{x}-\underset{\sim}{x}|$. The latter is an integral over a source function $f$ evaluated at a retarded time.

## Eg. 2. ADVANCED GREEN'S FONCTION ( $\alpha=\beta=1$ )

We have that

$$
\begin{equation*}
G_{A}^{(3,1)}(x)=(4 \pi r)^{-1} \delta\left(x_{0}+x\right) \tag{1.37}
\end{equation*}
$$

This is an incoming spherical pulse concentrated on a sphere of radius r=ct. Thus for a real pulse it must exist for negative times. It is called an advanced Green's function since any field can be written as the spatial integral over a source function $f$ evaluated at an advanced time $x_{0}+r$ in analogy to the previous discussion.

## Bg. 3. PRINCIPAL VALOR GRERN' $\mathcal{S}$ PONCTIO $(\alpha=\beta=0)$

From (1.20) this is directly related to $I_{1}$ so that

$$
\begin{equation*}
G_{P}^{(3,1)}(x)=(4 \pi)^{-1} \delta\left(x^{2}\right) \tag{1.38}
\end{equation*}
$$

We also note that it can be written as a linear combination of (1.36) and (1.37)

$$
\begin{equation*}
G_{P}^{(3,1)}(x)=\frac{1}{2}\left[G_{R}^{(3,1)}(x)+G_{A}^{(3,1)}(x)\right] \tag{1.39}
\end{equation*}
$$

so that another representation is

$$
\begin{equation*}
G_{P}^{(3,1)}(x)=(8 \pi r)^{-1}\left[\delta\left(x_{0}-r\right)+\delta\left(x_{0}+r\right)\right] \tag{1.40}
\end{equation*}
$$

which al so follows from (1.29) and (1.30). In a sense it is a standing wave Green's function since it balances both incoming and outgoing wave Green's functions.

## Eg. 4. DYSON GRREN'S FONCTION $(\alpha=-1, \beta=1)$

By relating the $I_{1}$ integral to the principal value term (1.38) we get that

$$
\begin{equation*}
G_{D}^{(3,1)}(x)=G_{P}^{(3,1)}(x)+\frac{i}{4 \pi^{2}} P \frac{1}{x^{2}} . \tag{1.41}
\end{equation*}
$$

## Eg. 5. CAOSAL GRERN'S PONCTION ( $\alpha=1, \beta=-1$ )

This is just the complex conjugate of the Dyson function

$$
\begin{equation*}
G_{C}^{(3,1)}(x)=G_{P}^{(3,1)}(x)-\frac{i}{4 \pi^{2}} P \frac{1}{x^{2}} \tag{1.42}
\end{equation*}
$$

Finally we can easily conclude either from the explicit forms of the five functions (1.36), (1.37), (1.38), (1.41) and (1.42) or from the definitions (1.16) and (1.20) that we have the relations

$$
G_{P}^{(3,1)}(x)=\frac{1}{2}\left[G_{R}^{(3,1)}(x)+G_{A}^{(3,1)}(x)\right],
$$

and

$$
\begin{equation*}
G_{P}^{(3,1)}(x)=\frac{1}{2}\left[G_{R}^{(3,1)}(x)+G_{C}^{(3,1)}(x)\right] \tag{1.43}
\end{equation*}
$$

so that only three of the five functions are linearly independent. In addition we also note that the difference of any two of these Green's functions is a solution of the homogeneous equation, i.e. for example

$$
\begin{equation*}
g=G_{R}^{(3,1)}-G_{A}^{(3,1)} \tag{1.44}
\end{equation*}
$$

satisfies

$$
\begin{equation*}
\square \mathrm{g}=0 \tag{1.45}
\end{equation*}
$$

### 1.1.2 (2.1)-DIMBNSIONS

Here we compute the Green's functions in two spatial and one temporal dimension. Fe do this by identifying one spatial coordinate and integrating the (3,1) Green's functions over this coordinate. We choose the z-direction as our direction of integration and write the radins as

$$
\begin{equation*}
r=\left[P^{2}+\left(z-z^{\prime}\right)^{2}\right]^{1 / 2} \tag{1.46}
\end{equation*}
$$

where

$$
\begin{equation*}
\underset{\sim}{\mathbf{P}}=R^{-} \boldsymbol{R}^{\prime}, \quad \mathbf{R}=(x, y) \tag{1.47}
\end{equation*}
$$

will be the remaining two-dimensional vector. The Green's functions we define all satisfy the equation

$$
\begin{equation*}
\left(\partial_{x}^{2}+\partial_{y}^{2}-c^{-2} \partial_{t}^{2}\right) G^{(2,1)}(\underset{\sim}{P}, \tau)=-\delta(\tau) \delta(\underset{\sim}{P}) \tag{1.48}
\end{equation*}
$$

Eg. 1. RETARDED GREEN'S FUNCTION
We define $G_{R}^{(2,1)}$ as the spatial integral over $G_{R}^{(3,1)}$ given by (1.36) where $r=\left|x-x^{\prime}\right|$ and $\tau=x_{0}-x_{0}^{\prime}$. It is

$$
\begin{align*}
\mathbf{G}_{\mathbf{R}}^{(2,1)}(\underset{\sim}{P}, \tau) & =\int_{-\infty}^{\infty} \mathbf{G}_{\mathbf{R}}^{(3,1)}\left(x, x^{\prime}\right) \mathrm{dz}^{\prime}  \tag{1.49}\\
& =\frac{1}{4 \pi} \int_{-\infty}^{\infty} \frac{\delta(x-\tau)}{r} d z^{\prime}
\end{align*}
$$

We define

$$
\begin{align*}
& \zeta=z-z^{\prime} \quad d \zeta=-d z^{\prime} \\
& \mathbf{r}^{2}=\mathbf{P}^{2}+\zeta^{2} \\
& d \zeta / \mathbf{r}=d \mathbf{r} / \zeta=d \mathbf{r}\left(\mathbf{r}^{2}-\mathbf{P}^{2}\right)^{-1 / 2} . \tag{1.50}
\end{align*}
$$

so that

$$
G_{R}^{(2,1)}(\underset{\sim}{P}, \tau)=\frac{1}{2 \pi} \int_{0}^{\infty} \frac{\delta(\mathbf{r}-\tau) d r}{\left(\mathbf{r}^{2}-\mathbf{P}^{2}\right)^{1 / 2}} .
$$

and, evaluating the $\delta$-function, we get

$$
\begin{equation*}
G_{R}^{(2,1)}(\underset{\sim}{P}, \tau)=\frac{1}{2 \pi} \frac{\theta(\tau-P)}{\left(\tau^{2}-P^{2}\right)^{1 / 2}} \tag{1.51}
\end{equation*}
$$

This illustrates the fact that, in two dimensions, the effect of an impulse after a (scaled) time $\tau$ has elapsed has spread over a region of spatial extent $P<\tau$. A line source in three dimensions produces a field which at any given point has a tail.

We can analogously define the other Green's functions as integrals over the corresponding (3,1) dimensional Green's functions. From (1.37), (1.40), (1.41) and (1.42) we get

Bg. 2. ADVANCED GREEN'S PUNCTION

$$
\begin{equation*}
G_{A}^{(2,1)}(\underset{\sim}{P}, \tau)=\frac{1}{2 \pi} \frac{\theta(-\tau-P)}{\left(\tau^{2}-P^{2}\right)^{1 / 2}} . \tag{1.52}
\end{equation*}
$$

## Eg. 3. PRINCIPAL VALOB GRERN'S FUNCTION

$$
\begin{equation*}
G_{P}^{(2,1)}(\underset{\sim}{P}, \tau)=\frac{1}{4 \pi} \frac{1}{\left(\tau^{2}-P^{2}\right)^{1 / 2}}[\theta(\tau-P)+\theta(-\tau-P)] \tag{1.53}
\end{equation*}
$$

Bg- 4. DYSON GRERN' S FDNCTION

$$
\begin{equation*}
\left.\left.G_{D}^{(2,1)} \underset{\sim}{P}, \tau\right)=G_{P}^{(2,1)} \underset{\sim}{P}, \tau\right)+\frac{i}{4 \pi} \frac{\theta(P-|\tau|)}{\left[P^{2}-\tau^{2}\right]^{1 / 2}} \tag{1.54}
\end{equation*}
$$

and finally
Eg. 5. CAUSAL GREEN' S PONCTION

$$
\begin{equation*}
\left.G_{C}^{(2,1)}(\underset{\sim}{P}, \tau)=G_{P}^{(2,1)} \underset{\sim}{P}, \tau\right)-\frac{i}{4 \pi} \frac{\theta(P-|\tau|)}{\left[P^{2}-\tau^{2}\right]^{1 / 2}} . \tag{1.55}
\end{equation*}
$$

$1.1 .3(1,1)$-DIMENSION
The Green's function in one spatial and one temporal dimension is defined as

$$
\begin{equation*}
G^{(1,1)}(\xi, \tau) \quad \xi=x-x^{\prime}, \tau=x_{0}-x_{0}^{\prime}=c\left(t-t^{\prime}\right) \tag{1.56}
\end{equation*}
$$

and satisfies the equation

$$
\begin{equation*}
\left(\partial_{I}^{2}-\partial_{0}^{2}\right) G^{(1,1)}(\xi, \tau)=-\delta(\xi) \delta(\tau) \tag{1.57}
\end{equation*}
$$

We can compute the five Green's functions either by integrating over the $y$ coordinate results $G^{(2,1)}(\underset{\sim}{(P, \tau)}$ from Sec. 1.1 .2 , or by using pole shifting and complex integration techniques as we used in Sec. 1.1.1. We choose the 1atter. Introduce the Fourier transform

$$
G^{(1,1)}(\xi, \tau)=(2 \pi)^{-2} \iint \exp \left[i\left(k_{x} \xi-k_{0} \tau\right)\right] \tilde{G}^{(1,1)}\left(k_{x}, k_{0}\right) d k_{x} d k_{0} \cdot(1.58)
$$

and apply the differential operator in (1.57) to it. We can thus solve for the Fourier transform and write (1.58) as

$$
\begin{equation*}
G^{(1,1)}(\xi, \tau)=-(2 \pi)^{-2} \int \exp \left(i k_{x} \xi\right) d k_{x} \int \frac{e \operatorname{xp}\left(-i k_{0} \tau\right)}{\left(k_{0}^{2}-k_{z}^{2}\right)} d \tau \tag{1.59}
\end{equation*}
$$

Shift the poles of the integrand as in Sec. 1.1.1. That is we have that

$$
\begin{align*}
\frac{1}{\left(k_{0}-k_{x}\right)\left(k_{0}+k_{x}\right)} & \rightarrow \frac{1}{\left(k_{0}-k_{x}-i \alpha \varepsilon\right)\left(k_{0}+k_{x}-i \beta \varepsilon\right)} \\
= & P\left[\frac{1}{k_{0}^{2}-k_{x}^{2}}\right] \\
& +\frac{\pi i}{2 k_{x}}\left[\alpha \delta\left(k_{0}-k_{x}\right)-\beta \delta\left(k_{0}+k_{x}\right)\right] \tag{1.60}
\end{align*}
$$

where we have used (1.17). The result for (1.59) is

$$
\begin{equation*}
G^{(1,1)}(\xi, \tau)=-(2 \pi)^{-2}\left[I_{2}(\xi, \tau)+(\pi i / 2)\left[\alpha I_{2}(\xi,-\tau)-\beta I_{2}(\xi, \tau)\right]\right] . \tag{1.61}
\end{equation*}
$$

in analogy to (1.20) where here

$$
\begin{equation*}
I_{1}(\xi, \tau)=P \int_{-\infty}^{\infty} \int_{0}^{\exp \left[i\left(k_{x} \xi-k_{0} \tau\right)\right]} \underset{k_{0}^{2}-k_{x}^{2}}{d k_{x}} d k_{0} . \tag{1.62}
\end{equation*}
$$

and

$$
\begin{equation*}
I_{2}(\xi, \tau)=\int_{-\infty}^{\infty} \frac{\exp \left[i k_{x}(\xi+\tau)\right]}{k_{x}} d k_{x} . \tag{1.63}
\end{equation*}
$$

The values $\alpha$ and $\beta$ are given by (1.16).
The integrals can be easily evaluated using complex variable techniques. The results are

$$
\begin{equation*}
I_{1}(\xi, \tau)=-\left(\pi^{2} / 2\right) \operatorname{sgn} \tau[\operatorname{sgn}(\xi+\tau)-\operatorname{sgn}(\xi-\tau)] \tag{1.64}
\end{equation*}
$$

or

$$
\begin{equation*}
I_{1}(\xi, \tau)=-\pi^{2}[\theta(\tau-|\xi|+\theta(-\tau-|\xi|)] \text {. } \tag{1.65}
\end{equation*}
$$

and

$$
\begin{equation*}
I_{2}(\xi, \tau)=\pi i \quad \operatorname{sgn}(\xi+\tau) \tag{1.66}
\end{equation*}
$$

We combine these results using (1.64)-(1.66) and (1.16) in (1.61) to get the results:

Ege 1. RETARDED GRRRN'S PDNCTION

$$
\begin{equation*}
G_{R}^{(1,1)}(\xi, \tau)=\frac{1}{2} \theta(\tau-|\xi|) \tag{1.67}
\end{equation*}
$$

Eg. 2. ADVANCED GRBRN' S FONCTIOH

$$
\begin{equation*}
G_{A}^{(1, x)}(\xi, \tau)=\frac{1}{2} \theta(-\tau-|\xi|) \tag{1.68}
\end{equation*}
$$

Eg. 3. PRINCIPAL VALDE GREEN' S FDNCTION

$$
\begin{equation*}
G_{P}^{(1,1)}(\xi, \tau)=\frac{1}{4}[\theta(\tau-|\xi|)+\theta(-\tau-|\xi|)] \tag{1.69}
\end{equation*}
$$

Eg. 4. DYSON GRERN' S FONCTION

$$
\mathbf{G}_{\mathbf{D}}^{(1,1)}(\xi, \tau)=-\frac{1}{4}[\theta(\tau) \operatorname{sgn}(\xi-\tau)+\theta(-\tau) \operatorname{sgn}(\xi+\tau)]
$$

(1.70)
and

Bg. 5. CADSAL GRBENS FONCTION

$$
\begin{equation*}
\mathbf{G}_{C}^{(1,1)}(\xi, \tau)=\frac{1}{4}[\theta(\tau) \operatorname{sgn}(\xi+\tau)+\theta(-\tau) \operatorname{sgn}(\xi-\tau)] \tag{1.71}
\end{equation*}
$$

Other compact values can al so be derived, for example

$$
\begin{equation*}
G_{D}^{(1,1)}(\xi, \tau)=-\frac{1}{4} \operatorname{sgn}[\xi-|\tau|] \tag{1.72}
\end{equation*}
$$

and

$$
\begin{equation*}
G_{C}^{(1,1)}(\xi, \tau)=\frac{1}{4} \operatorname{sgn}[\xi+|\tau|] \tag{1.73}
\end{equation*}
$$

### 1.2 GREEN'S FONCTIONS FOR TRE BRLMHOLTZ EQDATION

## 1.2 .1 (3)-DIMENSIONS

We can define the Green's functions for the Helmholtz equation as the temporal Fourier transform of the Green's functions for the wave equation derived in Sec. 1. Osing the four-dimensional formulation we have

$$
\begin{equation*}
G^{(3)}\left(x, X^{\prime}\right)=\int_{-\infty}^{\infty} G^{(3,1)}\left(x, x^{\prime}\right) \exp \left(i k_{0} \tau\right) d \tau \tag{2.1}
\end{equation*}
$$

where $\tau=x_{0}-x_{0}^{\prime}$. They satisfy the Helmholtz equation given by the same Fourier transform operating on (1.6). It is

$$
\begin{equation*}
\text { v } \quad\left(\nabla_{x}^{2}+k_{0}^{2}\right) G^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)=-\delta\left(\underset{\sim}{x}-x^{\prime}\right), \tag{2.2}
\end{equation*}
$$

where $k_{0}=\omega / c$ and $\omega$ is circular frequency. We compate each of the Green's functions corresponding to the pole shifts in Sec. 1.

Eg. 1. RETARDED GRERN'S FDNCTION
From (1.36) we have that for a difference of argnments (r=|x-x|)

$$
\begin{equation*}
G_{R}^{(3,1)}\left(x, x^{\prime}\right)=G_{R}^{(3,1)}\left(x-x^{\prime}\right)=(4 \pi r)^{-1} \delta(\tau-r) \tag{2,3}
\end{equation*}
$$

Substitute this in (2.1) to get

$$
\begin{equation*}
G_{R}^{(3)}\left(x, x^{\prime}\right)=(4 \pi x)^{-1} e x p\left(i k_{0} r\right) \tag{2.4}
\end{equation*}
$$

which is an outgoing spherical wave, i.e. for harmonic time dependence

$$
\begin{equation*}
\exp (-i \omega t)=\exp \left(-i k_{0} x_{0}\right) \tag{2.5}
\end{equation*}
$$

the wave travels in a positive radial direction and satisfies an out going radiation condition of the form

$$
\begin{equation*}
\lim _{r \rightarrow \infty}\left[\frac{\partial}{\partial r}-i k_{0}\right] G_{R}^{(3)}=0\left(r^{-2}\right) \tag{2.6}
\end{equation*}
$$

It expresses the field at the receiver point $\underset{\sim}{x}$ due to a point source located at $]^{\prime}$ in a homogeneous medium.

Eg. 2. ADVANCBD GREEN' S PDNCTION
From (1.37) we have that

$$
\begin{equation*}
G_{A}^{(3,1)}\left(x, x^{\prime}\right)=(4 \pi r)^{-1} \delta(\tau+r) \tag{2.7}
\end{equation*}
$$

Which when substituted into (2.1) yields

$$
\begin{equation*}
G_{A}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=(4 \pi r)^{-1} \exp \left(-i k_{0} r\right) \tag{2.8}
\end{equation*}
$$

For harmonic time dependence this is an incoming radial wave satisfying the radiation condition

$$
\begin{equation*}
\lim _{r \rightarrow \infty}\left[\frac{\partial}{\partial r}+i k_{0}\right] G_{A}^{(3)}=0\left(r^{-2}\right) \tag{2.9}
\end{equation*}
$$

## Eg. 3. PRINCIPAL VALOE GRERN'S FONCTION

This can be computed either directly from (1.40) using a difference of argaments or from (1.39) represented here as half the sum of (2.3) and (2.8). The result is

$$
\begin{equation*}
G_{P}^{(3)}\left(\underline{x}, x^{\prime}\right)=(4 \pi r)^{-1} \cos \left(k_{0} r\right) \tag{2.10}
\end{equation*}
$$

which for harmonic time dependence represents a standing wave. Note al so that in contrast to the retarded and advanced functions, the principal value Green's function is real.

## Eg. 4. DYSON AND CADSAL GREEN' S FONCTIONS

From (1.41) and (1.42) we have that

$$
\begin{equation*}
G_{D, C}^{(3,1)}\left(x, x^{\prime}\right)=G_{P}^{(3,1)}\left(x, x^{\prime}\right) \pm \frac{i}{4 \pi^{2}} P \frac{1}{\left(x-x^{\prime}\right)^{2}} \tag{2.11}
\end{equation*}
$$

We substitute this into (2.1) and use (2.10) for the evaluation of the principal value term. The remaining integral can be evaluated using residue calculus methods. The result is

$$
\begin{equation*}
G_{D, C}^{(3)}\left(X, x^{\prime}\right)=(4 \pi r)^{-1}\left[\cos \left(k_{0} r\right) \pm i \operatorname{sgn}\left(k_{0}\right) \sin \left(k_{0} r\right)\right], \tag{2.12}
\end{equation*}
$$

which can be rewritten as

$$
G_{D}^{(3)}\left(\underset{\sim}{\left.x, x^{\prime}\right)}= \begin{cases}G_{R}^{(3)}\left(\underset{\sim}{\left.x, I_{\sim}^{\prime}\right)}\right. & k_{0}>0  \tag{2.13}\\ G_{A}^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right) & k_{0}<0,\end{cases}\right.
$$

and

$$
G_{C}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\left\{\begin{array}{ll}
G_{A}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right) & k_{0}>0  \tag{2.14}\\
G_{R}^{(3)}\left(x, x^{\prime}\right) & k_{0}<0
\end{array} .\right.
$$

Becanse the Dyson and Causal Green's functions mix representations and do not satisfy either a well-defined radiation condition or a standing wave interpretation except for positive or negative frequency separately, they are not useful for our purposes. In addition negative frequency results are usually folded into positive frequency ones in applications, and neither the Dyson or Causal functions yield new results over and above those found from the retarded, advanced, and principal value functions. We do not compute them for (2) and (1) dimensions.

### 1.2.2 (2)-DIMENSIONS

The two-dimensional Helmholtz Green's functions are the temporal Fourier transforms of the Green's functions for the two-dimensional wave equation in Sec. 1.1.2. They are defined as

$$
\begin{equation*}
G^{(2)}\left(R, R^{\prime}\right)=\int_{-_{\infty}}^{\infty} G^{(2,1)}(\underset{\sim}{P}, \tau) \exp \left(i k_{0} \tau\right) d \tau \tag{2.15}
\end{equation*}
$$

and satisfy the Helmholtz equation given by the corresponding transform of (1.48) which becomes

$$
\begin{equation*}
\left(\partial_{z}^{2}+\partial_{y}^{2}+k_{0}^{2}\right) G^{(2)}\left(R, R^{\prime}\right)=-\delta\left(x-x^{\prime}\right) \delta\left(y-y^{\prime}\right) . \tag{2,16}
\end{equation*}
$$

Eg. 1. RETARDED GREEN'S FUNCTION
For this case we substitute (1.51) in (2.15) to get

$$
\begin{equation*}
G_{R}^{(2)}\left(R, R^{\prime}\right)=\frac{1}{2 \pi} \int_{P}^{\infty} \frac{\exp \left(i k_{0} \tau\right)}{\left(\tau^{2}-P^{2}\right)^{1 / 2}} d \tau \tag{2.17}
\end{equation*}
$$

If we make the substitution

$$
\tau=P \cosh \phi \quad ; \quad d \tau=P \sinh \phi d \phi
$$

the integral becomes

$$
\begin{equation*}
G_{R}^{(3)}\left(\rho \cdot R^{\prime}\right)=\frac{1}{4 \pi} \int_{-\infty}^{\infty} e \operatorname{xp}\left(i k_{0} P \cosh \phi\right) d \phi \quad . \tag{2.18}
\end{equation*}
$$

written from $-\infty$ to $\infty$ since the integrand is an even function of $\phi$. The integral (2.18) is a representation of the Bankel function

$$
\begin{equation*}
G_{R}^{(3)}\left(R \cdot R^{\prime}\right)=\frac{i}{4} H_{0}^{(1)}\left(k_{0} P\right) \tag{2.19}
\end{equation*}
$$

which is an outgoing cylindrical wave, i.e. asymptotically

$$
\begin{equation*}
H_{0}^{(1)}\left(k_{0} P\right) \sim\left(2 / \pi i k_{0} P\right)^{1 / 2} \exp \left(i k_{0} P\right), \tag{2.20}
\end{equation*}
$$

which spreads like a cylindrical wave with amplitude factor $\mathbf{P}^{\mathbf{- 1 / 2}}$.

## Eg. 2. ADVANCED GRREN'S FUNCTION

Substitute (1.52) into (2.15). The integration can be performed as above with an additional sign change in $\tau$. The result is the incoming Hankel function

$$
\begin{equation*}
G_{A}^{(2)}\left(\rho, \rho^{\prime}\right)=-\frac{i}{4} H_{0}^{(2)}\left(k_{0} P\right) \tag{2.21}
\end{equation*}
$$

which represents a cylindrical wave propagating in the direction of decreasing $r$.

## Eg. 3. PRINCIPAL VALOR GRREN'S PUNCTION

This can be computed either directly by substituting (1.53) in (2.15) or as half the sum of (2.19) and (2.21). The result is

$$
\begin{equation*}
G_{P}^{(2)}\left(p, P^{\prime}\right)=-\frac{1}{4} N_{0}\left(k_{0} P\right) \tag{2,22}
\end{equation*}
$$

Where $N_{0}$ is the Nemann function.

### 1.2.3 (1)-DIMRNSION

We define the one-dimensional Helmholtz Green's function as the temporal Fourier transform of the one-dimensional Green's functions for the wave equation in Sec. 1.1.3. It is

$$
\begin{equation*}
G^{(1)}\left(x, x^{\prime}\right)=\int_{-\infty}^{\infty} G^{(1,1)}(\xi, \tau) \exp \left(i k_{0} \tau\right) d \tau \tag{2.23}
\end{equation*}
$$

where $\xi=x-x^{\prime}$. Fourier transformation of the differential equation (1.57) yields the ordinary differential equation

$$
\begin{equation*}
\left[\frac{d^{2}}{d x^{2}}+k_{0}^{2}\right] G^{(1)}\left(x, x^{\prime}\right)=-\delta\left(x-x^{\prime}\right) \tag{2.24}
\end{equation*}
$$

which is the one-dimensional version of the Helmholtz equation satisfied by all the one-dimensional Green's functions below.

Substitute (1.67) in (2.23). Evaluation of the step function yields

$$
\begin{equation*}
G_{R}^{(1)}\left(x, z^{\prime}\right)=\frac{1}{2} \int_{|\xi|}^{\infty} \exp \left(i k_{0} \tau\right) d \tau \tag{2.25}
\end{equation*}
$$

Recall that this Green's function for the wave equation was computed with $\mathbf{k}_{0}$ shifted to $k_{0+\text {. We can then directly evaluate the integral in (2.25) since }}$ the contribution at $\infty$ vanishes. The result is

$$
\begin{equation*}
G_{R}^{(1)}\left(x, x^{\prime}\right)=-\left(2 i k_{0}\right)^{-1} \exp \left(i k_{0}|x-x|\right) \tag{2.26}
\end{equation*}
$$

which is a one-dimensional wave which travels to the right.

## Eg. 2. ADVANCED GREBN' S FUNCTION

Substitute (1.69) in (2.23). The integral evaluation proceeds in the same manner as the previous example except that here we note that the advanced Green's function is computed with $k_{0}$ shifted to $k_{0}$. . The result is

$$
\begin{equation*}
G_{A}^{(1)}\left(x, x^{\prime}\right)=\left(2 i k_{0}\right)^{-1} \exp \left(-i k_{0}\left|x-x^{\prime}\right|\right) \tag{2.27}
\end{equation*}
$$

which for harmonic time dependence is a one-dimensional wave travelling to the 1 eft.

## Eg. 3. PRINCIPAL VALOE GREBN' S FONCTION

This can be computed either directly from (1.69) or by combining half the $s$ wn of (2.26) and (2.27) to give

$$
\begin{equation*}
G_{P}^{(1)}\left(x, x^{\prime}\right)=-\left(4 k_{0}\right)^{-1} \sin \left(k_{0}\left|x-x^{\prime}\right|\right) \tag{2.28}
\end{equation*}
$$

which represents a standing wave.

### 1.3 CAUSAL PARABOLIC GREEN'S FONCTION

The causal Green's function which satisfies the parabolic wave equation

$$
\begin{equation*}
\nabla_{x}^{2} g\left(\underset{\sim}{x}, x_{\sim}^{\prime}, t, t^{\prime}\right)-2 \gamma \frac{\partial g}{\partial t}=-\delta\left(\underset{\sim}{x}-x^{\prime}\right) \delta\left(t-t^{\prime}\right) \tag{3.1}
\end{equation*}
$$

which contains only the first derivative in time and where $\gamma$ is a constant can be computed as follows. Causality means that there is no measurable effect until the source tarns on, i.e. g must vanish for times $t$ less than the source turn-on time t'. This is

$$
\begin{equation*}
g\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime} ; t, t^{\prime}\right)=0 \quad t<t^{\prime} \tag{3.2}
\end{equation*}
$$

Introduce the Fourier transform in $\underset{\sim}{x}$

$$
\begin{equation*}
\tilde{g}\left(\underset{\sim}{k}, x^{\prime} ; t, t^{\prime}\right)=\iiint e x p(-i k \cdot \underset{\sim}{x}) g\left(\underset{\sim}{x}, x^{\prime} ; t, t^{\prime}\right) d \underset{\sim}{x}, \tag{3.3}
\end{equation*}
$$

and correspondingly Fourier transform (3.1) to get

$$
\begin{equation*}
2 \gamma \frac{\partial \tilde{g}}{\partial t}+k^{2} \tilde{g}=e \operatorname{xp}\left(-i \underset{\sim}{k} \cdot{\underset{\sim}{x}}^{\prime}\right) \delta\left(t-t^{\prime}\right) \tag{3.4}
\end{equation*}
$$

where $k$ is the Fourier transform variable and $k^{2}=k \cdot k \cdot$ We still require the condition (3.2) and the one-dimensional equation (3.4) has solution given by

$$
\begin{equation*}
\tilde{g}\left(\underset{\sim}{k},{\underset{\sim}{x}}^{\prime} ; t, t^{\prime}\right)=(2 \gamma)^{-1} e x p\left[-i \underset{\sim}{n}{\underset{\sim}{x}}^{\prime}-k^{2}\left(t-t^{\prime}\right) / 2 \gamma\right] \theta\left(t-t^{\prime}\right) \tag{3.5}
\end{equation*}
$$

where the step function defines causality. The inverse transform is

$$
\begin{equation*}
g\left(\underset{\sim}{x}, x^{\prime} ; t, t^{\prime}\right)=(2 \pi)^{-3} \iiint e \operatorname{xp}(i k \cdot \underset{\sim}{x}) \underset{\sim}{\sim}\left(\underset{\sim}{x}, x^{\prime} ; t, t^{\prime}\right) d \underset{\sim}{k}, \tag{3.6}
\end{equation*}
$$

and if we substitute (3.5) into (3.6) we note that the result is the Fourier
transform of a Gaussian. The latter may be treated by cartesian coordinates to yield

$$
\begin{equation*}
g\left(x, x^{\prime} ; t, t^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \frac{\pi^{3 / 2}(2 \gamma)^{1 / 2}}{\left(t-t^{\prime}\right)^{3 / 2}} e x p\left[-\frac{\gamma}{2} \frac{\left|x_{\sim}-z^{\prime}\right|^{2}}{t-t^{\prime}}\right] \theta\left(t-t^{\prime}\right) \tag{3.7}
\end{equation*}
$$

It can be shown that this result can be generalized to n-spatial dimensions to yield

$$
g^{(n)}\left(\frac{x}{\sim}, x^{\prime} ; t, t^{\prime}\right)=\frac{1}{(2 \pi)^{n}} \frac{1}{2 \gamma}\left[\frac{2 \gamma \pi}{t-t^{\prime}}\right]^{n / 2} \exp \left[-\frac{\gamma}{2} \frac{\left|\frac{x-x^{\prime}}{\sim}\right|^{2}}{t-t^{\prime}}\right] \theta\left(t-t^{\prime}\right)
$$

where for

$$
\begin{array}{ll}
n=1 & \underset{\sim}{x}=(x) \\
n=2 & \underset{\sim}{x}=(x, y) \\
n=3 & \underset{\sim}{\sim}=(x, y, z) \\
\vdots & \underset{\sim}{n}=\left(x_{1}, x_{2}, \ldots, x_{n}\right)
\end{array}
$$

### 1.4 RERPRESENTATIONS

There are several useful integral representations of the Green's function for the Helmholtz equation. The latter is a spherical wave and the representations amount to expanding a spherical wave into either plane waves (spectral integrals) or cylindrical waves (Sommerfeld and Weyrich representations).

### 1.4.1 MEYL REPRESENTATION

The first representation is an expansion of a spherical wave into plane waves. The Helmholtz Green's function in three dimensions satisfies the differential equation

$$
\begin{equation*}
\left.\nabla^{2}+k_{0}^{2}\right) G^{(3)}\left(x, x^{\prime}\right)=-\delta\left(x_{\sim}^{-x^{\prime}}\right) \tag{4.1}
\end{equation*}
$$

Fourier transform this equation with respect to $\underset{\sim}{x}$, i.e. maltiply the equation by

$$
\begin{equation*}
\iiint e \exp \left[-i\left(k_{x} x+k_{y} y+k_{z} z\right)\right] d x d y d z \tag{4.2}
\end{equation*}
$$

to get the mixed representation

$$
\begin{equation*}
\tilde{G}^{(3)}\left(\underset{\sim}{k},{\underset{\sim}{N}}^{\prime}\right)=\exp \left(i \underset{\sim}{k} \cdot{\underset{\sim}{x}}^{\prime}\right)\left(k^{2}-k_{0}^{2}\right)^{-1} . \tag{4.3}
\end{equation*}
$$

where

$$
\begin{equation*}
k^{2}=|k|^{2}=k_{z}^{2}+k_{y}^{2}+k_{z}^{2}=k_{t}^{2}+k_{z}^{2} . \tag{4.4}
\end{equation*}
$$

and $\underset{\sim}{k}$ is the Fourier transform variable. Note that in (4.4), by defining the transverse part of the wavenumber

$$
\begin{equation*}
k_{t}=\left(k_{x}^{2}+k_{y}^{2}\right)^{1 / 2} \tag{4.5}
\end{equation*}
$$

we have essentially picked ont the z-direction as special. Of course we could do this with any of the three directions. Osing (4.3) the inverse Fonrier transform is thas

$$
\begin{align*}
& G^{(3)}\left(x \cdot x^{\prime}\right)=(2 \pi)^{-3} \iiint \exp \left[i\left(k_{z} x+k_{y} y+k_{z} z\right)\right] \tilde{G}\left(\underset{\sim}{x}, x^{\prime}\right) d \underset{\sim}{x}  \tag{4.6}\\
& =\frac{1}{(2 \pi)^{3}} \iiint \frac{\exp \left[i\left[k_{x}\left(x-x^{\prime}\right)+k_{y}\left(y-y^{\prime}\right)+k_{z}\left(z-z^{\prime}\right)\right]\right]}{\left(k_{z}-K\right)\left(k_{z}+K\right)} d k_{x} d k_{y} d k_{z} \tag{4.7}
\end{align*}
$$

where

$$
\begin{equation*}
K=\left(k_{0}^{2}-k_{t}^{2}\right)^{1 / 2} \tag{4.8}
\end{equation*}
$$

and where we have distinguished the poles in the integrand of (4.7) as poles in $k_{z}$. We evaluate (4.7) cylindrically, i.e. we do the $k_{z}$ integral first using complex variables. We shift the poles by adding a small positive imaginary part to $k_{0}$ and hence to $K$. That is

$$
\mathbf{k}_{0} \rightarrow \mathbf{k}_{0}+i \varepsilon \doteq \mathbf{k}_{0+} ; \quad K \rightarrow K+i \varepsilon \doteq K_{+}
$$

Then in the complex $k_{z}$-plane we have


$$
\begin{aligned}
& z-z^{\prime}>0 \\
& z-z^{\prime}<0
\end{aligned}
$$

We evaluate the $k_{z}$-integral in (4.7) by closing the contonr in the upper half plane ( $\left.z-z^{\prime}\right\rangle 0$ ) or in the 1 ower half plane ( $z-z^{\prime}\langle 0)$. The result is

$$
\begin{equation*}
\int \frac{e \operatorname{xp}\left[i k_{z}\left(z-z^{\prime}\right)\right]}{\left(k_{z}-\bar{K}_{+}\right)\left(k_{z}+\bar{K}_{+}\right)} d k_{z}=\frac{\pi i}{K} \exp \left[i K\left|z-z^{\prime}\right|\right] \tag{4.9}
\end{equation*}
$$

Using this in (4.7) and writing the remaining two integrals using the twodimensional vectors

$$
\begin{equation*}
{\underset{\sim}{k}}_{t}=\left(k_{x}, k_{y}\right) \quad \underset{\sim}{x}=(x, y) \tag{4.10}
\end{equation*}
$$

we get

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)=\frac{\pi i}{(2 \pi)^{3}} \iint \frac{\exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)+i K_{+}\left|z^{-z}\right|\right]}{K_{t}}{\underset{\sim}{z}}^{k_{t}} \tag{4.11}
\end{equation*}
$$

where $\operatorname{Im} K>0$.

Equation (4.11) is the two-dimensional plane-wave spectral representation or Weyl representation. In deriving it we have singled out the $z$-direction as special. This is appropriate if the z-direction in the application is special, for example if there is a discontinuity in $z$ or if the variability in the medium is in the z-direction. We treat this further in the next section. Also notice that here we shifted both poles by tie, so that effectivey one shifted above the axis and one below. For the retarded Green's function is Sec. 1 we shifted both $\omega_{k}$ poles down. This is equivalent to what we have done here since we had

$$
\begin{aligned}
& k_{0}-\omega_{k} \rightarrow k_{0}-\left(\omega_{k}-i \varepsilon\right)=k_{0}+i \varepsilon-\omega_{k} . \\
& k_{0}+\omega_{k} \rightarrow k_{0}+\left(\omega_{k}+i \varepsilon\right)=k_{0}+i \varepsilon+\omega_{k} .
\end{aligned}
$$

In both terms we give a positive shift to the $k_{0}$ term. By shifting the poles in the manner above we have derived the Weyl representation for the retarded Green's function. Other pole shifts can be done to form a Weyl representation for the advanded or principal value Green's functions for examp1e.

### 1.4.2 SOMHERFBLD REPRBSENTATION

For this case we expand the spherical wave in cylindrical waves. The expansion is essentially over the horizontal wave number. We begin by representing the Weyl representation (4.11) in cylindrical polar coordinates defined about $x-x^{\prime}$ and $y-y^{\prime}$. We have that

$$
\begin{equation*}
\exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)\right]=\exp \left(i k_{t} \rho \cos \theta\right) \tag{4.12}
\end{equation*}
$$

where

$$
\begin{equation*}
\rho=\left[\left(x-x^{\prime}\right)^{2}+\left(y-y^{\prime}\right)^{2}\right]^{1 / 2}=\left|x_{t}-x_{t}^{\prime}\right| \tag{4.13}
\end{equation*}
$$

and $\theta$ is the angle between $k_{t}$ and ${\underset{z}{t}}^{-} x_{t}$. Dsing the cylindrical differential area element

$$
\begin{equation*}
d k_{t}=k_{t} d k_{t} d \theta \tag{4.14}
\end{equation*}
$$

and the definition of the Bessel function (cylindrical wave)

$$
\begin{equation*}
J_{0}\left(k_{t} \rho\right)=(2 \pi)^{-1} \int_{0}^{2 \pi} e x p\left(i k_{t} \rho \cos \theta\right) d \theta \tag{4.15}
\end{equation*}
$$

we evaluate the $\theta$-integral as above to get from (4.11)

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)=\frac{i}{4 \pi} \int_{0}^{\infty} \frac{J_{0}\left(k_{t} \rho\right) \exp \left[i\left(k_{0+}^{2}-k_{t}^{2}\right)^{1 / 2}\left|z-z^{\prime}\right|\right]}{\left(k_{0+}^{2}-k_{t}^{2}\right)^{1 / 2}} k_{t} d k_{t} \tag{4.16}
\end{equation*}
$$

where we have explicitly written out the square root $K$. The result (4.16) is a one-dimensional integral representation of a spherical wave in terms of cylindrical waves called the Sommerfeld representation. It is an integral written over the horizontal wavenamber $k_{t}$, and is only useful for problems which contain an analogous horizontal symmetry (i.e. a parametric independence of $\boldsymbol{\theta}$ ).

Alternatively, we can write (4.16) in terms of the Bankel function. The Bessel function $J_{0}$ can be written in terms of Hankel functions $H_{0}^{(1)}$ and $\mathrm{H}_{0}^{(2)}$ as

$$
\begin{equation*}
J_{0}\left(k_{t} \rho\right)=1 / 2\left[H_{0}^{(1)}\left(k_{t} \rho\right)+H_{0}^{(2)}\left(k_{t^{2}}\right)\right] \tag{4.17}
\end{equation*}
$$

Substitute this into (4.16), and use, in the $H_{o}^{(2)}$ integral, the result

$$
\begin{equation*}
H_{0}^{(2)}\left(k_{t} r\right)=-H_{0}^{(1)}\left(e^{\pi i} k_{t} r\right), \tag{4.18}
\end{equation*}
$$

and, in this integral rotate the contour by defining a new variable

$$
\begin{equation*}
k_{t}^{\prime}=e^{\pi i} k_{t} \tag{4.19}
\end{equation*}
$$

so that the limits of integration go from $(0, \infty)$ to $\left(0, \infty e^{\pi i}\right)=(0,-\infty)$. The result is an integral over only $\mathrm{H}_{0}^{(1)}$ given by

$$
\begin{equation*}
G^{(3)}\left(x_{N}, x_{\sim}^{\prime}\right)=\frac{i}{8 \pi} \int \frac{H_{0}^{(1)}\left(k_{t} \rho\right) \exp \left[i\left[k_{0+}^{2}-k_{t}^{2}\right]^{1 / 2}\left|z-z^{\prime}\right|\right]_{k_{t}} d k_{t} .}{\left[k_{0+}^{2}-k_{t}^{2}\right]^{1 / 2}} \tag{4.20}
\end{equation*}
$$

Since $H_{0}^{(1)}$ behaves like an outgoing wave asymptotically, the representation (4.20) is useful in problems which contain this type of geometry, e.g. the exterior problem of scattering from a bounded object. Equation (4.16) on the other hand is useful for an interior representation, i.e. one which contains standing waves rather than outgoing waves.

### 1.4.3 EXPLICIT EVALDATION OR $\theta^{(3)}$

We mentioned that our representations were for the retarded Green's function. We can explicitly exhibit this by evaluating all the integrals in $G^{(3)}$. From (4.7) we have that

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \iiint \frac{\exp \left[i \underset{\sim}{k} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right]}{k^{2}-k_{0}^{2}} d \underline{\sim} \tag{4.21}
\end{equation*}
$$

Using spherical polar coordinates defined as in Fig. 1.2


Fig. 1.2
we have that

$$
\begin{align*}
\underset{\sim}{k} \cdot\left(\underset{\sim}{x}-x^{\prime}\right) & =\left.k\right|_{x}-x^{\prime} \mid \cos \theta=k r \cos \theta  \tag{4.22}\\
d \underline{\sim} & =k^{2} d k \sin \theta d \theta d \phi
\end{align*}
$$

The $\phi$-integral just yields $2 \pi$. The $\theta$-integral is

$$
\begin{equation*}
\int_{0}^{\pi} e \operatorname{xp}(i k r \cos \theta) \sin \theta d \theta=[e x p(i k r)-e x p(-i k r)](i k r)^{-1} \tag{4.23}
\end{equation*}
$$

We evaluate the positive exponential integral using complex variables and pole shifts to $k_{0}+i \varepsilon$ to get (close in uhp)

$$
\begin{equation*}
\int_{\infty_{\infty}}^{\infty} \frac{k e \exp (i k r)}{\left(k-k_{0+}\right)\left(k+k_{0+}\right.} d k=\pi i \exp \left(i k_{0} r\right) \tag{4.24}
\end{equation*}
$$

The negative exponential is evaluated by closing in the lower half plane with the same pole shifts ( $\left.k_{0+}+i \varepsilon\right)$ to yield

$$
\begin{equation*}
\int_{-\infty}^{\infty} \frac{k \exp (-i k r)}{\left(k-k_{0+}\right)\left(k+k_{0+}\right)}=-\pi i \exp \left(i k_{0} r\right) \tag{4.25}
\end{equation*}
$$

Combining all these results we get

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)=\exp \left(i k_{0} r\right) / 4 \pi r \tag{4.26}
\end{equation*}
$$

which was the same result as we found using the retarded contour in Sec. 2 .

### 1.4.4 WETRICE REPRESENTATION

An alternative representation of spherical waves expanded in cylindrical waves can be found by expanding in the vertical ( $k_{z}$ ) wavenumber
rather than the horizontal wavenumber as in the Sommerfeld representation. Starting with (4.21) we don't do the $k_{z}$-integral. Instead do the $k_{x}$ and $k_{y}$ integrals using cylindrical symmetry. Dsing the definition

$$
\begin{equation*}
k_{z}=\left(k_{0}^{2}-k_{z}^{2}\right)^{1 / 2} \tag{4.27}
\end{equation*}
$$

## (4.21) is written as

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \int_{-\infty}^{\infty} \exp \left[i k_{z}\left(z^{-}-z^{\prime}\right)\right] d k_{z} \iint \frac{\exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-x_{t}^{\prime}\right)\right]}{k_{t}^{2}-K_{z}^{2}} d k_{z} d k_{y^{\prime}} \tag{4.28}
\end{equation*}
$$

The latter two integrals can be evalated using the cylindrical coordinate and Bessel function definitions in (4.12)-(4.15). The $\theta$-integral again yields a Bessel function so that we have

$$
\begin{equation*}
G^{(3)}\left(x, x^{\prime}\right)=\frac{1}{(2 \pi)^{2}} \int_{-\infty}^{\infty} e x p\left[i k_{z}\left(z-z^{\prime}\right)\right] d k_{z} \int_{0}^{\infty} \frac{k_{t} J_{0}\left(k_{t} r\right)}{k_{t}^{2}-k_{z}^{2}} d k_{t} \tag{4.29}
\end{equation*}
$$

We use (4.17) for the Bessel function and in the integral for $H_{0}^{(2)}$ we again rotate the contour $\left(k_{t} \rightarrow-k_{t}\right)$ so that using (4.18) we again have an integral only over $H_{0}^{(1)}$. We get

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)=\frac{1}{2(2 \pi)^{2}} \int_{-\infty}^{\infty} \exp \left[i k_{z}\left(z-z^{\prime}\right)\right] d k_{z} \int_{-\infty}^{\infty} \frac{k_{t} H_{0}^{(1)}\left(k_{t} \rho\right)}{k_{t}^{2}-k_{z}^{2}} d k_{t} \tag{4.30}
\end{equation*}
$$

We have shifted our poles so that $\operatorname{Im} Z_{z}>0$. We evaluate the $k_{t}$-integral using complex variables and closing the contour in the uhp. Explicitly writing out $K_{z}$ we get

$$
\begin{equation*}
G^{(3)}\left(x, x^{\prime}\right)=\frac{i}{8 \pi} \int_{-\infty}^{\infty} e x p\left[i k_{z}\left(z-z^{\prime}\right)\right] H_{0}^{(1)}\left[\left[k_{0+}^{2}-k_{z}^{2}\right]^{1 / 2} \rho\right] d k_{z} \tag{4.31}
\end{equation*}
$$

which is Feyrich's formpla, a one-dimensional integral representation in terms of the vertical wavenumber $k_{z}$. The result (4.31) is of ten quoted using (4.26) as

$$
\begin{equation*}
\frac{\exp \left[i k_{0}\left(\rho^{2}+z^{2}\right)^{1 / 2}\right]}{\left(\rho^{2}+z^{2}\right)^{1 / 2}}=\frac{i}{2} \int_{-\infty}^{\infty} e^{i \alpha z} H_{0}^{(1)}\left[\rho\left(k_{0}^{2}-\alpha^{2}\right)^{1 / 2}\right] d \alpha \tag{4.32}
\end{equation*}
$$

where $\rho, z$ are real, $r=\left(\rho^{2}+z^{2}\right)^{1 / 2}$ and $0 \leq$ arg $\left(k_{0}^{2}-\alpha^{2}\right)^{1 / 2}<\pi$.

### 1.4.5 PLANE-WAVE DECOMPOSITION OR $0^{(2)}$

It is possible to derive two meyl-type representations for the Helmholtz Green's function in two dimensions. In two dimensions the latter satisfies the Helmholtz equation

$$
\begin{equation*}
\left(\nabla_{2}^{2}+k_{0}^{2}\right) G^{(2)}\left({\underset{\sim}{x}}_{t} \cdot{\underset{\sim}{x}}_{t}^{\prime}\right)=-\delta\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right) \tag{4.33}
\end{equation*}
$$

We Fourier transform the equation by multiplying by

$$
\iint \exp \left(-i k_{t} \cdot{\underset{\sim}{x}}_{t}\right) d{\underset{\sim}{x}}_{t}
$$

where $\underset{\sim}{k_{t}}=\left(\mathbf{k}_{x}, \mathbf{k}_{\mathbf{y}}\right)$ is the Fourier transform variable. The result is the mixed representation

$$
\begin{equation*}
\tilde{G}^{(2)}\left({\underset{\sim}{k}}_{t} \cdot{\underset{\sim}{x}}_{t}^{\prime}\right)=\exp \left(-i \underset{\sim}{k} \cdot{\underset{\sim}{x}}_{t}^{\prime}\right)\left(k_{t}^{2}-k_{0}^{2}\right)^{-1} . \tag{4.34}
\end{equation*}
$$

The inverse transform thus becomes

$$
\begin{align*}
& G^{(2)}\left(\underset{\sim}{x} t^{\prime}{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{(2 \pi)^{2}} \iint \exp \left(i k_{t} \cdot{\underset{\sim}{x}}\right) \tilde{G}^{(2)}\left({\underset{\sim}{k}}_{t} \cdot{\underset{\sim}{x}}_{t}^{\prime}\right) d{\underset{\sim}{t}} \quad .  \tag{4.35}\\
& =\frac{1}{(2 \pi)^{2}} \iint \frac{\exp \left[i{\underset{\sim}{t}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)\right]}{\left(k_{x}-K_{y}\right)\left(k_{x}+Z_{y}\right)} d z_{t} \quad . \tag{4.36}
\end{align*}
$$

where

$$
\begin{equation*}
k_{y}=\left(k_{0}^{2}-k_{y}^{2}\right)^{1 / 2} \tag{4.37}
\end{equation*}
$$

We expressed the integrand of (4.36) in such a way as to do the $\mathbf{k}_{x^{-}}$ integration using complex variable techniques just as we did the $k_{z}$ integration in (4.9). The result is

$$
\begin{equation*}
G^{(2)}\left(\underset{\sim}{x}, x^{\prime}\right)=\frac{\pi i}{(2 \pi)^{2}} \int_{-\infty}^{\infty} \frac{\exp \left[i k_{y}\left(y-y^{\prime}\right)+i K_{y}\left|x-x^{\prime}\right|\right]_{y}}{K_{y}} \tag{4.38}
\end{equation*}
$$

Alternatively we could define

$$
\begin{equation*}
k_{x}=\left(k_{0}^{2}-k_{x}^{2}\right)^{1 / 2} \tag{4.39}
\end{equation*}
$$

so that the denominator of (4.36) becomes

$$
\left(\mathbf{k}_{\mathbf{y}}-\mathbf{K}_{\mathbf{x}}\right)\left(\mathbf{k}_{\mathbf{y}}+\mathbf{K}_{\mathbf{x}}\right)
$$

and the obvious choice is to carry ont the $k_{y}$-integration. The result is

$$
\begin{equation*}
G^{(2)}\left(x_{t}, x_{i}^{\prime}\right)=\frac{\pi i}{(2 \pi)^{2}} \int_{-\infty}^{\infty} \frac{\exp \left[i k_{x}\left(x-x^{\prime}\right)+i K_{x}\left|y^{\prime} y^{\prime}\right|\right]}{K_{x}} d k_{x} \tag{4.40}
\end{equation*}
$$

The representations (4.38) and (4.40) are both Weyl-type representations for
the two-dimensional Green's function. Their use depends on exploiting the geometry of the particular application.

### 1.4.6 BXPLICIT BVALUATION OF $\mathrm{G}^{(2)}$

Osing (4.34) and (4.35) we can explicitly evaluate $G^{(2)}$ using cy lindrical coordinates and complex integration. We have

$$
\begin{equation*}
G^{(2)}\left({\underset{\sim}{x}}_{t}, x_{t}^{\prime}\right)=\frac{1}{(2 \pi)^{2}} \iint \frac{\exp \left[i{\underset{\sim}{c}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)\right]}{k_{t}^{2}-k_{0}^{2}} d k_{x} d k_{y} \tag{4.41}
\end{equation*}
$$

In cylindrical coordinates we have that

$$
\begin{aligned}
& d k_{x} d k_{y}=k_{t} d k_{t} d \theta
\end{aligned}
$$

The $\theta$-integration yields $2 \pi J_{0}\left(k_{t} r\right)$, and we replace $J_{0}$ using (4.17). The integral involving $H_{0}^{(2)}$ is rewritten using (4.18) so that we only have an integral over $H_{0}^{(1)}$. It is

$$
\begin{equation*}
G^{(2)}\left(x_{t}, x_{t}^{\prime}\right)=\frac{1}{4 \pi} \int_{-\infty}^{\infty} \frac{H_{0}^{(1)}\left(k_{t} r\right)}{k_{t}^{2}-k_{0}^{2}} k_{t} d k_{t} \tag{4.43}
\end{equation*}
$$

Since $H_{0}^{(1)}$ behaves like an outgoing wave, we evalnate (4.43) using complex integration by closing the contour in the upper half plane. The poles are shifted by $k_{0} \rightarrow k_{0}+i \varepsilon$. The result is the cylindrical Green's function

$$
\begin{equation*}
G^{(2)}\left(x_{t} \cdot x_{t}^{\prime}\right)=(i / 4) H_{0}^{(1)}\left(k_{0} \rho\right) \tag{4.44}
\end{equation*}
$$

It is most useful for problems in cylindrical coordinates which have no angalar dependence.

### 1.4.7 EXPLICIT RELATION BETVEEN $\mathbf{G}^{(3)}$ nd $\mathbf{G}^{(2)}$

In three dimensions we have that

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \iiint \frac{\exp \left[i \underset{\sim}{i k} \cdot\left(\underset{\sim}{x-x^{\prime}}\right)\right]}{k^{2}-k_{0}^{2}} d k_{x} d k_{y} d k_{z} \tag{4.45}
\end{equation*}
$$

Break up these three integrals into a $k_{z}$-integral and a two-dimensional transverse integral as

$$
\begin{align*}
G^{(3)}\left(\underset{\sim}{x}, \mathbb{x}^{\prime}\right)= & \frac{1}{2 \pi} \int \exp \left[i k_{z}\left(z-z^{\prime}\right)\right] d k_{z} \cdot \\
& \frac{1}{(2 \pi)^{2}} \iint \frac{\exp \left[i k_{t} \cdot\left({\underset{\sim}{\sim}}_{t}-z_{t}^{\prime}\right)\right]}{k_{t}^{2}-\left[k_{0}^{2}-k_{z}^{2}\right]} d k_{z} d k_{y} \tag{4.46}
\end{align*}
$$

The latter two integrals are just (4.41) but with $k_{0}^{2}$ replaced by $K_{z}^{2}=k_{0}^{2}-k_{z}^{2}$. Dsing (4.44) we thins have that

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)=\frac{i}{8 \pi} \int_{-\infty}^{\infty} \exp \left[i k_{t}\left(z-z^{0}\right)\right] H_{0}^{(1)}\left[\left[z_{0}^{2}-k_{t}^{2}\right]^{1 / 2} \rho\right] d k_{z} \tag{4.47}
\end{equation*}
$$

which is just Meyrich's formula (4.31).

### 1.4.8 TARER-DIMRNSIONAL REPRESENTATIONS IN THE HALR-PLANB

In the previous sections we presented several integral representations for the Green's functions. We placed no restrictions on the regions of validity of these representations, and consequently they are valid in all space. Here we treat representations valid in one or the other half space.

We have singled out the z-direction as special, and we define the halfplanes using this. The Weyl representation for the retarded Green's function is from (4.11)

If we restrict the region to $z-z^{\prime} \geq 0$ so that the absolute value can be dropped we write the result as a three-dimensional integral as (the + sign indicates the region $z-z^{\prime} \geq 0$ ).

$$
\begin{equation*}
\left[G_{R}^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)\right]_{+}=(2 \pi)^{-3} \iiint A_{R}^{+}(\underset{\sim}{k}) \exp \left[i \underset{\sim}{k} \cdot\left(\underset{\sim}{x}-x_{\sim}^{\prime}\right)\right] d \underline{\sim} \tag{4.49}
\end{equation*}
$$

where the amplitude function is defined by

$$
\begin{equation*}
A_{R}^{+}(k)=(\pi i / K) \delta\left(k_{z}^{-K}\right) \tag{4.50}
\end{equation*}
$$

The advantage of this representation is in a three-dimensional problem where however the boundary is planar.

Similarly, a representation for $z-z^{\prime} \leq_{0}$ can be written as

$$
\begin{equation*}
\left[G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)\right]_{-}=(2 \pi)^{-3} \iiint A_{R}^{-}(\underset{\sim}{k}) \exp \left[\underset{\sim}{i} \underset{\sim}{x} \cdot\left(\underset{\sim}{x}-x^{\prime}\right)\right] d \underset{\sim}{k} \tag{4.51}
\end{equation*}
$$

where the amplitude is defined by

$$
\begin{equation*}
\overline{A_{R}^{-}}(\underline{z})=(\pi i / K) \delta\left(k_{z}+K\right) \tag{4.52}
\end{equation*}
$$

We can combine the two representations (4.49) and (4.51) to yield a quasi-three-dimensional Fourier representation where however the amplitude is
spatially dependent viz.

$$
\begin{equation*}
G_{R}^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)=(2 \pi)^{-3} \iiint A_{R}\left(z-z^{\prime}, k\right) \exp \left[i \underset{\sim}{i k} \cdot\left(\underset{\sim}{x-x^{\prime}}\right)\right] d k, \tag{4.53}
\end{equation*}
$$

and is given by

$$
\begin{align*}
A_{R}\left(z-z^{\prime}, k\right) & =\theta\left(z-z^{\prime}\right) A_{R}^{+}(k)+\theta\left(z^{\prime}-z\right) A_{R}^{-}(k) \\
& =(\pi i / K)\left[\theta\left(z-z^{\prime}\right) \delta\left(k_{z}-K\right)+\theta\left(z^{\prime}-z\right) \delta\left(k_{t}+K\right)\right] . \tag{4.54}
\end{align*}
$$

We could al so derive a Weyl-representation for the advanced Green's function. It is
with analogous representations for $z-z^{\prime} \geq 0$ and $z-z^{\prime} \leq 0$ given by

$$
\begin{equation*}
\left[G_{A}^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)\right]_{ \pm}=(2 \pi)^{-3} \iiint_{a}^{ \pm}(\underset{\sim}{k}) \exp \left[i \underset{\sim}{k} \cdot\left(\underset{\sim}{x}-x^{\prime}\right)\right] d \underset{\sim}{k} \tag{4.56}
\end{equation*}
$$

where

$$
\begin{equation*}
A_{a}^{ \pm}(k)=-(\pi i / K) \delta\left(k_{z}^{ \pm} K\right), \tag{4.57}
\end{equation*}
$$

with the full three-dimensional representation given by

$$
\begin{equation*}
G_{A}^{(3)}\left(x, x^{\prime}\right)=(2 \pi)^{-3} \iiint_{a}\left(z-z^{\prime}, \frac{k}{\sim}\right) e x p\left[i \underset{\sim}{x} \cdot\left(\underset{\sim}{x-x^{\prime}}\right)\right] d z \tag{4.58}
\end{equation*}
$$

where the amplitude is

$$
\begin{equation*}
A_{a}\left(z-z^{\prime}, \frac{k}{\sim}\right)=-(\pi i / K)\left[\theta\left(z-z^{\prime}\right) \delta\left(k_{z}+K\right)+\theta\left(z+z^{\prime}\right) \delta\left(k_{z}-K\right)\right] \tag{4.59}
\end{equation*}
$$

From these representations we can also compate the representation for the principal value using

$$
\begin{equation*}
G_{P}^{(3)}\left(x, x^{\prime}\right)=1 / 2\left[G_{R}^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)+G_{A}^{(3)}\left(x, x^{\prime}\right)\right], \tag{4.60}
\end{equation*}
$$

so that

$$
\begin{equation*}
G_{P}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=(2 \pi)^{-3} \iiint A_{p}\left(z-z^{\prime} \cdot k\right) \exp \left[i \underset{\sim}{i k} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right] d k . \tag{4.61}
\end{equation*}
$$

where

$$
\begin{align*}
A_{P}\left(z-z^{\prime}, k\right) & =1 / 2\left[A_{R}\left(z-z^{\prime}, \underline{\sim}\right)+A_{a}\left(z-z^{\prime}, k\right)\right]  \tag{4.62}\\
& =(\pi i / 2 K) \operatorname{sgn}\left(z-z^{\prime}\right)\left[\delta\left(k_{z}-K\right)-\delta\left(k_{z}+K\right)\right] . \tag{4.63}
\end{align*}
$$

### 1.5 ANALTTIC PROPERTIRS OF THE GRRRN'S PDNCTIONS

### 1.5.1 ANALITIC PROPBRTIES OR $G_{R}{ }^{(3)}$

We discuss the properties of $G_{R}^{\left({ }^{(3)}\right.}$ using distributions. We have that

$$
\begin{equation*}
G_{R}^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)=\frac{\exp \left[i k_{0}\left|x-x_{\sim}^{\prime}\right|\right]}{4 \pi\left|x-x^{\prime}\right|} \tag{5.1}
\end{equation*}
$$

We begin with the Weyl representation presented in Sec. 4. Here we have
$K=\left[k_{0}^{2}-k_{t}^{2}\right]$ and $k_{t}^{2}=k_{x}^{2}+k_{y}^{2}$. It is

$$
\begin{equation*}
G_{R}^{(3)}\left(\underset{\sim}{x}-x_{\sim}^{\prime}\right)=\frac{\pi i}{(2 \pi)^{3}} \iint \frac{e x p\left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)+i K_{t}\left|z_{-z}^{\prime}\right|\right]}{K_{t}} d k_{t} . \tag{5.2}
\end{equation*}
$$

where ${\underset{\sim}{x}}^{\prime}$ is the source point and $x$ the receiver point. We keep the representation as a difference in these coordinates. We use the term $K_{+}=K+i \varepsilon$ to distinguish the square root having a positive imaginary part. The properties are as follows:

PROPRETY 1. $G_{R}{ }^{(3)}\left(z_{0}-X^{0}\right)$ is continnous as $z-z^{0} \rightarrow 0$.

The proof is obvious. There are two cases, when $z-z^{\prime}>0$ and $z-z^{\prime}<0$. Both 1 imits are the same. They are

$$
\underset{z-z^{\prime} \rightarrow 0^{+}}{1 i^{\prime}} G_{R}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)=\lim _{z^{\prime} \rightarrow 0^{-}} G_{R}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)
$$

and equating the 1 imits of (5.1) and (5.2) we get

$$
\begin{equation*}
\frac{e^{i k_{0} \rho}}{4 \pi \rho}=\frac{\pi i}{(2 \pi)^{3}} \iint \exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{z}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)\right] \frac{d \underline{k}_{t}}{K_{t}}, \tag{5.3}
\end{equation*}
$$

where $\rho=\left|x_{t}-x_{t}\right|$. Simply put the argament of the exponential in (5.2) vanishes independent of direction because of the absolute value.

## PROPBRTY 2. The first derivative in depth is discontinuous.

To prove this differentiate (5.2) with respect to $z$

$$
\partial_{z} G_{R}^{(3)}\left(\underset{\sim}{x}-\underset{\sim}{x}{ }^{\prime}\right)=\frac{-\pi s g n\left(z^{\prime} z^{\prime}\right)}{(2 \pi)^{3}} \iint \exp \left[i{\underset{\sim}{t}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-\underset{\sim}{x}{ }_{t}^{\prime}\right)+i K_{+}\left|z-z^{\prime}\right|\right] d k_{t} .
$$

Note that the factor $\mathbb{K}$ cancels in the integrand. In the limit as $z-z^{\prime} \rightarrow 0$ the exponent vanishes independent of direction bat the antisymmetric signum function remains. A1so, the integral for $z-z^{\prime}=0$ is just the two dimensional delta function multiplied by $(2 \pi)^{2}$. The result is

$$
\begin{aligned}
& \lim _{z-z^{\prime} \rightarrow 0} \partial_{z} G_{R}^{(3)}\left(\underset{\sim}{x}-\underset{\sim}{x}{ }^{\prime}\right)=\frac{-\pi}{(2 \pi)^{3}} \quad \underbrace{1 i m}_{z-z^{\prime} \rightarrow 0} \operatorname{sgn}\left(z-z^{\prime}\right)
\end{aligned}
$$

$$
\begin{align*}
& \lim _{z-z^{\prime} \rightarrow 0} \partial_{z} G_{R}^{(3)}\left(\underset{\sim}{x-x_{\sim}^{\prime}}\right)=-\frac{1}{2} \delta\left({\underset{\sim}{t}}_{t}-{\underset{\sim}{x}}^{\prime}\right)\left[\begin{array}{ll}
1 & z-z^{\prime} \rightarrow 0^{+} \\
-1 & z-z^{\prime} \rightarrow 0^{-}
\end{array} .\right. \tag{5.5}
\end{align*}
$$

Note that the 1 imits are independent of $k_{0}$, so the same discontinuous derivative behavior holds for static potential theory. Define:

$$
\begin{equation*}
\underset{z-z^{\prime} \rightarrow 0 \pm}{1 \operatorname{im}_{z}} \partial_{R} G^{(3)}\left(\underset{\sim}{x}-z^{\prime}\right)=\left[\partial_{z} G_{R}^{(3)}\left(\underset{\sim}{x} t^{\prime}{\underset{\sim}{x}}^{\prime}\right)\right]_{ \pm} \tag{5.6}
\end{equation*}
$$

$$
\left.\begin{array}{rl}
\text { disc }\left[\partial_{z} G_{R}^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right)\right.
\end{array}\right]_{z=z^{\prime}}=\left[\partial_{z} G_{R}^{(3)}\right]_{+}-\left[\partial_{z} G_{R}^{(3)}\right]_{-} .
$$

We show later that this is analogous to the discontinuous behavior for onedimensional Green's functions. Here we also have an additional twodimensional delta function.

PROPRRTY 3.- The transverse derivatives are continuous.
Define the transverse differential operator as

$$
\partial_{j t}=\left[\begin{array}{ll}
\partial / \partial x & j=1 \\
\partial / \partial y & j=2
\end{array}\right.
$$

Differentiating (5.2) we get

$$
\begin{equation*}
\partial_{j t} G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\frac{-\pi}{(2 \pi)^{3}} \iint \exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)+i K_{+}\left|z-z^{\prime}\right|\right] \frac{k_{j t}}{K_{+}} d k_{t} \tag{5.8}
\end{equation*}
$$

This again approaches a finite value in the 1 imit $z-z^{\prime} \rightarrow \mathbf{0}$ independent of direction and is

$$
\begin{equation*}
\lim _{z \rightarrow z^{\prime} \rightarrow 0} \partial_{j t} G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{z}}^{\prime}\right)=\frac{-\pi}{(2 \pi)^{3}} \iint \exp \left[i \underset{\sim}{k_{t}} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}_{t}^{\prime}\right)\right]{\underset{j}{k}}_{K_{t}}^{d z}{\underset{\sim}{t}} \tag{5.9}
\end{equation*}
$$

which can be evaluated directly by doing the integral or simply by noting that me can interchange the derivative and the 1 imiting process to yield

$$
\begin{equation*}
\lim _{z \rightarrow z^{\prime} \rightarrow 0} \partial_{j} G_{R}^{(3)}\left(\underset{\sim}{\left.x-z^{\prime}\right)}=\partial_{j t} \frac{e^{i k_{0} \rho}}{4 \pi \rho} \quad ; p=\left|x-x_{t}^{\prime}\right|\right. \tag{5.10}
\end{equation*}
$$

Next we want a representation for the full vector derivative. We will need this later to find the normal derivative of the function. For reasons
which will be clear later it is convenient to have this in three-dimensional integral form which is regularized, i.e. which has no derivative singularity, plus a singular term. Differentiate (5.2) (where $j=1,2,3$ and $\left.\partial_{3}=\partial / \partial_{z}\right)$

$$
\begin{align*}
& \partial_{j} G_{R}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right)=\frac{\pi i}{(2 \pi)^{3}} \iint\left[i k_{j t}+i \delta_{j 3} K_{t} \operatorname{sgn}\left(z-z^{\prime}\right)\right] . \\
& \cdot \exp \left[i{\underset{\sim}{k}}_{t} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}_{\prime}^{\prime}\right)+i K_{+}\left|z-z^{\prime}\right|\right] \frac{d k_{t}}{K_{+}} \quad . \tag{5.11}
\end{align*}
$$

Now regularize the singularity in the $z$-derivative as follows:

$$
\begin{align*}
& \partial_{j} G_{R}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)=\frac{\pi i}{(2 \pi)^{3}} \iint i k_{j t} e \operatorname{xp}\left[i \underset{\sim}{t} \underset{\sim}{ } \cdot(\underset{\sim}{x}-\underset{\sim}{x})+i K_{+}\left|z-z^{\prime}\right|\right] \frac{d k_{t}}{K_{+}} \\
& +i \delta_{j 3} \operatorname{sgn}\left(z-z^{\prime}\right) . \\
& \cdot\left[\frac { \pi i } { ( 2 \pi ) ^ { 3 } } \int \int d { \underset { \sim } { t } } \operatorname { e x p } [ i { \underset { \sim } { k } } _ { t } \cdot ( { \underset { \sim } { x } } _ { t } - { \underset { \sim } { x } } _ { t } ^ { \prime } ) ] \left[\exp \left(i K_{+}\left|z-z^{\prime}\right|-1\right]\right.\right. \\
& \left.+\frac{\pi i}{(2 \pi)^{3}} \quad(2 \pi)^{2} \delta\left({\underset{\sim}{t}}-{\underset{\sim}{z}}_{t}^{\prime}\right)\right] \quad . \tag{5.12}
\end{align*}
$$

In the singular term we subtracted and added the term 1 . This brought out the $\delta$ function explicitly. We could have subtracted any function of $z-z$ " which has the 1 imit 1 as $z-z^{\prime} \rightarrow 0$, as for example $\cos \left[k_{0}\left(z-z^{\prime}\right)\right]$. Regularization is not unique, and this subtraction, or any appropriate subtraction, is a regularization in the sense that the resulting integral is not singular.

Me next want to reintroduce the $k_{z}$-integration in (5.12). Recall that we eliminated this integration by evaluating it to derive the Weyl
representation. In the first integral it is simple to recover the $k_{z}$ integra1. We use the result that $\left(k_{0} \rightarrow k_{0}+i \varepsilon, K \rightarrow K+i \varepsilon\right)$

$$
\begin{equation*}
e^{i K\left|z-z^{\prime}\right|}=\frac{K}{\pi i} \int_{-\infty}^{\infty} \frac{\exp \left[i k_{z}\left(z-z^{\prime}\right)\right]}{k_{z}^{2}-K_{+}^{2}} d k_{z} \tag{5.13}
\end{equation*}
$$

In the second integral we use

$$
\begin{equation*}
\operatorname{sgn}\left(z-z^{\prime}\right)\left[\exp \left(i K\left|z-z^{\prime}\right|\right)-1\right]=\frac{1}{\pi i} \int_{-\infty}^{\infty} \frac{\exp \left[i k_{z}\left(z^{-} z^{\prime}\right)\right]}{k_{z}^{2}-K_{+}^{2}} p \frac{k^{2}}{k_{z}} d k_{z} \tag{5.14}
\end{equation*}
$$

which can al so be derived using residue calculus methods. The integrand in (5.14) has three poles, the one at $k_{z}=0$ evaluated using the principal value ( $P$ ), and the other two using the shifts $k_{0}+i \varepsilon$ or $K+i \varepsilon$. The result in (5.12) is, noting that $k_{z}^{2}-k^{2}=k^{2}-k_{0}^{2}$

$$
\begin{align*}
& \partial_{j} G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \iiint i k_{j t} \frac{\exp \left[i k \cdot\left(\frac{\left.x-x^{\prime}\right)}{\sim}\right]\right.}{k^{2}-k_{0}^{2}} d k \\
& +i \delta_{j 3} \iiint \frac{\exp \left[i k^{\cdot}\left(\underset{\sim}{x}-x^{\prime}\right)\right]}{k^{2}-k_{0}^{2}} P \frac{k^{2}}{k_{z}} d k \\
& -\frac{1}{2} \operatorname{sgn}\left(z-z^{\prime}\right) \delta\left({\underset{\sim}{t}}^{t}-\underset{\sim}{x^{\prime}}\right) \delta_{j} \quad . \tag{5.15}
\end{align*}
$$

Noting that the Fourier transform of $G_{R}^{(a)}$ is

$$
\begin{equation*}
\tilde{\mathbf{G}}_{\mathrm{R}}^{(3)}(\mathrm{k})=\left[\mathbf{k}^{2}-\mathbf{k}_{0+}^{2}\right]^{-1} \tag{5.16}
\end{equation*}
$$

with $k_{0} \rightarrow k_{0}+i \varepsilon$, we can write (5.15) as

$$
\begin{align*}
& -\frac{1}{2} \delta_{j 3} \operatorname{sgn}\left(z-z^{\prime}\right) \delta(\underset{\sim}{x}-\underset{\sim}{x}) \quad \text {, }
\end{align*}
$$

where

$$
\begin{equation*}
P_{j}(\underset{\sim}{k})=2 i\left[k_{j t}+\delta_{j 3} P\left[\frac{\mathbb{K}^{2}}{k_{z}}\right]\right] \text {. } \tag{5.18}
\end{equation*}
$$

The integral term is not singular. (In fact note that if we set $z^{\prime} z^{\prime}=0$ in the integral we get that the $j=3$ term vanishes since the resulting integrand is an odd function of $k_{z^{\prime}}$ ) It is a Cauchy principal value integral. The subtraction of the term 1 has led to this. Subtraction of another term will lead to an alternate principal value integral. The full discontinuity is proportional to the $j=3$ term, i.e.

$$
\begin{equation*}
\left[\partial_{j} G_{R}^{(3)}\left(\underset{\sim}{x}-x^{\prime}\right)\right]_{+}-\left[\partial_{j} G_{R}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right]_{-}=-\delta_{j 3} \delta\left({\underset{\sim}{t}}_{t}-x_{i}^{\prime}\right), \tag{5.19}
\end{equation*}
$$

Where the + and - signs refer to the 1 imits as $z-z^{\prime}$ approaches zero from positive or negative values respectively. An alternative way of writing (5.17) is

$$
\begin{equation*}
\partial_{j} G_{R}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{2} R_{j}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)-\frac{1}{2} \delta_{j}{ }_{3} \operatorname{sgn}\left(z-z^{\prime}\right) \delta\left(\underset{\sim}{x}-\underset{\sim}{x}{ }_{t}^{\prime}\right), \tag{5.20}
\end{equation*}
$$

where

$$
\begin{equation*}
R_{j}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \iiint \exp \left[i \underset{\sim}{k} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right]{\underset{\mathrm{G}}{R}}_{(3)}^{(k)} P_{j}(k) d k \quad, \tag{5.21}
\end{equation*}
$$

is the regular part of the derivative term. We can thus set both vectors $x$
and ${\underset{\sim}{x}}^{\prime}$ onto the surface in $\mathbb{R}_{j}$ and have a well defined limit.

### 1.5.2 ANALYTIC PROPBRTIES OR G $\AA^{3}$ )

The full three-dimensional Fourier representation for $G^{(x)}\left(x^{\prime} x^{\prime}\right)$ is given by

$$
\begin{equation*}
G^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \iint \exp \left[i{\underset{\sim}{x}}_{t}\left({\underset{\sim}{t}}_{t}-x_{t}^{\prime}\right)\right] d{\underset{\sim}{t}}_{t} \int \frac{\exp \left[i k_{z}\left(z-z^{\prime}\right)\right]}{\left(k_{z}-K_{+}\right)\left(k_{z}+K_{+}\right)} d k_{z} \tag{5.22}
\end{equation*}
$$

where $\quad K=\left[k_{0}^{2}-k_{t}^{2}\right]^{1 / 2}$ and $k_{t}^{2}=k_{x}^{2}+k_{y}^{2}$. The retarded Green's function was computed by shifting these poles using $\mathbf{k}_{0} \rightarrow \mathbf{k}_{0+}$ or $\mathrm{K} \rightarrow \mathbf{K}_{+}$. The advanced Green's function is computed using the shift $\mathbf{k}_{\mathbf{0}} \rightarrow \mathbf{k}_{\mathbf{0}}$ - ie $=\mathbf{k}_{\mathbf{0}}$ or $K \rightarrow$ $K-i \varepsilon=K_{-}$. The singularity structure in the complex $\mathbf{k}_{\mathbf{z}}$-plane is thus


$$
z-z^{\prime}>0
$$

Fig. 1.3

We close the contour in the upper half plane for $z-z^{\prime}>0$ and in the lower half plane for $z-z^{\prime}<0$. The result is

$$
\begin{equation*}
\int \frac{\exp \left[i k_{z}\left(z-z^{\prime}\right)\right]}{\left(k_{z}-K\right)\left(k_{z}+K\right)} d k=\frac{-\pi i}{K} \exp \left[-\left.i K\right|_{z-z^{\prime}} \mid\right] \text {. } \tag{5.23}
\end{equation*}
$$

so that the Weyl representation for the advanced Green's function

$$
\begin{equation*}
G_{A}^{(3)}\left(x_{\sim}^{-x} x^{\prime}\right)=\frac{\exp \left(-i k_{0} r\right)}{4 \pi x} ; r=\left|x^{\prime} x^{\prime}\right| \text {, } \tag{5.24}
\end{equation*}
$$

is given by

$$
\begin{equation*}
G_{A}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right)=\frac{-\pi i}{(2 \pi)^{3}} \iint \exp \left[i k_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)-i K_{-}\left|z-z^{\prime}\right|\right] \frac{\underset{\sim}{\sim} t}{K_{-}} \tag{5.25}
\end{equation*}
$$

where we distinguish the square root term using $K_{-}=K$-ie. Its properties axe

1. G $A^{3}$ ) $\left(\underset{\sim}{x}-x^{\prime}\right)$ is continuous as $z-z^{\prime} \rightarrow 0$. The limits from both directions are the same. Note that from the functional form the 1 imit is

$$
\begin{equation*}
\frac{e^{-i k_{0} \rho}}{4 \pi \rho}=\frac{-\pi i}{(2 \pi)^{3}} \iint \exp \left[i{\underset{\sim}{t}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}^{\prime}\right)\right] \frac{d k_{t}}{K_{-}} . \tag{5.26}
\end{equation*}
$$

with $\rho=\left|{\underset{\sim}{x}} t^{-x} t^{\prime}\right|$. The square root distinguishes the contribution.
2. The first derivative in depth is discontinuous. Differentiating we get from (5.25)

$$
\begin{equation*}
\partial_{z} G_{A}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)=-\pi \frac{\operatorname{sgn}\left(z-z^{\prime}\right)}{(2 \pi)^{3}} \iint e \exp \left[i{\underset{\sim}{t}}_{t} \cdot\left({\underset{\sim}{x}}^{t}-{\underset{\sim}{x}}^{\prime}\right)+i k_{-}\left|z-z^{\prime}\right|\right] d k_{t} . \tag{5.27}
\end{equation*}
$$

which in the 1 imit as $z-z^{\prime} \rightarrow \mathbf{O}$ from the two directions is

which is the same discontinuity as the derivative of the retarded Green's function in (5.5).
3. The transverse derivatives are continuous. From (5.23) we have that

$$
\begin{equation*}
\lim _{z-z^{\prime} \rightarrow 0} \partial_{j t_{A}} G_{A}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right)=\frac{\pi}{(2 \pi)^{3}} \iint \exp \left[i{\underset{\sim}{t}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)\right] \frac{k_{j t}}{K_{-}} d k_{t}, \tag{5.29}
\end{equation*}
$$

which from the functional form (5.24) equals

$$
\begin{equation*}
\partial_{j t}\left[e \operatorname{xp}\left(-i k_{0} \rho\right) / 4 \pi \rho\right] \tag{5.30}
\end{equation*}
$$

Hence we can write the full vector derivative as

$$
\begin{align*}
\partial_{j} G_{A}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)= & \frac{-\pi i}{(2 \pi)^{3}} \iint\left[i k_{j t}-i \delta_{j} s_{-} \operatorname{sgn}\left(z-z^{\prime}\right)\right] \\
& \quad \exp \left[i{\underset{\sim}{t}}_{t} \cdot\left[{\underset{\sim}{t}}_{t}-x_{t}^{\prime}\right]-i K_{-}\left|z-z^{\prime}\right|\right] \frac{d k}{\sim_{t}} \tag{5.31}
\end{align*}
$$

Now regularize in the $z$ - derivative term. Rewrite (5.31) as

$$
\begin{align*}
& \partial_{j} G_{A}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right)=\frac{-\pi i}{(2 \pi)^{3}} \iint_{j t} \exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)-i K_{-}\left|z-z^{\prime}\right|\right] \frac{\underset{\sim}{d} t}{K_{-}} \\
& -i \delta_{j} \operatorname{sgn}\left(z-z^{\prime}\right) \cdot \\
& {\left[\frac { - \pi i } { ( 2 \pi ) ^ { 3 } } \int \int \operatorname { e x p } \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-\underset{\sim}{x}{ }_{t}^{\prime}\right]\left[\exp \left(-i K_{\_}\left|z-z^{\prime}\right|\right)-1\right] d k_{t}\right.\right.} \\
& \left.-\frac{\pi i}{(2 \pi)^{3}}(2 \pi)^{2} \delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right] \text {. } \tag{5.32}
\end{align*}
$$

Reintroduce the $k_{z^{-}}$integration (here $\mathbf{k}_{0} \rightarrow \mathbf{k}_{\mathbf{0}}-\mathrm{ie}$ and $\mathrm{K} \rightarrow \mathrm{K}-\mathrm{ie}=\mathrm{K}_{\mathbf{\prime}}$ ) using

$$
\begin{equation*}
\exp \left[-i \mathbb{K}\left|z-z^{\prime}\right|\right]=\frac{-\mathbb{R}}{\pi i} \int_{-\infty}^{\infty} d k_{z} \frac{\exp \left[i k_{z}\left(z-z^{\prime}\right)\right]}{k_{z}^{2}-\mathbb{K}_{-}^{2}} \tag{5.33}
\end{equation*}
$$

and

$$
\begin{equation*}
\operatorname{sgn}\left(z-z^{\prime}\right)\left[e \exp \left[-i K\left|z-z^{\prime}\right|\right]-1\right]=\frac{1}{\pi i} \int_{-\infty}^{\infty} \frac{\exp \left[i k_{z}\left(z-z^{\prime}\right)\right]}{k_{z}^{2}-K_{-}^{2}} p\left[\frac{k^{2}}{k_{z}}\right] d k_{z} \tag{5.34}
\end{equation*}
$$

Note the minus sign in front of (5.33) and the plus sign in front of (5.34). The principal value term in (5.34) is an odd function. Osing the Fourier transform of G $\mathbf{A}^{3}$ ) which is

$$
\begin{equation*}
\tilde{G}_{A}^{(3)}(k)=\left[k^{2}-k_{0-}^{2}\right]^{-1} \tag{5.35}
\end{equation*}
$$

where $k_{0} \rightarrow k_{0}$-i $\varepsilon$ the result using (5.33) and (5.34) in (5.32) is

$$
\begin{align*}
\partial_{j} G_{A}^{(3)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)= & \frac{1}{2} \frac{1}{(2 \pi)^{3}} \iiint \exp \left[\underset{\left.\underset{\sim}{k} \cdot\left(x-x^{\prime}\right)\right]}{ } \tilde{G}_{A}^{(3)}(k) P_{j}(k) d k\right. \\
& -\frac{1}{2} \operatorname{sgn}\left(z-z^{\prime}\right) \delta_{j}{ }_{j} \delta\left(\underset{\sim}{x}-\underset{\sim}{x}{ }_{t}^{\prime}\right) \quad . \tag{5.36}
\end{align*}
$$

Note that the result is similar to that for the retarded Green's function, (5.17), except that here the Fourier transform of the advanced Green's function is under the integral. $P_{j}(\underset{\sim}{k})$ is defined by (5.18).

### 1.5.3 REGULARIZATION OF $\partial_{j} \mathbf{G}_{\mathrm{P}}{ }^{(1)}$

Me have that

$$
G_{P}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right)=\frac{1}{2}\left[G_{R}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right)+G_{A}^{(3)}\left(\underset{\sim}{x}-x_{\sim}^{\prime}\right)\right],
$$

so that combining (5.17) and (5.36) we get

$$
\begin{align*}
\partial_{j} G_{P}^{(3)}\left(\underset{\sim}{x}-x^{\prime}\right)= & \frac{1}{2} \frac{1}{[2 \pi]^{3}} \iiint \exp \left[i \underset{\sim}{i k} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right] \tilde{G}_{P}^{(3)}(k) P_{j}(\underset{\sim}{k}) d \underset{\sim}{x} \\
& -\frac{1}{2} \operatorname{sgn}\left(z-z^{\prime}\right) \delta_{j 3} \delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right) \tag{5.38}
\end{align*}
$$

where

$$
\begin{equation*}
\tilde{G}_{P}^{(3)}(k)=P\left[k^{2}-k_{0}^{2}\right]^{-1} \tag{5.39}
\end{equation*}
$$

Hence the principal value Green's function has the same discontinuity as the retarded and advanced Green's functions.
1.5.4 REGOLARIZATION OR $\partial_{m}{ }^{\partial}{ }_{j} G_{R}^{(3)}$

We begin with the Weyl representation for the retarded Green's function in three dimensions

$$
\begin{equation*}
G_{R}^{(3)}(\underset{\sim}{x})=\frac{\pi i}{(2 \pi)^{3}} \iint \exp \left[i \underset{\sim}{i k_{t}} \cdot \underset{\sim}{x}+i K_{+}|z|\right] \frac{d k_{t}}{\mathbb{K}_{+}} . \tag{5.40}
\end{equation*}
$$

Differentiate this representation to get
and differentiate a second time to yield

$$
\begin{align*}
& \partial_{m^{2}}{ }_{j} G_{R}^{(3)}(\underset{\sim}{x})=\frac{\pi i}{(2 \pi)^{3}} \iint\left[2 i K_{+} \delta(z) \delta_{j^{3}} \delta_{m^{3}}+\right. \\
& \left.+i^{2}\left[k_{j t}+\delta_{j} z^{K} \operatorname{sgn}(z)\right]\left[k_{m t}+\delta_{m} X_{+} s g n(z)\right]\right] . \\
& \cdot \exp \left[i{\underset{\sim}{k}}_{t} \cdot \underset{\sim}{x}+i k_{+}|z|\right] \frac{d{\underset{\sim}{t}}^{( }}{K_{+}} \quad .
\end{align*}
$$

which can be witten as four terms

$$
\begin{align*}
& \partial_{m} \partial_{j} G_{R}^{(3)}(\underset{\sim}{x})=\frac{\pi i}{(2 \pi)^{3}} \iint \exp \left[i \underset{\sim}{k}{ }_{t} \cdot{\underset{\sim}{z}}_{t}+i K_{+}|z|\right] . \\
& {\left[\frac{i^{2} k_{j t^{k}} \mathrm{k}^{\prime}}{\mathrm{K}_{+}}+i^{2} \operatorname{sgn}(z)\left[k_{m t^{3}} \delta_{j}+k_{j t^{3}} \delta^{3}\right]\right.} \\
& \left.+i^{2} K_{+} \delta_{j^{3}} \delta_{m^{3}}+2 i \delta_{j^{3}} \delta_{m^{3}} \delta(z)\right] d k_{t} . \tag{5.43}
\end{align*}
$$

The first term is not singular, and using the relation

$$
\begin{align*}
\frac{\exp \left(i K_{+}|z|\right)}{\mathbf{K}_{+}} & =\frac{1}{\pi i} \int \frac{e \operatorname{xp}\left(i z k_{z}\right)}{k^{2}-k_{0+}^{2}} d k_{z} \\
& =\frac{1}{\pi i} \int e \operatorname{xp}\left(i z k_{z}\right) \tilde{G}_{R}^{(3)}(k) d k_{z} \tag{5.44}
\end{align*}
$$

we can write it as a three dimensional integral

$$
\begin{equation*}
I_{1}=\frac{1}{(2 \pi)^{3}} \iiint i^{2} k_{j t^{\prime}} k_{m} \exp (i \underset{\sim}{k} \cdot \underset{\sim}{x}){\underset{R}{(3)}}_{(k)}^{(k k} \tag{5.45}
\end{equation*}
$$

The second term containing the $s g n(z)$ function has a possible delta function singularity in the 1 imit as $z \rightarrow 0$. We regularize it as follows. First rewrite it as

$$
\begin{gather*}
I_{z}=\frac{\pi i}{(2 \pi)^{3} \iint i^{2}} \operatorname{sgn}(z)\left[k_{m t} \delta_{j 3}+k_{j t} \delta_{m^{3}}\right] \exp \left(i \xi_{t} \cdot{\underset{\sim}{t}}_{t}\right)  \tag{5.46}\\
\cdot\left[\exp \left(i K_{+}|z|\right)-1+1\right] d k_{t}
\end{gather*}
$$

The term involving the +1 in the bracket can be written as the derivatives of a two dimensional delta function. The remaining term can be written as a three dimensional integral using the relation (5.14) as

$$
\begin{equation*}
\operatorname{sgn}(z)\left[\exp \left(i K_{+}|z|\right)_{-1}\right]=\frac{1}{\pi i} \int \exp \left(i k_{z} z\right) \tilde{G}_{R}^{(3)}(k) P\left[\frac{k_{+}^{2}}{k_{z}}\right] d k_{z} \tag{5.47}
\end{equation*}
$$

The result is

$$
\begin{align*}
I_{2}= & -\frac{1}{2} s g n(z)\left[\delta_{j^{3}} \partial_{m t}+\delta_{m 3} \partial_{j t}\right] \delta(\underset{\sim}{x}) \\
& +\frac{1}{(2 \pi)^{3}} \iiint i^{2}\left[k_{m t} \delta_{j 3}+k_{j 3} \delta_{m 3}\right] P\left[\frac{K_{+}^{2}}{k_{z}}\right] e \operatorname{xp}(i \underset{\sim}{k} \cdot \underset{\sim}{x}) G_{R}^{(3)}(k) d \underset{\sim}{k}{ }_{i 5} \tag{5.48}
\end{align*}
$$

The third term in (5.43) is not singular. Using the relation

$$
\begin{equation*}
{K_{+}}^{e} \operatorname{xp}\left(i K_{+}|z|\right)=\frac{\mathbf{K}_{+}^{2}}{\pi i} \int \exp \left(i k_{z} z\right) \tilde{G}_{R}^{(3)}(k) d k_{z} \tag{5.49}
\end{equation*}
$$

we can write it as a three dimensional integral

$$
\begin{equation*}
I_{3}=\frac{1}{(2 \pi)^{3}} \iiint i^{-2}{K_{+}^{2}}_{\tilde{G}_{R}^{(3)}(k) \exp (i \underset{\sim}{x} \cdot \underline{z}) d \underset{\sim}{k}} \delta_{j^{3}} \delta_{m^{3}}, \tag{5.50}
\end{equation*}
$$

which can be written using the identity

$$
\begin{align*}
\mathbf{K}_{+}^{2} \tilde{G}_{R}^{(3)}(k) & =\left(k_{0+}^{2}-k_{t}^{2}\right) /\left(k^{2}-k_{0+}^{2}\right) \\
& =\left[k_{0+}^{2}-\left(k_{t}^{2}+k_{z}^{2}\right)+k_{z}^{2}\right] /\left(k^{2}-k_{0+}^{2}\right) \\
& =-1+k_{z}^{2} \tilde{G}_{R}^{(3)}(k) \tag{5.51}
\end{align*}
$$

to yield

$$
\begin{equation*}
I_{3}=\frac{1}{(2 \pi)^{2}} \iiint \int_{\sim}^{d k} i^{2} k_{z}^{2} \tilde{G}_{R}^{(3)}(k) \exp (i k \cdot x) \delta_{j 3} \delta_{m 3}+\delta_{j 3} \delta_{m 3} \delta(x) . \tag{5.52}
\end{equation*}
$$

The fourth term in (5.43) is a delta function. It is

$$
\begin{align*}
I_{4} & =\frac{\pi i}{(2 \pi)^{3}} \iiint 2 i \delta_{j{ }^{3}} \delta_{m^{3}} \delta(z) \exp \left[\underset{\sim}{i k}{ }_{t}{\underset{\sim}{x}}_{t}+i K_{+}|z|\right] d{\underset{\sim}{k}}_{t} \\
& =-\delta_{j_{3}{ }_{m} \delta^{3} \delta(\underset{\sim}{x})} . \tag{5.53}
\end{align*}
$$

The result (5.43) is given by the sum of $I_{1}$ thru $I_{4}$ from (5.45), (5.48), (5.52) and (5.53). The result can be written

$$
\begin{equation*}
\partial_{m} \partial_{j} G_{R}^{(3)}(\underset{\sim}{x})=-\frac{1}{2} R_{m j}(\underset{\sim}{x})-\frac{1}{2} s g n(z)\left[\delta_{j 3} \partial_{m t}+\delta_{m 3} \delta_{j t}\right] \delta\left({\underset{\sim}{x}}_{t}\right) . \tag{5.54}
\end{equation*}
$$

where $R_{m j}(x)$ is the regalar part of this mixed second derivative given by

$$
\begin{equation*}
R_{m j}(\underset{\sim}{x})=\frac{1}{(2 \pi)^{3}} \iiint \exp (i \underset{\sim}{k} \cdot \underset{\sim}{x}) G_{R}^{(3)}(k) P_{m j}(\underset{\sim}{k}) \underset{\sim}{d k} . \tag{5.55}
\end{equation*}
$$

and where

$$
\begin{equation*}
\frac{1}{2} P_{m j}\left(\frac{k}{\sim}\right)=k_{m t} k_{j t}+\left[k_{m t} \delta_{j j^{3}}+k_{j t} \delta_{m^{3}}\right] p\left[\frac{K^{2}}{k_{z}}\right]+\underline{k}_{z^{2} \delta_{m}^{3}} \delta_{j^{3}} \tag{5.56}
\end{equation*}
$$

Note that from this representation it is easy to show that

$$
\begin{equation*}
\left(\partial_{1} \partial_{1}+\partial_{2} \partial_{2}+\partial_{3} \partial_{3}\right) G_{R}^{(3)}(\underset{\sim}{x})=-\frac{1}{(2 \pi)^{3}} \iiint \exp (i \underset{\sim}{k} \cdot \underset{\sim}{x}){\underset{G}{R}}_{(3)}^{(k) k^{2} d k}, \tag{5.57}
\end{equation*}
$$

where $k^{2}=k_{z}^{2}+\mathbf{k}_{\mathbf{y}}^{2}+k_{z}^{2}$. From the identity

$$
k^{2} /\left(k^{2}-k_{0}^{2}\right)=1+k_{0}^{2} \tilde{G}_{R}^{(3)}(k)
$$

we see that

$$
\begin{equation*}
V^{2} G_{R}^{(3)}(\underset{\sim}{x})=-\delta(\underset{\sim}{x})-k_{0}^{2} \tilde{G}_{R}^{(3)}(k) \tag{5.58}
\end{equation*}
$$

which serves as a check on our results.
Some properties of this representation are obvious. The first is the symmetry of the derivative operation

$$
\begin{equation*}
\partial_{m} \partial_{j} G_{R}^{(3)}(\underset{\sim}{x})=\partial_{j} \partial_{m} G_{R}^{(3)}(\underset{\sim}{x}) \tag{5.59}
\end{equation*}
$$

Since $G_{R}^{(3)}$ is a homogeneous function we can exchange derivatives with respect to field and source coordinates up to minus sign so that

$$
\begin{equation*}
\partial_{m} \partial_{j} G_{R}^{(3)}\left(x_{N}-x^{\prime}\right)=\partial_{m}^{\prime} \partial j_{j}^{\prime} G_{R}^{(3)}\left(\underset{\sim}{x-x^{\prime}}\right) \tag{5.60}
\end{equation*}
$$

From (5.54) it is obvious that $P$ is symmetric

$$
\begin{equation*}
P_{m j}(\underline{k})=P_{j m}(\underset{\sim}{k}) \tag{5.61}
\end{equation*}
$$

Since $m$ and $j$ run from 1 to 3 we thus have at most six independent
components. Evaluating terms we get

$$
\begin{align*}
& P_{11}=2 k_{x}^{2}, P_{22}=2 k_{y}^{2}, P_{33}=2 k_{z}^{2} \\
& P_{12}=2 k_{x} k_{y} \\
& P_{13}=2 k_{x} P\left[\frac{k^{2}}{k_{z}}\right] \quad, \quad P_{23}=2 k_{y} p\left[\frac{\mathbf{k}^{2}}{k_{z}}\right] \tag{5.62}
\end{align*}
$$

Other constraints are possible. Note that

$$
\begin{equation*}
\mathbf{P}_{11} \mathbf{P}_{22}=\mathbf{P}_{12}^{2} \tag{5.63}
\end{equation*}
$$

and

$$
\begin{equation*}
\left[P_{13} / P_{23}\right]^{2}=P_{11} / P_{22} \tag{5.64}
\end{equation*}
$$

Finally, note that the jump discontinaity in the representation (5.54) occurs only in the off-diagonal components. The discontinuity across the surface is

$$
\begin{equation*}
\left[\partial_{m} \partial_{j} G_{R}^{(3)}\right]_{+}-\left[\partial_{m} \partial_{j} G_{R}^{(3)}\right]_{-}=-\left[\delta_{j 3} \partial_{m t}+\delta_{m 3} \partial_{j t}\right] \delta\left({\underset{\sim}{x}}_{t}\right) \tag{5.65}
\end{equation*}
$$

Note also that if we satially integrate these dipole terms by themselves the result is zero. For example

$$
\begin{equation*}
\int \partial_{j t} \delta\left({\underset{\sim}{x}}_{t}\right) \mathrm{d} \mathrm{x}_{\mathrm{t}}=0_{j t} \tag{5.66}
\end{equation*}
$$

but with an additional term in the integral we get for example

$$
\begin{equation*}
\iint f\left({\underset{\sim}{x}}_{t}\right) \partial_{j t} \delta\left({\underset{\sim}{x}}_{t}\right) d{\underset{\sim}{x}}_{t}=-\left.\partial_{j t} f\left({\underset{\sim}{x}}_{t}\right)\right|_{{\underset{\sim}{x}}_{t}}={\underset{\sim}{t}}_{t} . \tag{5.67}
\end{equation*}
$$

Note that if we have the difference of arguments we get

$$
\begin{aligned}
\partial_{m}^{\prime} \partial_{j}^{\prime} G_{R}^{(3)}(\underset{\sim}{x}-\underset{\sim}{x})= & -\frac{1}{2} R_{m j}\left({\underset{\sim}{x}}^{\prime}-\underset{\sim}{x}\right) \\
& -\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right)\left(\delta_{j z^{\prime}} \partial_{m t}^{\prime}+\delta_{m^{3}} \partial_{j \mathrm{j}}^{\prime}\right) \delta\left(\underset{\sim}{x} t^{\prime}-\underset{\sim}{x}\right)
\end{aligned}
$$

To differentiate on the second argument note that

$$
\begin{equation*}
\partial_{m}^{\prime} \partial_{j}^{\prime} G_{R}^{(3)}\left(\underset{\sim}{x} x^{\prime}-\underset{\sim}{x}\right)=\partial_{m} \partial_{j} G_{R}^{(3)}\left(x_{\sim}^{\prime}-\underset{\sim}{x}\right) \quad . \tag{5.69}
\end{equation*}
$$

by the homogeneity of the Green's function. Substituting this in (5.68) and Writing the partial derivatives in terms of the unprimed coordinate on the rhs we get a sign change in the latter term. The result is

$$
\begin{align*}
\partial_{m} \partial_{j} G_{R}^{(3)}(\underset{\sim}{x}-\underset{\sim}{x})= & -\frac{1}{2}{\underset{m j}{ }(\underset{\sim}{x}-\underset{\sim}{x})} \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right)\left(\delta_{j 3^{3}} \partial_{m t}+\delta_{m 3} \partial_{j t}\right) \delta\left(\underset{\sim}{x}{ }_{t}^{\prime}-\underset{\sim}{x}\right) \tag{5.70}
\end{align*} .
$$

### 1.6 ONE-DIMENSIONAL PROBLEUS

### 1.6.1 GRREN'S FUNCTION IN 1-DIMBNSION

We begin with a general second order linear differential equation with a delta function source term

$$
\begin{equation*}
\frac{d^{2} \phi}{d z^{2}}+p(z) \frac{d \phi}{d z}+q(z) \phi=-\delta\left(z-z^{\prime}\right) \tag{6.1}
\end{equation*}
$$

The function $\varnothing$ is thas the Green's function for this one-dimensional problem. The source point $z^{\prime}$ is singled out and we have different solutions in the regions $z\rangle z^{\prime}$ and $z\left\langle z^{\prime}\right.$. Thus the source point can be interpreted as introducing another boundary layer into the problem. We thus have two solutions and must say how they match at the layer interface. We assume that
(a) $P$ is continuous across the layer
(b) dø/dz has a discontinuity across the 1 ayer.

We use the continuity property to find the discontinnity as follows. First rewrite (6.1) in the form

$$
\begin{equation*}
\frac{d^{2} \phi}{d z^{2}}+\frac{d}{d z}[p(z) \phi]+\left[q(z)-p^{\prime}(z)\right] \phi(z)=0 \tag{6.2}
\end{equation*}
$$

Next integrate ( 6.2 ) across the 1 ayer from $z^{\prime \prime}-\varepsilon$ to $z^{\prime}+\varepsilon$ where $\varepsilon$ is small and shrinks to zero. For the second derivative term in (6.2) we get

$$
\begin{equation*}
\lim _{\varepsilon \rightarrow 0} \int_{z^{\prime}-\varepsilon}^{z^{\prime}+\varepsilon} \frac{d^{2} \phi}{d^{2} z} d z=\left.\frac{d \phi}{d z}\right|_{z^{\prime}-\varepsilon} ^{z^{\prime}+\varepsilon}=\frac{d \phi^{+}}{d z}-\frac{d \phi^{-}}{d z} \tag{6.3}
\end{equation*}
$$

where we have defined

$$
\begin{equation*}
\lim _{\varepsilon \rightarrow 0} \frac{d \phi}{d z}\left(z^{\prime} \pm \varepsilon\right)=\frac{d \phi^{ \pm}}{d z} . \tag{6.4}
\end{equation*}
$$

The second term on the 1 hs of (6.2) becomes

$$
\int_{z^{\prime}-\varepsilon}^{z^{\prime}+\varepsilon} \frac{d}{d z}[p(z) \phi(z)] d z=\left.p(z) \phi(z)\right|_{z^{\prime}-\varepsilon} ^{z^{\prime}+\varepsilon} \rightarrow 0
$$

which vanishes as $\varepsilon \rightarrow 0$ since $\rho$ is continuous and we assume $p$ is also continuons. It is not necessary to assume the latter in which case our result is

$$
\left(p^{+}-p^{-}\right) \phi\left(z^{\prime}\right)
$$

where

$$
p^{ \pm}=\lim _{\varepsilon \rightarrow 0} p\left(z^{\prime} \pm \varepsilon\right)
$$

The third term on the 1 hs of (6.2) is

$$
\int_{z^{\prime}-\varepsilon}^{z^{\prime}+\varepsilon}\left[q(z)-p^{\prime}(z)\right] \phi(z) d z
$$

Which vanishes in the 1 imit as $\varepsilon \rightarrow 0$ unless $q(z)$ or $p^{\prime}(z)$ are discontinuous. If $q(z)$ is discontinnous it becomes

$$
\left(q^{+}-q^{-}\right) \phi\left(z^{\prime}\right)
$$

where

$$
q^{ \pm}=\lim _{\varepsilon \rightarrow 0} q\left(z^{\prime} \pm \varepsilon\right)
$$

and if $p$ is discontinuous so that

$$
p(z)=p_{0}(z) \theta\left(z-z^{\prime}\right)+p_{i}(z) \theta\left(z^{\prime}-z\right)
$$

where $\theta$ is the step function

$$
\theta(x)=\left[\begin{array}{ll}
1 & x>0 \\
0 & x<0
\end{array}\right.
$$

we have that

$$
\begin{aligned}
p^{\prime}(z)= & p_{0}^{\prime}(z) \theta\left(z-z^{\prime}\right)+p_{1}^{\prime}(z) \theta\left(z-z^{\prime}\right) \\
& +p_{0}(z) \delta\left(z-z^{\prime}\right)-p_{1}(z) \delta\left(z^{-} z^{\prime}\right)
\end{aligned}
$$

so that we get for the result

$$
\begin{aligned}
\int_{z^{\prime}-\varepsilon}^{\prime}+\varepsilon & p^{\prime}(z) \phi(z) d z \\
& =\left.p(z) \phi(z)\right|_{z^{\prime}-\varepsilon} ^{z^{\prime}+\varepsilon}-\int_{z^{\prime}-\varepsilon}^{z^{\prime}+\varepsilon} p(z) \frac{d \phi}{d z} d z \\
& \rightarrow\left[p_{0}\left(z^{\prime}\right)-p_{1}\left(z^{\prime}\right)\right] \phi\left(z^{\prime}\right)-\left[p_{0}\left(z^{\prime}\right) \frac{d \phi}{d z}-p_{1}\left(z^{\prime}\right) \frac{d \phi}{d z}\right] \varepsilon
\end{aligned}
$$

The second term vanishes as $e \rightarrow 0$ provided neither of $d \phi^{ \pm} / d z$ is singular. Essentially all of these discontinuous properties merely complicate our al gebra and we drop them. That is, we assume $p, p^{\prime}$, and $q$ are continuous at the interface. The result is, from (6.3)

$$
\begin{equation*}
\frac{d \phi^{+}}{d z}-\frac{d \phi^{-}}{d z}=-1 \quad\left(z=z^{\prime}\right) \tag{6.5}
\end{equation*}
$$

where the -1 results from integrating the delta function. The continuity of $\phi$ is expressed as

$$
\begin{equation*}
\phi^{+}-\phi^{-}=0 \quad\left(z=z^{\prime}\right), \tag{6.6}
\end{equation*}
$$

and it is these latter two equations we use in the analysis. Now we must define the field boundary value problem. We do two examples.

## Eg. 1. INPINITE SPACE EXAMPLB

Here the boundary layer is the only finite boundary. Above that boundary 1 ayer we have solutions of the homogeneous version of (6.1) (no delta function term). For wave like solntions asymptotically we choose the solution which satisfies an outgoing radiation condition. To insure the se wave-1ike solutions $p(z)$ and $q(z)$ are required to have certain asymptotic properties. The simplest are that as $z \rightarrow \infty$

$$
p(z) \rightarrow 0 \text { and } q(z) \rightarrow \text { constant }>0
$$

Strictly speaking we al so have to require that $p(z)$ and $q(z)$ are monotonic functions. If there are any kinks in these profiles, waves can be trapped, and we must effectively introduce further 1 ayers into the problem. For simplicity we assume these properties are satisfied.

Thus, in the upper layer ( 0 ) where $z z^{\prime}$ the solution of the homogeneous version of (6.1) satisfying the outgoing radiation condition can be written 2 S

$$
\begin{equation*}
\phi_{0}(z)=A \mathbf{u}_{+}(z) \quad z>z^{\prime}, \tag{6.7}
\end{equation*}
$$

where $A$ is an unknown constant. Similarly in the lower (L) region (zくz) the solution satisfying the outgoing radiation condition is

$$
\begin{equation*}
\phi_{L}(z)=B u_{-}(z) \quad z<z^{\prime}, \tag{6.8}
\end{equation*}
$$

where $B$ is an unknown constant. Our conditions (6.5) and (6.6) are then satisfied by (6.7) and (6.8) provided that at $z=z^{\prime}$

$$
\begin{equation*}
A \frac{d u_{+}}{d z}-B \frac{d u_{-}}{d z}=-1, \tag{6.9}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{Au}_{+}-\mathrm{Bu} \mathbf{u}_{-}=0 \text {. } \tag{6.10}
\end{equation*}
$$

The se are coupled equations for $A$ and $B$ whose solution is

$$
\begin{equation*}
A=-\mathbf{n}_{-}\left(z^{\prime}\right) / W, \quad B=-\mathbf{u}_{+}\left(z^{\prime}\right) / W \text {, } \tag{6.11}
\end{equation*}
$$

where $W$ is the Mronskian

$$
\begin{equation*}
W=\mathbf{n}_{+}^{\prime} \mathbf{u}_{-}-\mathbf{u}_{+} \mathbf{u}_{-}^{\prime} . \tag{6.12}
\end{equation*}
$$

The fall solution can thus be written as

$$
\phi(z)=\left[\begin{array}{ll}
-u_{+}(z) u_{-}\left(z^{\prime}\right) / W & z>z^{\prime}  \tag{6.13}\\
-u_{+}\left(z^{\prime}\right) u_{-}(z) / W & z<z^{\prime} .
\end{array}\right.
$$

To specify the solution further we must know the eigenfunctions. As a simple example choose $p=0$ and $q=k_{0}^{2}$ in (6.1). Then the eigenfunctions of

$$
\begin{equation*}
\frac{d^{2} \phi}{d z^{2}}+k_{0}^{2} \phi=0 \tag{6.14}
\end{equation*}
$$

are either $\exp \left({ }^{\left(1 k_{0}\right.} \mathbf{z}\right)$ or $\left[\cos k_{0} z, \sin k_{0} z\right]$. Te choose the former set since they correspond to outgoing waves. We thas have that

$$
\begin{equation*}
u_{+}(z)=e^{i k_{0} z} ; u_{-}(z)=e^{-i k_{0} z} ; W=2 i k_{0} . \tag{6.15}
\end{equation*}
$$

so that

$$
\phi(z)=\left[\begin{array}{ll}
-\frac{e^{i k_{0}\left(z-z^{\prime}\right)}}{2 i k_{0}} & z>z^{\prime}  \tag{6.16}\\
-\frac{e^{-i k_{0}\left(z-z^{\prime}\right)}}{2 i k_{0}} & z\left\langle z^{\prime},\right.
\end{array}\right.
$$

or rewriting

$$
\begin{equation*}
\phi(z)=-\frac{1}{2 i k_{0}} e^{i k_{0}\left|z^{\prime}-z^{\prime}\right|}, \tag{6.17}
\end{equation*}
$$

which is our standard retarded one-dimensional Green's function. Note that we got the retarded Green's function because we chose outgoing radiation solutions.

Eg. 2. BODNDARY EXAMPLR
In this example we introduce upper and lower boundaries at a finite distance from the source plane. Geometrically we have that


Now we must choose our solutions as

$$
\begin{equation*}
\phi_{0}(z)=A_{+} \mathbf{n}_{+}^{(1)}(z)+A_{-} \mathbf{n}_{-}^{(1)}(z) \tag{6.18}
\end{equation*}
$$

incluaing both 1 inearly independent solutions $u^{(1)}$ in region 1 , and in the lower region 2

$$
\begin{equation*}
\phi_{L}(z)=B_{+} \mathbf{u}_{+}^{(2)}(z)+B_{-} \mathbf{u}_{-}^{(2)}(z) \tag{6.19}
\end{equation*}
$$

al so including both 1 inearly independent solutions in this region $u^{(2)}$. We have the same continuity and jump conditions at the interface as before.

Equations (6.5) and (6.6) become at $z=z^{\prime}$

$$
\begin{equation*}
\frac{d \phi_{0}}{d z}-\frac{d \phi_{L}}{d z}=-1 \tag{6.20}
\end{equation*}
$$

and

$$
\begin{equation*}
\phi_{\mathbf{D}}-\phi_{\mathbf{L}}=0 \tag{6.21}
\end{equation*}
$$

The se are two conditions on the four constants $A_{ \pm}$and $B_{ \pm}$. In addition there are boundary conditions at $z_{0}$ and $z_{L}$. Assume for simplicity that

$$
\begin{align*}
& \phi_{D}\left(z_{D}\right)=0  \tag{6.22}\\
& \frac{d \phi_{L}}{d z}\left(z_{L}\right)=0 \tag{6.23}
\end{align*}
$$

Again our formalism can accommodate much more complicated impedance type boundary conditions at these surfaces. We again take our previons example where $p=0$ and $q=k_{0}^{2}$ in (6.1). Because of the form of the boundary conditions it is convenient (but not necessary) to choose our eigenfanctions as

$$
\mathbf{u}_{ \pm}^{(1)}(z)=\left[\begin{array}{l}
\sin k_{0}\left(z-z_{0}\right)  \tag{6.24}\\
\cos k_{0}\left(z-z_{0}\right)
\end{array}\right.
$$

and

$$
\mathbf{u}_{ \pm}^{(2)}(z)=\left[\begin{array}{l}
\sin k_{0}\left(z-z_{L}\right)  \tag{6.25}\\
\cos k_{0}\left(z-z_{L}\right)
\end{array}\right.
$$

Using (6.18) and (6.22) yields $A_{-}=0$. Dsing (6.19) and (6.23) yields $B_{+}=0$. We thus have satisfied the boundary conditions with the fields

$$
\begin{equation*}
\phi_{0}(z)=A_{+} \sin \left[k_{0}\left(z-z_{0}\right)\right] . \tag{6.26}
\end{equation*}
$$

and

$$
\begin{equation*}
\phi_{L}(z)=B_{-} \cos \left[k_{0}\left(z-z_{L}\right)\right] \tag{6.27}
\end{equation*}
$$

We now must satisfy conditions (6.20) and (6.21) which become

$$
\begin{align*}
& A_{+} k_{0} \cos \left[k_{0}\left(z^{\prime}-z_{0}\right)\right]+B_{-} k_{0} \sin \left[k_{0}\left(z^{\prime}-z_{L}\right)\right]=-1  \tag{6.28}\\
& A_{+} \sin \left[k_{0}\left(z^{\prime}-z_{0}\right)\right]-B_{-} \cos \left[k_{0}\left(z^{\prime}-z_{L}\right)\right]=0, \tag{6.29}
\end{align*}
$$

whose solution is

$$
\begin{align*}
& A_{+}=\frac{\mathbf{u}_{-}^{(z)}\left(z^{\prime}\right)}{W}=\cos \left[k_{0}\left(z^{\prime}-z_{L}\right)\right] / W,  \tag{6.30}\\
& B_{-}=\frac{\mathbf{u}_{+}^{(1)}\left(z^{\prime}\right)}{W}=\sin \left[k_{0}\left(z^{\prime}-z_{u}\right)\right] /^{W}, \tag{6.31}
\end{align*}
$$

where

$$
\begin{equation*}
W=-k_{0} \cos \left[k_{0}\left(z_{L}-z_{u}\right)\right], \tag{6.32}
\end{equation*}
$$

so that

$$
\varphi(z)=\left[\begin{array}{cc}
\frac{-\sin \left[k_{0}\left(z-z_{0}\right)\right] \cos \left[k_{0}\left(z^{\prime}-z_{L}\right)\right]}{k_{0} \cos \left[k_{0}\left(z_{L}-z_{0}\right)\right]} & z>z^{\prime} \\
\frac{-\sin \left[k_{0}\left(z^{\prime}-z_{0}\right)\right] \cos \left[k_{0}\left(z-z_{L}\right)\right]}{k_{0} \cos \left[k_{0}\left(z_{L} z_{0}\right)\right]} & z<z^{\prime}
\end{array},\right.
$$

or

$$
\varphi(z)=\left[\begin{array}{cc}
\frac{\sin \left[k_{0}\left(z_{0}-z\right)\right] \cos \left[k_{0}\left(z^{\prime}-z_{L}\right)\right]}{k_{0} \cos \left[k_{0}\left(z_{L} z_{0} z_{0}\right)\right]} & z>z^{\prime}  \tag{6.33}\\
\frac{\sin \left[k_{0}\left(z_{0}-z^{\prime}\right)\right] \cos \left[k_{0}\left(z-z_{L}\right)\right]}{k_{0} \cos \left[k_{0}\left(z_{L}-z_{0}\right)\right]} & z<z^{\prime}
\end{array} .\right.
$$

### 1.6.2 SOLVABLE PRORILRS - INROMOGBNEOUS MRDIA

The examples in the previous development in this section were for homogeneous media. Here we develop a general method to find the eigenfunctions for various one-dimensionally inhomogeneous media. In a sense we do an inverse problem, first choosing the eigenfunctions and then finding the medium index of refraction.

We begin with a general linear second order differential equation

$$
\begin{equation*}
\frac{d^{2} h}{d x^{2}}+p(x) \frac{d h}{d x}+q(x) h(x)=0 \tag{6.34}
\end{equation*}
$$

whose solutions are assumed to be known in terms of special functions for example. Transform both independent and dependent variables as

$$
\begin{array}{lr}
\mathbf{x}=\mathbf{u}(z) & \mathbf{n}^{\prime}(z) \neq 0 \\
\mathbf{h}(\mathbf{x})=w(z) \phi(z) & w(z) \neq 0 \tag{6.36}
\end{array}
$$

so that the new equation on $\phi$ is given by

$$
\begin{equation*}
\frac{d^{2} \phi}{d z^{2}}+A(z) \frac{d \phi}{d z}+B(z) \phi(z)=0 \tag{6.37}
\end{equation*}
$$

where $A$ and $B$ are given by

$$
\begin{equation*}
A(z)=2 \frac{\nabla^{\prime}}{\nabla}-\frac{\mathbf{n}^{\prime \prime}}{\mathbf{u}^{\prime}}+\mathbf{n}^{\prime} P(z), \tag{6.38}
\end{equation*}
$$

and

$$
\begin{equation*}
B(z)=\frac{w^{\prime \prime}}{w}+\frac{w^{\prime}}{w}\left[u^{\prime} P(z)-\frac{u^{\prime \prime}}{\mathbf{u}^{\prime}}\right]+\left(u^{\prime}\right)^{2} Q(z) \tag{6.39}
\end{equation*}
$$

where we have defined

$$
\begin{equation*}
P(z)=p(n(z)) ; Q(z)=q(n(z)) \tag{6.40}
\end{equation*}
$$

To prove this note that

$$
\frac{d}{d x} h(x)=\frac{d z}{d x} \frac{d}{d z}[w(z) \phi(z)]=\frac{1}{u^{\prime}}\left[w^{\prime} \phi+w \phi^{\prime}\right]
$$

and

$$
\begin{aligned}
\frac{d^{2}}{d x^{2}} h(x) & =\frac{d z}{d x} \frac{d}{d x}\left[\frac{d f}{d x}\right] \\
& =\frac{1}{u^{\prime}}\left[-\frac{\mathbf{u}^{\prime \prime}}{\left(u^{\prime}\right)^{2}}\left[w^{\prime} \phi+w \phi^{\prime}\right]+\frac{1}{u^{\prime}}\left[w^{\prime \prime} \phi+2 w^{\prime} \phi^{\prime}+w \phi^{\prime \prime}\right]\right] .
\end{aligned}
$$

Combining these in (6.34) yields

$$
\begin{aligned}
\frac{\left[w \phi^{\prime \prime}+2 w^{\prime} \phi^{\prime}+w^{\prime \prime} \phi\right]}{\left(u^{\prime}\right)^{2}} & -\frac{u^{\prime \prime}}{\left(u^{\prime}\right)^{3}}\left(w \phi^{\prime}+w^{\prime} \phi\right) \\
& +\frac{P(z)}{u^{\prime}}\left[w \phi^{\prime}+w^{\prime} \phi\right]+Q(z) w \phi=0 .
\end{aligned}
$$

Multiplying by $\left(u^{\prime}\right)^{2}$ yields

$$
\begin{aligned}
w \phi^{\prime \prime}+2 w^{\prime} \phi+w^{\prime \prime} \phi & -\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}\left[w \phi^{\prime}+w^{\prime} \phi\right] \\
& +P(z) u^{\prime}\left[w \phi^{\prime}+w^{\prime} \phi\right]+Q(z)\left(u^{\prime}\right)^{2} w \phi=0
\end{aligned}
$$

Dividing by yields

$$
\begin{aligned}
\phi^{\prime \prime}+\phi^{\prime} & {\left[2 \frac{w^{\prime}}{w}-\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}+\mathbf{u}^{\prime} P(z)\right] } \\
& +\phi\left[\frac{w^{\prime \prime}}{w}-\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}} \frac{w^{\prime}}{w}+P(z) u^{\prime} \cdot \frac{w^{\prime}}{w}+Q(z)\left(u^{\prime}\right)^{2}\right]=0 .
\end{aligned}
$$

The basic idea of this development is as follows:

1. Choose $p(x)$ and $q(x)$ such that the differential equation (6.34) has known solutions in terms of special functions.
2. Choose $A(z)=0$ to find one of the functions or $w$ in terms of the other.
3. Then $B(z)$ is known in terms of one of the transformation functions from (6.38) and (6.39) as

$$
\begin{align*}
B(z) & =\frac{w^{\prime \prime}}{w}-2\left[\frac{w^{\prime}}{w}\right]^{2}+\left[u^{\prime}\right]^{2} Q(z)  \tag{6.41}\\
& =\frac{d}{d z}\left[\frac{w^{\prime}}{w}\right]-\left[\frac{w^{\prime}}{w}\right]^{2}+\left[u^{\prime}\right]^{2} Q(z),
\end{align*}
$$

with say $u$ known as a function of $w$ from the condition $A(z)=0$.
4. Then choose the second transformation function, w say, so that $B(z)$ is known.
5. Write $B(z)=k_{0}^{2} n^{2}(z)$ where $n(z)$ is the index of refraction in the inhomogeneous media described by the differential equation ( $A=0$ and $B$ $=k_{0}^{2} n^{2}(z)$ from (6.37)) as

$$
\begin{equation*}
\frac{d^{2} \phi}{d z^{2}}+k_{0}^{2} n^{2}(z) \phi=0 \text {. } \tag{6.42}
\end{equation*}
$$

We can integrate the equation $A(z)=0$ which is

$$
\begin{aligned}
& u^{\prime} P(z)+2 \frac{w^{\prime}}{w}-\frac{u^{\prime \prime}}{u^{\prime}}=0, \\
& \frac{d}{d z} \ln \left[u^{\prime}\right]-2 \frac{d}{d z} \ln [w]=a^{\prime} P(z),
\end{aligned}
$$

and

$$
\begin{equation*}
\frac{d}{d z} \ln \left[\frac{u^{\prime}}{w^{2}}\right]=u^{\prime} P(z) \tag{6.43}
\end{equation*}
$$

However, rather than integrate the equation for a general $P(z)$ we will find that the differential form is most usef al as we choose particular values of $p(x)$ and hence $P(z)$.

## Eg. 1. HYPRRGBOMETRIC EQOATION

We begin with our differential equation (6.34) where we choose

$$
\begin{equation*}
p(x)=\frac{c-[a+b+1] x}{x(1-x)}, \quad q(x)=\frac{-a b}{x(1-x)} . \tag{6.44}
\end{equation*}
$$

Here $a, b$ and $c$ are constants. For convenience we define $a=a+b+1$. The result is a second order linear differential equation with three regular
singular points which is called the hypergeometric equation. Solutions can be found in the form of power series, convergent for $|x|<1$. The two linearly independent solutions can be witten as

$$
\begin{equation*}
h_{1}=F(a, b ; c ; x), \tag{6.45}
\end{equation*}
$$

and

$$
\begin{equation*}
h_{2}=x^{1-c} F(a+1-c, b+1-c ; 2-c ; x), \tag{6.46}
\end{equation*}
$$

Where the notation for the hypergeometric function is (Ref. 1.1)

$$
\begin{equation*}
F(a, b, c, x)=1+\sum_{n=1}^{\infty} \frac{(a)_{n}(b)_{n}}{(c)_{n} n!} x^{n} . \tag{6.47}
\end{equation*}
$$

with

$$
\begin{align*}
& (a)_{n}=a(\alpha+1)(a+2) \cdots(a+n-1) \quad n \geq 1  \tag{6.48}\\
& (a)_{0}=1 .
\end{align*}
$$

Note that the series are finite for negative integers or zero.
From (6.40) we have that

$$
\begin{equation*}
P(z)=\frac{c-a u(z)}{a(z)[1-u(z)]}, \quad Q(z)=\frac{-a b}{u(z)[1-a(z)]} . \tag{6.49}
\end{equation*}
$$

so that (6.43) can be written as

$$
\begin{aligned}
\frac{d}{d z} \ln \left[\frac{u^{\prime}}{w^{2}}\right] & =n^{\prime} \frac{c-\alpha u}{u(1-u)} \\
& =c \frac{u^{\prime}}{n}-(\alpha-c) \frac{n^{\prime}}{1-u} \\
& =\frac{d}{d z} \ln \left[u^{c}(1-n)^{\alpha-c}\right]
\end{aligned}
$$

which can be integrated to yield a relation between $w$ and $n$

$$
\begin{equation*}
w^{2}(z)=D_{0} \frac{u^{\prime}}{u^{c}(1-\mathfrak{u})^{a-c}} \tag{6.50}
\end{equation*}
$$

where $D_{0}$ is an integration constant. Alternatively we could write from (6.38) with $A(z)=0$

$$
\begin{equation*}
\frac{w^{\prime}}{w}=\frac{1}{2}\left[\frac{u^{\prime \prime}}{u^{\prime}}-\frac{n^{\prime}(c-a u(z))}{u(1-u)}\right] \text {, } \tag{6.51}
\end{equation*}
$$

so that our representation for $B(z)$ from (6.41) is

$$
\begin{equation*}
B(z)=\frac{d}{d z}\left[\frac{w^{\prime}}{w}\right]-\left[\frac{w^{\prime}}{w}\right]^{2}-\frac{a b\left[u^{\prime}\right]^{2}}{u(1-u)} \tag{6.52}
\end{equation*}
$$

which, using (6.51), expresses B entirely in terms of $n(z)$, which we now choose.

To motivate a choice of $\mathfrak{n}(\mathrm{z})$, note that $\mathrm{from}(6.52)$ we would like a constant background term in the inder of refraction. The differential equation

$$
\begin{equation*}
\frac{\left(u^{\prime}\right)^{2}}{u(1-u)}=4 f^{2} \tag{6.53}
\end{equation*}
$$

where $f$ is a constant, has the solution

$$
\begin{equation*}
\mathfrak{u}(z)=\sin ^{2}(f z+g), \tag{6.54}
\end{equation*}
$$

where $g$ is an integration constant. From (6.50) we get

$$
\begin{equation*}
w^{2}(z)=2 D_{0} f[\sin (f z+g)]^{1-2 c}[\cos (f z+g)]^{1+2 c-2 a}, \tag{6.55}
\end{equation*}
$$

so that

$$
\frac{W^{\prime}}{W}=\frac{f}{2}[(1-2 c) \operatorname{ctn}(f z+g)-(1+2 c-2 a) \tan (f z+g)]
$$

and hence

$$
\frac{d}{d z}\left[\frac{w^{\prime}}{w}\right]=-\frac{f^{2}}{2}\left[(1-2 c)[\csc (f z+g)]^{2}+(1+2 c-2 \alpha)[\sec (f z+g)]^{2}\right]
$$

and

$$
\begin{gathered}
{\left[\frac{w^{\prime}}{w}\right]^{2}=\frac{f^{2}}{4}\left[(1-2 c)^{2}[\operatorname{ctn}(f z+g)]^{2}+(1-2 c-2 \alpha)^{2}[\tan (f z+g)]^{2}\right.} \\
-2(1-2 c)(1+2 c-2 \alpha)] .
\end{gathered}
$$

so that from (6.52) we get

$$
\begin{align*}
B(z) & =f^{2}\left[\frac{1}{2}(1-2 c)(1+2 c-2 \alpha)-4 a b\right]  \tag{6.56}\\
& -\left(f^{2} / 2\right)\left[(1-2 c)[\csc (f z+g)]^{2}+(1+2 c-2 \alpha)[\sec (f z+g)]^{2}\right] \\
& -\left(f^{2} / 4\right)\left[(1-2 c)^{2}[\operatorname{ctn}(f z+g)]^{2}+(1+2 c-2 \alpha)^{2}[\tan (f z+g)]^{2}\right]
\end{align*}
$$

An alternate version of this expression is

$$
\begin{equation*}
B(z)=f^{2}\left[a_{1}+\frac{a_{2}}{[\sin (f z+g)]^{2}}+\frac{a_{3}}{[\cos (f z+g)]^{2}}\right] \tag{6.57}
\end{equation*}
$$

where

$$
\begin{align*}
& a_{1}=(a-b)^{2},  \tag{6.58}\\
& a_{2}=-(c-1 / 2)(c-3 / 2), \tag{6.59}
\end{align*}
$$

and

$$
\begin{equation*}
a_{3}=1 / 4-(c-a-b)^{2} \tag{6.60}
\end{equation*}
$$

## Some simple properties of the profile (6.57) are:

(a) All the coefficients, hence the profile, are positive if we require $a+b, 1 / 2<c<3 / 2$, and $-1 / 2<c-a-b<1 / 2$.
(b) The derivative of $B$ vanishes at $z=z_{0}$ if

$$
\begin{equation*}
\left[\tan \left(f z_{0}+g\right)\right]^{4}=a_{2} / a_{3} \tag{6.61}
\end{equation*}
$$

An a priori choice of $z_{0}$ can be used to fix the other parameters. The values of the function and its second derivative at $z=z_{0}$ are

$$
\begin{equation*}
B\left(z_{0}\right)=f^{2}\left[a_{1}+a_{2}+a_{3}+2 \sqrt{a_{2} a_{3}}\right] \tag{6.62}
\end{equation*}
$$

and

$$
\begin{equation*}
B^{\prime \prime}\left(z_{0}\right)=8 f^{4}\left[a_{2}+a_{3}+2 \sqrt{a_{2} a_{3}}\right]=8 f^{2}\left[B\left(z_{0}\right)-a_{1} f^{2}\right] . \tag{6.63}
\end{equation*}
$$

A profile minimum occurs at $z_{0}$ if $B^{\prime \prime}\left(z_{0}\right)>0$ or if $B\left(z_{0}\right) \geqslant a_{1} f^{2}$ which implies $a_{2}+a_{3}+2 \sqrt{a_{2} a_{3}}>0$. A profile maximum occurs at $z_{0}$ if $B^{n}\left(z_{0}\right)<0$ or if $a_{2}+a_{3}+2 \sqrt{a_{2} a_{3}}<0$, in which case both $a_{2}$ and $a{ }_{3}$ must be negative in order for the rhs of (6.61) to be positive.
(c) From (6.45) and (6.46) the solutions corresponding to the profile (6.57) are

$$
\begin{equation*}
h_{1}=F\left(a, b ; c ; \sin ^{2}(f z+g)\right), \tag{6.64}
\end{equation*}
$$

and

$$
\begin{equation*}
h_{2}=[\sin (f z+g)]^{2(1-c)} F\left(a+1-c, b+1-c ; 2-c ; \sin ^{2}(f z+g)\right) \tag{6.65}
\end{equation*}
$$

(d) Note that from (6.57) if we choose $f=k_{0}$ to cancel the $k_{0}$ terms in $B(z)=k_{0}^{2} n^{2}(z)$, the resulting profile $n(z)$ is frequency dependent due to the remaining $f$ terms in the denominator sine and cosine functions.

However, we can scale this out. See the remarks in Appendix 6A.

## Bg._1a.

As a simple example and check on the method, we should recover the solutions for a constant profile. Let $c=1 / 2$ and $a=-b=1 / 2$ in (6.57) then

$$
\begin{equation*}
B(z)=k_{0}^{2} n^{2}(z)=f^{2} . \tag{6.66}
\end{equation*}
$$

The solutions are from (6.45)

$$
\begin{equation*}
h_{1}=F\left(1 / 2,-1 / 2 ; 1 / 2 ; \sin ^{2}(f z+g)\right)=\cos (f z+g) \text {. } \tag{6.67}
\end{equation*}
$$

and

$$
\begin{equation*}
h_{z}=\sin (f z+g) F\left(1,0,3 / 2, \sin ^{-2}(f z+g)\right)=\sin (f z+g), \tag{6.68}
\end{equation*}
$$

both evaluations of which can be found in Ref. 1.2, pg. 1040.

Bg. 1b.
Suppose in the representation (6.56) we choose $c=1 / 2$, then we get
$B(z)=f^{2}\left[-4 a b-(1-a)[\sec (f z+g)]^{2}-(1-\alpha)^{2}[\tan (f z+g)]^{2}\right]$.
For convenience let $g=0$ and choose $f$ to be pare imaginary, $f=$ if $_{1}$, $f_{1}$ real. Then using $\cos \left(i f_{1} z\right)=\cosh \left(f_{1} z\right)$ and $\tan \left(i f_{1} z\right)=i \tanh \left(f_{1} z\right)$ we get

$$
\begin{equation*}
B(z)=f_{1}^{2}\left[4 a b+(1-\alpha)\left[\operatorname{sech}\left(f_{1} z\right)\right]^{2}-(1-\alpha)^{2}\left[\tanh \left(f_{1} z\right)\right]^{2}\right] \tag{6.69}
\end{equation*}
$$

which bears a resemblance to the Epstein profile illustrated later. Here however, the eigenfunctions are

$$
h_{1}=F\left(a, b ; c ;-\sinh ^{2}\left(f_{1} z\right)\right), c=1 / 2,
$$

and

$$
h_{2}=\left[i \sinh \left(f_{1} z\right)\right] F\left(a+1-c, b+1-c ; 2-c ;-\sinh ^{2}\left(f_{1} z\right)\right)
$$

For example, for $a=b=c=1 / 2$ we get (Ref. 1.2, pg. 1042)

$$
\begin{equation*}
h_{1}=\operatorname{sech}\left(f_{1} z\right) \tag{6.70}
\end{equation*}
$$

and

$$
\begin{equation*}
h_{2}=\tanh \left(f_{1} z\right) \tag{6.71}
\end{equation*}
$$

each of which is a solution of $h^{\prime \prime}-f_{1}^{2} h=0$. This makes sense because $a=2$ and the only profile remaining is $B(z)=f_{1}^{2}$. The minus sign in the equation results from rotating $z$ to $i z$, or equivalently $f \rightarrow$ if $_{1}$.

## Bg. 1c.

For this example we directly pick the transformation function $w(z)$ as

$$
\begin{equation*}
w(z)=e^{\beta z} \tag{6.72}
\end{equation*}
$$

Then from (6.50) we have that

$$
u^{\prime}(z)=\frac{1}{D_{0}} e^{2 \beta z} u^{c}(1-u)^{a-c}
$$

Let $\alpha=c$ and choose

$$
u(z)=e^{\alpha z} .
$$

We have a solution provided

$$
\gamma=\frac{1}{D_{0}}=2 \beta+c \gamma \quad \text { or } \quad \beta=(1-c) \gamma / 2
$$

Then from (6.52) we get

$$
\begin{equation*}
B(z)=-\left[\frac{(1-c) \gamma}{2}\right]^{2}-a b \frac{\gamma^{2} e^{\gamma z}}{1-e^{\gamma z}} . \tag{6.73}
\end{equation*}
$$

If we want to shift the origin from $z=0$ define

$$
\begin{equation*}
u(z)=e^{\gamma\left(z-z_{0}\right)}, \tag{6.74}
\end{equation*}
$$

then the constraints are

$$
\gamma=2 \beta+c \gamma \quad \text { and } \quad \gamma D_{0}=e^{\gamma(1-c) z_{0}},
$$

with

$$
\begin{equation*}
B(z)=-\beta^{2}-\frac{a b \gamma^{2} e^{\gamma\left(z-z_{0}\right)}}{1-e^{\gamma\left(z-z_{0}\right)}} . \tag{6.75}
\end{equation*}
$$

Eg. 1d. Bpstein Profile
Here we choose our transformation function as

$$
\begin{equation*}
\mathfrak{n}(z)=1 / 2(1+\tanh (z / 2), \tag{6.76}
\end{equation*}
$$

so that

$$
1-n=1 / 2(1-\tanh (z / 2)),
$$

and

$$
u(1-u)=1 / 4[\operatorname{sech}(z / 2)]^{2},
$$

with the results

$$
\begin{aligned}
u^{\prime}(z) & =1 / 4 \operatorname{sech}^{2}(z / 2) \\
u^{\prime \prime}(z) & =-1 / 4 \operatorname{sech}^{2}(z / 2) \tanh (z / 2), \\
\frac{\mathbf{u}^{\prime \prime}(z)}{u^{\prime}(z)} & =-\tanh (z / 2) \\
\frac{u^{\prime}(z)}{u(z)} & =\frac{1}{2} \frac{\operatorname{sech}^{2}(z / 2)}{1+\tanh ^{\prime}(z / 2)}=\frac{1}{2} \frac{1-\tanh ^{2}(z / 2)}{1+\tanh (z / 2)} \\
& =\frac{1}{2}[1-\tanh (z / 2)] .
\end{aligned}
$$

and

$$
\frac{n^{\prime}}{1-u}=\frac{1 / 4 \operatorname{sech}^{2}(z / 2)}{1 / 2[1-\tanh (z / 2)]}=\frac{1}{2}[1+\tanh (z / 2)] .
$$

Me thas have from (6.51) that

$$
\begin{aligned}
\frac{w^{\prime}}{w} & =\frac{1}{2}\left[\frac{\mathbf{a}^{\prime \prime}}{\mathbf{a}^{\prime}}-c \frac{\mathbf{n}^{\prime}}{\mathbf{u}}-(c-a) \frac{\mathbf{n}^{\prime}}{1-\mathrm{u}}\right] \\
& =-\frac{1}{2}\left[c-\frac{a}{2}+\left(1-\frac{\alpha}{2}\right) \tanh (z / 2)\right] .
\end{aligned}
$$

and

$$
\frac{d}{d z}\left[\frac{w^{\prime}}{w}\right]=-\frac{1}{4}\left(1-\frac{\alpha}{2}\right) \operatorname{sech}^{2}(z / 2)
$$

$$
\begin{aligned}
{\left[\frac{w^{\prime}}{w}\right]^{2}=\frac{1}{4}\left[(c-a / 2)^{2}\right.} & +2(c-\alpha / 2)(1-\alpha / 2) \tanh (z / 2) \\
& \left.+(1-\alpha / 2)^{2}[\tanh (z / 2)]^{2}\right] .
\end{aligned}
$$

and if we replace

$$
[\tanh (z / 2)]^{2}=1-[\operatorname{sech}(z / 2)]^{2}
$$

we have

$$
\begin{aligned}
{\left[\frac{w^{\prime}}{w}\right]^{2}=\frac{1}{4} } & {\left[(c-\alpha / 2)^{2}+(1-\alpha / 2)^{2}\right.} \\
& +2(c-\alpha / 2)(1-\alpha / 2) \tanh (z / 2) \\
& \left.-(1-\alpha / 2)^{2}[\operatorname{sech}(z / 2)]^{2}\right]
\end{aligned}
$$

We can thus write from (6.52) that

$$
\begin{equation*}
B(z)=B_{0}+B_{1}[\operatorname{sech}(z / 2)]^{2}+B_{2} \tanh (z / 2), \tag{6.77}
\end{equation*}
$$

where

$$
\begin{align*}
& B_{0}=\frac{1}{4}\left[(c-\alpha / 2)^{2}+(1-\alpha / 2)^{2}\right],  \tag{6.78}\\
& B_{1}=-\frac{1}{4}\left(1-\frac{a}{2}\right)+\frac{1}{4}\left(1-\frac{a}{2}\right)^{2}-\frac{a b}{4}, \tag{6.79}
\end{align*}
$$

and

$$
\begin{equation*}
B_{2}=-\frac{1}{2}(c-a / 2)(1-a / 2) . \tag{6.80}
\end{equation*}
$$

Equation (6.77) has the form of the Epstein profile (Ref. 1.3). Some properties of the profile are:
(a)

$$
\begin{align*}
B(0) & =\frac{1}{4}(1-\alpha / 2)^{2}-\frac{1}{4}(1-\alpha / 2)-\frac{a b}{4}-\frac{1}{4}(c-\alpha / 2)^{2}-\frac{1}{4}(1-\alpha / 2)^{2} \\
& =-\frac{1}{4}\left[1-\frac{a}{2}+a b+[c-\alpha / 2]^{2}\right] \\
& =B_{0}+B_{1}, \tag{6.81}
\end{align*}
$$

(b)

$$
B^{\prime}(z)=-B_{1} \operatorname{sech}^{2}(z / 2) \tanh (z / 2)+1 / 2 B_{2} \operatorname{sech}^{2}(z / 2),
$$

$$
\begin{equation*}
B^{\prime}(z)=[\operatorname{sech}(z / 2)]^{2}\left[1 / 2 B_{2}-B_{1} \tanh (z / 2)\right], \tag{6.82}
\end{equation*}
$$

so that $B^{\prime}\left(z_{0}\right)=0$ if

$$
\begin{equation*}
T \equiv \tanh \left(z_{0} / 2\right)=B_{2} / 2 B_{1} \tag{6.83}
\end{equation*}
$$

Since the tanh is positive for $z_{0}>0, B_{2}$ and $B_{1}$ must have the same sign.
(c)

$$
\begin{align*}
& \mathrm{B}^{\prime \prime}(\mathrm{z})=[\operatorname{sech}(z / 2)]^{2}\left[2 \mathrm{~B}_{1}[\tanh (z / 2)]^{2}-\frac{B_{2}}{2} \tanh (z / 2)-B_{1}\right],(6.84) \\
& \text { so that using }(6.83) \text { we get } \\
& \mathrm{B}^{\prime \prime}\left(\mathrm{z}_{0}\right)=\mathrm{B}_{1}\left[\operatorname{sech}\left(\mathrm{z}_{0} / 2\right)\right]^{2}\left[\mathrm{~T}^{2}-1\right] \\
&  \tag{6.85}\\
& =-B_{1}\left[\operatorname{sech}\left(z_{0} / 2\right)\right]^{4}
\end{align*}
$$

so that $z_{0}$ is a minimum if $B_{1}\left\langle 0\right.$ and a maximum if $B_{1}>0$.

## Eg. 2. CONFLOENT BYPBRGBOMETRIC BOUATION

As our second example we again start with (6.34) where we now choose

$$
\begin{equation*}
p(x)=(c-x) / x, \quad q(x)=a / x, \tag{6.86}
\end{equation*}
$$

The resulting equation is the confluent hypergeometric equation with and c constants.

Bg. 2a.
Choose both transformation functions as

$$
\begin{equation*}
u(z)=\beta z \quad, \quad w(z)=z^{\gamma} e^{\delta z} \tag{6.87}
\end{equation*}
$$

where from (6.43) we have the constraint

$$
\begin{align*}
2 \frac{w^{\prime}}{w} & =\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}-\mathbf{u}^{\prime} P(z) \\
& =\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}-\frac{\mathbf{u}^{\prime}}{\mathbf{u}}(c-u)=\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}-c \frac{\mathbf{u}^{\prime}}{\mathbf{u}}+\mathbf{u}^{\prime} \\
& =-\frac{c}{z}+\beta \quad \tag{6.88}
\end{align*}
$$

with

$$
W^{\prime}=\gamma z^{\gamma-1} e^{\delta z}+\delta z^{\gamma} e^{\delta z}=\left(\frac{\gamma}{z}+\delta\right) W,
$$

so that

$$
2\left(\frac{\gamma}{z}+\delta\right)=\beta-\frac{c}{z} .
$$

and

$$
\begin{equation*}
\delta=\beta / 2 \quad \gamma=-c / 2 . \tag{6.89}
\end{equation*}
$$

We have that

$$
\left[\frac{W^{\prime}}{W}\right]^{2}=\frac{1}{4}\left(\beta^{2}-2 \beta \frac{c}{z}+\frac{c}{z}\right)
$$

and

$$
\frac{d}{d z}\left[\frac{w^{\prime}}{w}\right]=\frac{c}{2 z^{2}},
$$

so that from (6.41)

$$
\begin{align*}
B(z) & =\frac{c}{2 z^{2}}-\frac{1}{4}\left[\beta^{2}-\frac{2 \beta c}{z}+\frac{c}{z^{2}}\right]-\frac{\beta^{2}}{z} \\
& =-\frac{1}{4} \beta^{2}+\frac{1}{2} \beta \frac{c}{z}-\frac{\beta^{2}}{z}+\frac{c}{4 z^{2}} \\
& =B_{0}+\frac{B_{1}}{z}+\frac{B_{2}}{z^{2}}, \tag{6.90}
\end{align*}
$$

where

$$
\begin{align*}
& B_{0}=-\frac{1}{4} \beta^{2}  \tag{6.91}\\
& B_{1}=\beta(c / 2-\beta) \tag{6.92}
\end{align*}
$$

and

$$
\begin{equation*}
B_{2}=c / 4, \tag{6.93}
\end{equation*}
$$

and thus for an algebraic $u(z)$ we get an algebraic profile. If we want a profile in a region including $z=0$ we must shift this origin away.

Es. 2b.
As a second example we choose an exponential transformation function

$$
\begin{equation*}
u(z)=e^{-\beta z}, \tag{6.94}
\end{equation*}
$$

and from the constraint (6.43)

$$
2 \frac{w^{\prime}}{w}=\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}-\mathbf{u}^{\prime} \mathbf{P}(z)
$$

we have that

$$
\begin{equation*}
\frac{w^{\prime}}{w}=\frac{\beta}{2}\left(c-1-e^{-\beta z}\right) \tag{6.95}
\end{equation*}
$$

so that $B(z)$ from (6.41) can be written as

$$
\begin{equation*}
B(z)=a_{0}+a_{1} e^{-1 \beta z}+a_{2} e^{-2 b z} \tag{6.96}
\end{equation*}
$$

where

$$
\begin{align*}
& a_{0}=-\frac{1}{4} \beta^{2}(c-1)^{2} \\
& a_{1}=\beta^{2}\left(\frac{c}{2}-a\right) \tag{6.97}
\end{align*}
$$

and

$$
\begin{equation*}
a_{2}=-\beta^{2} / 4 \tag{6.99}
\end{equation*}
$$

Thus for an exponential choice of transformation function we get an exponential profile.

Eg-3. BESSRL BQUATION
As a third example, start with (6.34) with

$$
\begin{equation*}
p(x)=1 / x \quad, \quad q(x)=1-\alpha^{2} / x^{2} . \tag{6.100}
\end{equation*}
$$

which is the Bessel differential equation with solutions $Z_{a}(x)$ where $Z_{\alpha}$ represents the appropriate Bessel and Hankel function for a given boundary value problem.

## Bg. 3a.

Choose the transformation function as

$$
\begin{equation*}
\mathfrak{u}(z)=\beta z \quad \text {, } \tag{6.101}
\end{equation*}
$$

so that the constraint (6.43) is

$$
\begin{equation*}
2 \frac{\varpi^{\prime}}{w}=\frac{\mathbf{u}^{\prime \prime}}{\mathbf{u}^{\prime}}-{n^{\prime}}^{\prime} P(z)=\frac{u^{\prime \prime}}{u^{\prime}}-\frac{\mathbf{n}^{\prime}}{\mathbf{u}}=-\beta, \tag{6.102}
\end{equation*}
$$

and

$$
\begin{equation*}
w(z)=\bar{e}^{-\beta z / 2} \tag{6.103}
\end{equation*}
$$

The profile $B(z)$ from (6.41) is

$$
\begin{align*}
B(z) & =\frac{d}{d z} \frac{W^{\prime}}{W}-\left[\frac{W^{\prime}}{W}\right]^{2}+\left(u^{\prime}\right)^{2} Q(z) \\
& =-\frac{\beta^{2}}{4}+\beta^{2}+\beta^{2}\left(1-\alpha^{2} / \beta^{2} z^{2}\right)  \tag{6.104}\\
& =\frac{3}{4} \beta^{2}-\frac{\alpha^{2}}{z^{2}}
\end{align*}
$$

with solutions

$$
\begin{equation*}
\phi=Z_{\alpha}(\beta z) \tag{6.105}
\end{equation*}
$$

Eg. 3b.
Choose an exponential transformation function

$$
\begin{equation*}
x=u(z)=e^{-\beta z} \tag{6.106}
\end{equation*}
$$

and the constraint (6.43) is

$$
2 \frac{w^{\prime}}{w}=\frac{u^{\prime \prime}}{u^{\prime}}-\frac{u^{\prime}}{u}=-\beta+\beta=0
$$

so that

$$
\begin{equation*}
w(z)=w_{0}=\text { constant } \tag{6.107}
\end{equation*}
$$

Thns from (6.41)

$$
\begin{aligned}
B(z) & =\left(u^{\prime}\right)^{2} Q(z)=\beta^{2} e^{-2 \beta z}\left[1-\frac{\alpha^{2}}{e^{-2 \beta z}}\right] \\
& =\beta^{2}\left[e^{-2 \beta z}-\alpha^{2}\right]
\end{aligned}
$$

so that the solutions of

$$
\begin{equation*}
\frac{d^{2} \phi}{d z^{2}}+B(z) \phi=0 \tag{6.109}
\end{equation*}
$$

are

$$
\begin{equation*}
\phi=Z_{a}\left(e^{-\beta z}\right) \tag{6.110}
\end{equation*}
$$

## APPENDIX 1A. SCALING AND FREQUENCY INDBPENDENT $\mathrm{m}(\mathrm{z})$

In general we have that

$$
B(z)=k_{0}^{2} n^{2}(z)=r h s .
$$

where the rhs contains various constants. We want $n(z)$ to be frequency independent (i.e. independent of $k_{0}$ ), as we previously remarked. We can't do this for a direct choice of constants since the $k_{0}$ is buried in the functional dependence of the profile. We scale out this frequency dependence as follows: Our differential equation on $\phi$ is

$$
\frac{d^{2} \phi}{d z^{2}}+k_{0}^{2} n^{2}(z) \phi=0
$$

Scale the $z$-coordinate to give $z^{*}=k_{0} z$ so that

$$
\frac{d^{2} \phi}{d z^{*}}+n^{2}\left(z^{*}\right) \phi=0 \text {. }
$$

so we are really finding $B\left(z^{*}\right)=n^{2}\left(z^{*}\right)$. But this has the same functional dependence as $n^{2}(z)$. So once we find $n^{2}\left(z^{*}\right)$ simply replace $z^{*}$ by $z$ (not by $k_{0} z$ ) to get $n(z)$. The correct scaling occurs in the solution since there we automatically replace $z^{*}$ by $k_{0} z$.

## APPBNDIX 1B. Darboax Theorew

Once we know one profile and its solution it is possible to find other profiles in a systematic way. It goes like this:

If the general solution of the second order linear differential equation

$$
\frac{d^{2} f}{d z^{2}}=[h+a(z)] f
$$

is known for all values of $h$, and $v(z)$ is a particular solution of this equation for $h=h_{1}$, i.e.

$$
\frac{d^{2} v}{d z^{2}}=\left[h_{1}+a(z)\right] v,
$$

then the general solution of

$$
\frac{d^{2} g}{d z^{2}}=\left[h-h_{1}+v(z) \frac{d^{2}}{d z^{2}}\left[\frac{1}{v(z)}\right]\right] g
$$

for $h \neq h_{1}$ is given by

$$
g(z)=v(z) \frac{d}{d z}\left[\frac{f(z)}{v(z)}\right]
$$

## 2. SOLOTION OR INITIAL AND BOUNDARY VALOR PROBLEMS

In this chapter we study the initial and boundary value problems for the wave equation, Helmholtz equation, and the parabolic equation. We discuss the integral representations of the first two in detail. In addition we briefly mention the Rayleigh-Sommerfield integral representations, the extended boundary condition or extinction coefficient method, the $T$-matrix approach, and the Kirchhoff approximation.

### 2.1 VAVE BQOATION

We write the wave equation for the Green's function as

$$
\begin{equation*}
\left[V^{2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}\right] G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)=-\delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right) \delta\left(t-t^{\prime}\right), \tag{1.1}
\end{equation*}
$$

which is related to our four-dimensional formulation of the Green's function by a factor $c^{-1}$. We assume the Green's function satisfies causality given by

$$
\begin{equation*}
\left.G\left(\underline{\sim}, t ; \underline{x}^{\prime}, t^{\prime}\right)=0 \quad t^{\prime}\right\rangle t \tag{1.2}
\end{equation*}
$$

That is, no signal is present for measurement times in the field, $t$, less than the initial time of the source, $t^{\prime}$. We first prove reciprocity given by

$$
\begin{equation*}
G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)=G\left(\underset{\sim}{x},-t^{\prime} ; \underset{\sim}{x},-t\right) \tag{1,3}
\end{equation*}
$$

which yields a relation between exchanging source and receiver positions and times. The proof goes as follows. Since we have a second derivative in time we can write an equation analogous to (1.1) in the form

$$
\begin{equation*}
\left[\nabla^{2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}\right] G\left(\underset{\sim}{x},-t ;{\underset{\sim}{x}}^{n},-t^{n}\right)=-\delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{n}\right) \delta\left(t-t^{n}\right) \tag{1.4}
\end{equation*}
$$

Next, multiply (1.1) by $G\left(\underset{\sim}{x},-t ; x_{\sim}^{\prime \prime},-t^{\prime \prime}\right)$ and (1.4) by $G\left(\underset{\sim}{x}, t ; x^{\prime}, t^{\prime}\right)$ and subtract the resulting equations. Integrate the result over all space and time to yield

$$
\begin{align*}
& I_{1}\left(\underset{\sim}{x^{\prime}},{\underset{\sim}{x}}^{n} ; t^{\prime}, t^{n}\right)-\frac{1}{c^{2}} I_{2}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}^{\prime \prime} ; t^{\prime}, t^{n}\right) \\
& \quad=G\left({\underset{\sim}{x}}^{\prime \prime}, t^{\prime \prime} ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)-G\left(x_{\sim}^{\prime},-t^{\prime} ;{\underset{\sim}{x}}^{\prime \prime},-t^{n}\right) \tag{1.5}
\end{align*}
$$

where $I_{1}$ and $I_{2}$ are defined as

$$
\begin{array}{r}
I_{1}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime \prime} ; t^{\prime}, t^{\prime \prime}\right)=\int d t \iiint d x\left[G\left(\underset{\sim}{x},-t ; \underset{\sim}{x} x^{\prime \prime},-t^{\prime \prime}\right) \nabla^{2} G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)\right.  \tag{1.6}\\
\left.-G\left(\underset{\sim}{x}, t_{;}{\underset{\sim}{x}}^{\prime}, t^{\prime}\right) \nabla^{2} G\left(\underset{\sim}{x},-t ; x^{\prime \prime},-t^{\prime \prime}\right)\right],
\end{array}
$$

and

$$
\begin{align*}
I_{2}\left(x^{\prime}, x^{\prime \prime} ; t^{\prime}, t^{\prime \prime}\right)=\int & \iint d x \int d t\left[G\left(\underset{\sim}{x},-t ; x_{\sim}^{\prime \prime},-t^{\prime \prime}\right) \frac{\partial^{2}}{\partial t^{2}} G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)\right. \\
& \left.-G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right) \frac{\partial^{2}}{\partial t^{2}} G\left(\underset{\sim}{x},-t ; x_{\sim}^{\prime \prime},-t^{\prime \prime}\right)\right] \quad( \tag{1.7}
\end{align*}
$$

We prove both integrals vanish, and reciprocity follows from (1.5). In $I_{1}$ use Green's theorem on the volme integral part to yield a surface integral of the form

$$
\begin{align*}
\iint d S & {\left[G\left(x_{s},-t ;{\underset{\sim}{x}}^{\prime \prime},-t^{\prime \prime}\right) \frac{\partial}{\partial n} G\left({\underset{\sim}{x}}_{s^{\prime}}, t ; x_{\sim}^{\prime}, t^{\prime}\right)\right.} \\
& \left.-G\left(\underset{\sim}{x}, t,{\underset{\sim}{x}}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial n} G\left(\underset{\sim}{x},-t,{\underset{\sim}{x}}^{\prime \prime},-t^{\prime \prime}\right)\right], \tag{1.8}
\end{align*}
$$

where $n$ is the outward normal over any bounded closed surfaces in the problem. The contribution from the surface at infinity vanishes becanse the functions satisfy the radiation condition. The surface is specified by
 derivative, or a homogeneous linear combination of the two vanish on the surface, then (1.8) is identically zero. Next, write the temporal integral in $I_{2}$, using the fact that the integrand is an exact differential, as

$$
\begin{align*}
& \int_{-\infty}^{\infty} d t \frac{\partial}{\partial t}\left[G\left(\underset{\sim}{x},-t ;{\underset{\sim}{x}}^{\prime \prime},-t^{\prime \prime}\right) \frac{\partial}{\partial t} G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)\right.  \tag{1.9}\\
&\left.-G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial t} G\left(\underset{\sim}{x},-t ;{\underset{\sim}{x}}^{\prime \prime},-t^{\prime \prime}\right)\right]
\end{align*}
$$

When integrated and evaluated at $\mathrm{t}_{\mathrm{m}}$ the terms vanish by cansality. Thus both $I_{1}$ and $I_{2}$ vanish and reciprocity follows from (1.5).

In order to derive the integral relation for the wave equation we begin with equations on the field function $\varphi$ (arising from a source $S$ ) and the Green's function as follows

$$
\begin{align*}
& {\left[\forall^{\prime 2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime}}\right] \phi\left(\underset{\sim}{x^{\prime}}, t^{\prime}\right)=-S\left(\underset{\sim}{x^{\prime}}, t^{\prime}\right),}  \tag{1.10}\\
& {\left[\forall^{\prime 2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime}}\right] G\left(x, t ; x^{\prime}, t^{\prime}\right)=-\delta\left(\underset{\sim}{x}-x^{\prime}\right) \delta\left(t-t^{\prime}\right),} \tag{1.11}
\end{align*}
$$

where (1.11) follows from (1.1) by interchanging variables, viz.

$$
\begin{equation*}
\left[\nabla^{\prime 2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime}}\right] G\left(x_{\sim}^{\prime}, t^{\prime}, x, t\right)=-\delta\left(\underset{\sim}{x}-x_{\sim}^{\prime}\right) \delta\left(t-t^{\prime}\right) \tag{1.12}
\end{equation*}
$$

If we now let $t \rightarrow-t$ and $t^{\prime \prime} \rightarrow-t^{\prime}$ we get

$$
\begin{equation*}
\left[\forall^{\prime 2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime}}\right] G\left(x^{\prime},-t^{\prime} ; x,-t\right)=-\delta\left(\underset{\sim}{x}-x^{\prime}\right) \delta\left(t-t^{\prime}\right), \tag{1.13}
\end{equation*}
$$

and, by reciprocity, (1.11) follows from this. Next, maltiply (1.10) by $G\left(x, t ; x^{\prime}, t^{\prime}\right)$ and (1.11) by $\phi\left(x^{\prime}, t^{\prime}\right)$ and subtract the resulting equations.

Integrate the results over all space and from $t=0$ to $t=\infty$ in time. The result is witten as ( $t=0$ is the initial time).

$$
\begin{align*}
& \phi(x, t)=\int_{0}^{\infty} d t^{\prime} \iiint d x_{\sim}^{\prime} G\left(\underset{\sim}{x}, t_{;} x_{\sim}^{\prime}, t^{\prime}\right) S\left(x^{\prime}, t^{\prime}\right) \\
& +\int_{0}^{\infty} d t^{\prime} \iiint d x^{\prime}\left[G\left(\underset{\sim}{x}, t_{3} x_{\sim}^{\prime}, t^{\prime}\right) \nabla^{\prime}{ }_{\phi}\left(\underset{\sim}{x^{\prime}}, t^{\prime}\right)\right. \\
& \left.-\phi\left(x_{\sim}^{\prime}, t^{\prime}\right) V^{\prime 2} G\left(\underset{\sim}{x}, t ; x^{\prime}, t^{\prime}\right)\right] \\
& -\frac{1}{c^{2}} \iiint d x^{\prime} \int_{0}^{\infty} d t^{\prime} \frac{\partial}{\partial t^{\prime}},\left[G\left(x, t, x_{\sim}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial t}, \phi\left(x^{\prime}, t^{\prime}\right)\right. \\
& \left.-\phi\left({\underset{\sim}{x}}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial t}, G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)\right] \quad . \tag{1.14}
\end{align*}
$$

In the second integral term in (1.14) we use the spatial Green's theorem as in (1.8), and we integrate the temporal part of the third integral whose integrand is an exact differential. The infinite surface contributions vanish by the radiation condition, and the infinite time contributions vanish using cansality. The resulting integrals can be restricted using causality to give the final result

$$
\begin{align*}
& \phi(x, t)=\int_{0}^{t+} d t \cdot \iiint d x^{\prime} G\left(x, t ; x^{\prime}, t^{\prime}\right) S\left(x^{\prime}, t^{\prime}\right) \\
& +\int_{0}^{t+} d t^{\prime} \iint d S^{\prime}\left[G\left(x, t ; x_{S}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial n^{\prime}} p\left(x_{S}^{\prime}, t^{\prime}\right)\right. \\
& \left.-\phi\left(x_{s^{\prime}}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial n^{\prime}} G\left(x, t ; x_{s^{\prime}}^{\prime}, t^{\prime}\right)\right] \\
& +\frac{1}{c^{2}} \iiint d^{\prime}\left[G\left(x, t ; x^{\prime}, 0\right) \frac{\partial}{\partial t^{\prime}} \phi\left(x^{\prime}, 0\right)\right. \\
& \left.-\phi\left(x_{\sim}^{\prime}, 0\right) \frac{\partial}{\partial t^{\prime}} G\left(\underset{\sim}{x}, t ; x^{\prime}, 0\right)\right] . \tag{1.15}
\end{align*}
$$

in terms of a surface integral over finite surfaces having an outward normal $A^{\prime}$. The result for the field function is that it is expressed as a
superposition of wavelets from the source $S$, the boundary surface (or surfaces), and the initial conditions $\varphi\left(x^{\prime}, 0\right)$ and $\partial \phi\left(x^{\prime}, 0\right) / \partial t^{\prime}$. Both initial conditions and one boundary condition specify the problem uniquely.

We now want to examine the separate terms in (1.15). We choose for our Green's function the retarded Green's function in Ch. 1. It is written as

$$
\begin{align*}
G\left(\underset{\sim}{x}, t, x^{\prime}, t^{\prime}\right) & =c G_{R}^{(3,1)}\left(x, x^{\prime}\right) \\
& =\frac{c \delta(\tau-r)}{4 \pi r}, \tau=c\left(t-t^{\prime}\right), r=\left|x-x^{\prime}\right| \\
& =\frac{1}{4 \pi r} \delta\left(t-t^{\prime}-\frac{1}{c}\left|x-x^{\prime}\right|\right) \tag{1.16}
\end{align*}
$$

The first term in (1.15) is

$$
\begin{equation*}
\varphi_{1}(x, t)=\int_{\sim}^{t^{+}} d t^{\prime} \iiint d x_{\sim}^{\prime} G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, t^{\prime}\right) S\left(\underset{\sim}{x^{\prime}}, t^{\prime}\right) \tag{1.17}
\end{equation*}
$$

which is an integral over the domain of the source function $S$. Recall that G was a retarded Green's function and yielded the representation

$$
\begin{equation*}
\phi_{1}(x, t)=\frac{1}{4 \pi r} \iiint d x^{\prime} S\left(x^{\prime}, t-x / c\right), \tag{1.18}
\end{equation*}
$$

where we have used (1.16) in (1.17) and integrated over time. The result is a field $\varphi_{1}$ due to sources integrated over those spatial values such that its temporal values occur before the measurement time $t$, i.e. at retarded time values.

We next discuss several examples.

Bg. 1. MOVING POINT SOURCR
Although many applications relate to fixed sources, we can treat the signal received by a moving point source as follows. Choose

$$
\begin{equation*}
\delta\left({\underset{\sim}{x}}^{\prime}, t^{\prime}\right)=\delta\left(\underset{\sim}{x}{ }^{\prime}-\underset{\sim}{R}\left(t^{\prime}\right)\right) . \tag{1.19}
\end{equation*}
$$

which is a point source moving on a path $\underset{\sim}{\mathbb{R}}\left(\mathrm{t}^{\prime}\right)$. See Fig. 2.1 below.


Fig. 2.1

Osing (1.16) and (1.19) the result of (1.17) is

$$
\begin{equation*}
\phi_{1}(x, t)=\int_{0}^{t} d t^{\prime} \iiint d x_{\sim}^{\prime} \frac{\delta\left(t-t^{\prime}-\frac{1}{c}\left|\underset{\sim}{x}-x^{\prime}\right|\right)}{4 \pi\left|z-x^{\prime}\right|} \delta\left(\underset{\sim}{x} \cdot-\underset{\sim}{R}\left(t^{\prime}\right)\right) . \tag{1.20}
\end{equation*}
$$

The volume integral can be easily evaluated to yield

$$
\begin{equation*}
\varphi_{1}(x, t)=\int_{0}^{t^{+}} d t^{\prime} \frac{\delta\left(t-t^{\prime}-\frac{1}{c}\left|x-B\left(t^{\prime}\right)\right|\right)}{4 \pi\left|Z_{\sim}^{-R}\left(t^{\prime}\right)\right|} . \tag{1.21}
\end{equation*}
$$

To evaluate this final integral use the relation

$$
\begin{equation*}
\int g\left(t^{\prime}\right) \delta\left(f\left(t^{\prime}\right)\right) d t^{\prime}=\int g\left(t^{\prime}\right) \frac{\delta\left(t^{\prime}-t_{0}\right)}{\left|d f / d t^{\prime}\right|} d t^{\prime}=\frac{g\left(t_{0}\right)}{\mid d f / d t^{\prime} T_{t^{\prime}}=t_{0}} \tag{1.22}
\end{equation*}
$$

where

$$
\begin{align*}
& g\left(t^{\prime}\right)=\left[4 \pi\left|\underset{\sim}{x}-\underset{\sim}{R}\left(t^{\prime}\right)\right|\right]^{-1},  \tag{1.23}\\
& f\left(t^{\prime}\right)=t-t^{\prime}-\frac{1}{c}\left|x-R\left(t^{\prime}\right)\right|, \tag{1.24}
\end{align*}
$$

and $t_{0}$ is given by the solution $f\left(t_{0}\right)=0$ so that it satisfies

$$
\begin{equation*}
t_{0}=t-\frac{1}{c}\left|\underset{\sim}{x}-\underset{\sim}{R}\left(t_{0}\right)\right| \tag{1.25}
\end{equation*}
$$

We al so have the result

$$
\begin{equation*}
\frac{d f}{d t^{\prime}}=-1+\frac{1}{c} \frac{d R}{d t^{\prime}} \cdot \frac{\left[x-R\left(t^{\prime}\right)\right]}{\left|x-R\left(t^{\prime}\right)\right|} \cdot \tag{1.26}
\end{equation*}
$$

To carry out the integral we require that $0<t_{0}<t$. The result is the contribution from a moving point source

$$
\begin{equation*}
\phi_{1}(\underset{\sim}{x}, t)=\frac{1}{4 \pi} \frac{\theta\left(t-t_{0}\right) \theta\left(t_{0}\right)}{\left.\left|\underset{\sim}{x}-R_{\sim}\left(t_{0}\right)\right|-\frac{1}{c}\left[\frac{d R}{d t^{\prime}}\right]_{t_{0}} \cdot\left[\underset{\sim}{x}-R_{\sim}\left(t_{0}\right)\right] \right\rvert\,} \tag{1.27}
\end{equation*}
$$

which is called the Lienard-Wiechert potential. It can al so be witten as

$$
\begin{equation*}
\phi_{1}(\underset{\sim}{x}, t)=\frac{1}{4 \pi} \frac{\theta\left(t-t_{0}\right)}{\left|\underset{\sim}{x}-\underset{\sim}{R}\left(t_{0}\right)\right|} \quad \frac{\theta\left(t_{0}\right)}{\left.1-\frac{1}{c} \underset{\sim}{v}\left(t_{0}\right) \cdot n\left(t_{0}\right) \right\rvert\,} \tag{1.28}
\end{equation*}
$$

For the special case $\underset{\sim}{x}=0$, the result of the spatial and temporal superposition of spherical waves is a pulse-like solution. Note that we have evaluated the last delta function in (1.21) assuming only one solution. For a homogeneous medium this is almays true. The reason is that there is one arrival time and thos minimum path for the signal. A signal along any other path would arrive at a later time. However, in an inhomogeneous medium this might not be true. Even though the distance may be longer for a second path, the point source speed $v$ in an inhomogeneous medium might be faster than the wave speed $c$ in the homogeneous region. The result conld be the same arrival time for two waves, one which radiates solely into the homogeneous region to the receiver, the other where the source travels a distance in the inhomogeneous medium, then radiates into the homogeneous medium to the receiver. An example in electromagnetic theory is the Cerenkov effect, where the signal travels faster than the speed of light in the inhomogeneous medium. Note this is phase velocity.

## Eg. 2. MOVING DIPOLR SOURCE

As a second example, consider a moving dipole source term. For simplicity we write the dipole in only one direction, the x-direction, as

$$
\begin{equation*}
S\left({\underset{\sim}{x}}^{\prime}, t^{\prime}\right)=\delta^{\prime}\left(x^{\prime}-R\left(t^{\prime}\right)\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}\right) \tag{1.29}
\end{equation*}
$$

in the $x$-direction. A sum of dipoles in the $x$, $y$, and $z$ directions would correspond to an approximation of an explosive source. The moving dipole in a single direction could correspond to a very simple model of a fault or earthquake. The result for (1.21) is

$$
\phi_{1}(x, t)=\int_{0}^{t+} d t^{\prime} \iiint \int d x^{\prime} \frac{\delta\left(t-t^{\prime}-\frac{1}{c}\left|x^{\prime}-x^{\prime}\right|\right)}{4 \pi\left|x^{-}-x^{\prime}\right|} \delta^{\prime}\left(x^{\prime}-R\left(t^{\prime}\right)\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}\right) \quad(1.30)
$$

If we integrate by parts we get

$$
\phi_{1}(x, t)=-\int_{\delta}^{t} d t \cdot \iiint d x_{\sim}^{\prime} \frac{\partial}{\partial x} \cdot\left[\frac{\delta\left(t-t^{\prime}-\frac{1}{c}\left|\underset{\sim}{x}-x^{\prime}\right|\right)}{4 \pi\left|\underset{\sim}{x}-x^{\prime}\right|}\right] \delta\left(\underset{\sim}{x^{\prime}}-\underset{\sim}{R}\left(t^{\prime}\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}\right) .\right.
$$

The integrand has two terms

$$
\begin{align*}
\frac{\partial}{\partial x^{\prime}}\left[\frac{\left.\delta\left(t-t^{\prime}\right)-\frac{1}{c}\left|x-x^{\prime}\right|\right)}{\left|x_{\sim}-x^{\prime}\right|}\right]= & \frac{\delta^{\prime}\left(t-t^{\prime}-\frac{1}{c}\left|x-x^{\prime}\right|\right)}{\left|x-x^{\prime}\right|} \frac{1}{c} \frac{\left(x-x^{\prime}\right)}{\left|x-x^{\prime}\right|} \\
& +\delta\left(t-t^{\prime}-\frac{1}{c}\left|x^{\prime}-x^{\prime}\right|\right) \frac{\left(x-x^{\prime}\right)}{\left|x-x^{\prime}\right|^{3}} \tag{1.32}
\end{align*}
$$

The first term can be written as

$$
\begin{align*}
\phi_{a}(x, t) & =-\int d t^{\prime} \iiint d x^{\prime} \frac{\delta^{\prime}\left(t-t^{\prime}-\frac{1}{c}\left|\frac{x}{a}-x^{\prime}\right|\right)}{4 \pi\left|x-x^{\prime}\right|^{2}} \frac{\left(x-x^{\prime}\right)}{c} \delta\left(x^{\prime}-R\left(t^{\prime}\right)\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}\right) \\
& =-\frac{1}{4 \pi} \int d t^{\prime} \delta^{\prime}\left(f\left(t^{\prime}\right)\right) g\left(t^{\prime}\right) . \tag{1.33}
\end{align*}
$$

where

$$
\begin{align*}
f\left(t^{\prime}\right) & =t-t^{\prime}-\frac{1}{c}\left|x-x^{\prime}\right| \\
& =t-t^{\prime}-\frac{1}{c}\left[\left[x-R\left(t^{\prime}\right)\right]^{2}+y^{2}+z^{2}\right]^{1 / 2} \tag{1.34}
\end{align*}
$$

and

$$
\begin{equation*}
g\left(t^{\prime}\right)=\frac{x-R\left(t^{\prime}\right)}{c} \frac{1}{\left[x-R\left(t^{\prime}\right)\right]^{2}+y^{2}+z^{2}} \tag{1.35}
\end{equation*}
$$

We can evaluate this integral using integration by parts

$$
\begin{align*}
\int d t^{\prime} \delta^{\prime}\left(f\left(t^{\prime}\right)\right) g\left(t^{\prime}\right) & =\int d t^{\prime} \frac{d}{d t^{\prime}} \delta\left(f\left(t^{\prime}\right)\right) \frac{g\left(t^{\prime}\right)}{d f f f^{\prime}} \\
& =-\int d t^{\prime} \delta\left(f\left(t^{\prime}\right)\right) \frac{d}{d t^{\prime}}\left[\frac{g\left(t^{\prime}\right)}{f^{\prime}(t)}\right]  \tag{1.36}\\
& =-\left.\frac{1}{\mid d f / d t^{\prime} T_{t_{0}}} \frac{d}{d t^{\prime}}\left[\frac{g\left(t^{\prime}\right)}{f^{\prime}\left(t^{\prime}\right)}\right]\right|_{t=t_{0}},
\end{align*}
$$

where $t_{0}$ arises from $f\left(t_{0}\right)=0$ and hence solves (1.25). We thus have that

$$
\begin{equation*}
\phi_{a}(x, t)=\frac{1}{4 \pi} \frac{1}{T f^{\prime}\left(t_{0}\right) T}\left[\frac{g^{\prime}\left(t_{0}\right)}{f^{\prime}\left(t_{0}\right)}-\frac{g\left(t_{0}\right)}{\left[f^{\prime}\left(t_{0}\right)\right]^{2}} f^{\prime \prime}\left(t_{0}\right)\right] . \tag{1.37}
\end{equation*}
$$

To evaluate (1.37) define

$$
\begin{equation*}
D\left(t^{\prime}\right)=\left[\left[x-R\left(t^{\prime}\right)\right]^{2}+y^{2}+z^{2}\right]^{1 / 2} \tag{1.38}
\end{equation*}
$$

Then we can write

$$
f\left(t^{\prime}\right)=t-t^{\prime}-\frac{1}{c} D\left(t^{\prime}\right)
$$

and

$$
g\left(t^{\prime}\right)=\frac{1}{c}\left[x-R\left(t^{\prime}\right)\right] D^{-2}\left(t^{\prime}\right),
$$

so that

$$
\begin{aligned}
f^{\prime} & =-1+\frac{1}{c} \frac{R^{\prime}(x-R)}{D}=\frac{R^{\prime}(x-R)-c D}{c D}, \\
f^{\prime \prime} & =\left[R^{\prime}(x-R)-c D\right] \frac{(x-R) R^{\prime}}{c D^{3}}+\frac{1}{c D}\left[R^{\prime \prime}(x-R)-\left(R^{\prime}\right)^{2}+c \frac{R^{\prime}(x-R)}{D}\right] \\
& =\frac{R^{\prime \prime}(x-R)}{c D}-\frac{\left(R^{\prime}\right)^{2}}{c D}+\frac{\left[R^{\prime}(x-R)\right]^{2}}{c D^{3}},
\end{aligned}
$$

and

$$
\begin{aligned}
g^{\prime} & =\frac{1}{c}\left(-R^{\prime}\right) D^{-2}+\frac{x-R}{c}(-2) D^{-3}\left[\frac{-R^{\prime}(x-R)}{D}\right] . \\
& =\frac{-R^{\prime}}{c D^{2}}+\frac{2 R^{\prime}(x-R)^{2}}{c D^{4}} .
\end{aligned}
$$

Combining these we get

$$
\begin{align*}
\phi_{a}= & \frac{c^{2} R^{\prime} D}{4 \pi\left|R^{\prime}(x-R)-c D\right|^{3}} \\
& \cdot\left[1-\left[\frac{R^{\prime \prime}}{R^{\prime}}\right] \frac{(x-R)^{2}}{c D}-2 \frac{(x-R)^{2}}{D^{2}}+\frac{R^{\prime}(x-R)^{3}}{c D^{3}}\right] . \tag{1.39}
\end{align*}
$$

where each term is evaluated at $t^{\prime}=t_{0}$ with $t_{0}$ a solution to $f\left(t_{0}\right)=0$ or

$$
\begin{equation*}
t_{0}=t-\frac{1}{c}\left|x-R\left(t_{0}\right) T\right| \tag{1.40}
\end{equation*}
$$

with $D$ defined in (1.38). ©a has the dimensions (1ength) ${ }^{-2}$.
The second term in (1.31) is defined as

$$
\begin{align*}
\phi_{b}=\frac{-1}{4 \pi} \int_{0}^{t^{+}} d t^{\prime} \iiint d x^{\prime} & \delta\left(t-t^{\prime}-\frac{1}{c}\left|z-x^{\prime}\right|\right) \\
& \cdot \frac{\left(x-x^{\prime}\right)}{\left|x-x^{\prime}\right|^{3}} \delta\left(x^{\prime}-R\left(t^{\prime}\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}\right)\right. \tag{1.41}
\end{align*}
$$

Spatial integration yields

$$
\begin{equation*}
\rho_{b}=-\frac{1}{4 \pi} \int_{0}^{t} d t^{\prime} \delta\left(t-t^{\prime}-\frac{1}{c} D\left(t^{\prime}\right)\right) \frac{\left[x-R\left(t^{\prime}\right)\right]}{\left[D\left(t^{\prime}\right)\right]^{3}} . \tag{1.42}
\end{equation*}
$$

where $D\left(t^{\prime}\right)$ is defined in (1.38). To evaulate this use (1.22) to get

$$
\begin{align*}
\phi_{b} & =\frac{-1}{4 \pi} \frac{\left[x-R\left(t_{0}\right)\right]}{\left[D\left(t_{0}\right)\right]^{3}} \frac{1}{\left|d f / d t^{\prime}\right|_{t_{0}}} \\
& =-\frac{1}{4 \pi} \frac{c\left[x-R\left(t_{0}\right)\right] D\left(t_{0}\right)}{\left|R^{\prime}(x-R)-c D\right|} \frac{1}{\left[D\left(t_{0}\right)\right]^{3}} \\
& =-\frac{c}{4 \pi} \frac{x-R\left(t_{0}\right)}{D^{2}\left(t_{0}\right)\left|R^{\prime}(x-R)-c D\right|} \tag{1.43}
\end{align*}
$$

which al so has dimensions (length) ${ }^{-2}$. This can be rewritten as

$$
\begin{align*}
\phi_{b} & =-\frac{c^{2} R^{\prime} D}{4 \pi\left|R^{\prime}(x-R)-c D\right|^{3}}\left[\frac{\left[R^{\prime}(x-R)-c D\right]^{3}(x-R)}{c R^{\prime} D^{3}}\right] \\
& =-\frac{c^{2}}{4 \pi} \frac{R^{\prime} D(x-R)}{\left|R^{\prime}(x-R)-c D\right|^{3}}\left[\frac{R^{\prime}{ }^{2}(x-R)^{2}-2 c D R^{\prime}(x-R)+c^{2} D^{2}}{c R^{\prime} D^{3}}\right] \\
& =-\frac{c^{2}}{4 \pi} \frac{R^{\prime} D}{\left|R^{\prime}(x-R)-c D\right|^{3}}\left[\frac{c(x-R)}{R^{\prime} D}-\frac{2(x-R)^{2}}{D^{2}}+\frac{R^{\prime}(x-R)^{3}}{c D^{3}}\right] . \tag{1.44}
\end{align*}
$$

Thus $\phi_{1}$ from (1.31) and (1.32) is

$$
\phi_{1}=\phi_{a}+\phi_{b},
$$

and from (1.39) and (1.44) we get

$$
\begin{equation*}
\phi_{1}=\frac{c^{2}}{4 \pi} \frac{R^{\prime} D}{\left|R^{\prime}(x-R)-c D\right|^{3}}\left[1-\frac{x-R}{c R^{\prime} D}\left[R^{\prime \prime}(x-R)+c^{2}\right]\right] \tag{1.45}
\end{equation*}
$$

The first term in the brackets in (1.45) is the far field term and the second term a near-field term.

As a simple check on this result, suppose the dipole is fired at $x^{\prime}=0$, i.e. 1et $R\left(t^{\prime}\right)=0$. Then

$$
D=r=\left(x^{2}+y^{2}+z^{2}\right)^{1 / 2}
$$

and only the second term in (1.45) contributes. The result is

$$
\left.9_{1}\right|_{R=0}=-\frac{x}{4 \pi r^{3}}=\frac{1}{4 \pi} \frac{\partial}{\partial x}\left[\frac{1}{r}\right]
$$

which is the standard result for a fixed dipole source.
Dipoles moving in other directions can easily be computed using the form (1.45). We relable the sorrce in (1.29) indicating that it is in the $x$-direction as

$$
S_{x}\left(x^{\prime}, t^{\prime}\right)=\delta^{\prime}\left(x^{\prime}-R_{1}\left(t^{\prime}\right)\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}\right)
$$

If we al so define from (1.38)

$$
D_{1}\left(t^{\prime}\right)=\left[\left[x-R_{1}\left(t^{\prime}\right)\right]^{2}+y^{2}+z^{2}\right]^{1 / 2}
$$

Then (1.45) is for a moving dipole in the x-direction

$$
\phi_{1}^{D}=\frac{c^{2}}{4 \pi} \frac{R_{1}^{\prime} D_{1}}{\left|R_{1}^{\prime}\left(x-R_{1}\right)-c D_{1}\right|^{3}}\left[1-\frac{x-R_{1}}{C R_{1}^{\prime} D_{1}}\left[R_{1}^{n}\left(x-R_{1}\right)+c^{2}\right]\right]
$$

A moving dipole in the $y$-direction is

$$
S_{y}\left(x^{\prime}, t^{\prime}\right)=\delta\left(x^{\prime}\right) \delta\left(y^{\prime}-R_{2}\left(t^{\prime}\right)\right) \delta\left(z^{\prime}\right)
$$

and if we define

$$
D_{2}\left(t^{\prime}\right)=\left[x^{2}+\left[y-R_{2}\left(t^{\prime}\right)\right]^{2}+z^{2}\right]^{1 / 2}
$$

the result analogous to (1.45) is

$$
\phi_{1}^{D y}=\frac{c^{2}}{4 \pi} \frac{R_{2}^{\prime} D_{2}}{\left|R_{2}^{\prime}\left(y-R_{2}\right)-c D_{2}\right|^{3}}\left[1-\frac{y-R_{2}}{c R_{2}^{\prime} D_{2}}\left[R_{2}^{\prime \prime}\left(y-R_{2}\right)+c^{2}\right]\right]
$$

Similarly in the z-direction we have

$$
S_{z}\left(x_{\sim}^{\prime}, t^{\prime}\right)=\delta\left(x^{\prime}\right) \delta\left(y^{\prime}\right) \delta\left(z^{\prime}-R_{3}\left(t^{\prime}\right)\right)
$$

and

$$
D_{3}\left(t^{\prime}\right)=\left[x^{2}+y^{2}+\left[z-R_{3}\left(t^{\prime}\right)\right]^{2}\right]^{1 / 2}
$$

and the result we get is

$$
\phi_{1}^{D z}=\frac{c^{2}}{4 \pi} \frac{R_{3}^{\prime} D_{3}}{\left|R_{3}^{\prime}\left(z-R_{3}\right)-c D_{3}\right|^{3}}\left[1-\frac{z-R_{3}}{c R_{3}^{\prime} D_{3}}\left[R_{3}^{\prime \prime}\left(z-R_{3}\right)+c^{2}\right]\right]
$$

For the second term in (1.15) we write

$$
\begin{align*}
\phi_{2}(x, t)=\int_{0}^{t+} d t^{\prime} \iint d S^{\prime} & {\left[G\left(x, t ; x_{s}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial n}, \phi\left(x_{s}^{\prime}, t^{\prime}\right)\right.} \\
& \left.-\phi\left(\underset{\sim}{x_{s}^{\prime}}, t^{\prime}\right) \frac{\partial}{\partial n^{\prime}}, G\left(\underset{\sim}{x}, t ;{\underset{\sim}{s}}_{s}^{\prime}, t^{\prime}\right)\right] . \tag{1.46}
\end{align*}
$$

For the boundary condition we have either the Dirichlet boundary value problem

$$
\begin{equation*}
\phi\left(x_{N}^{\prime}, t^{\prime}\right)=0 \tag{1.47}
\end{equation*}
$$

or the Neumann boundary valne problem

$$
\begin{equation*}
\frac{\partial}{\partial n^{\prime}} p\left(x_{s}^{\prime}, t^{\prime}\right)=0 \tag{1.48}
\end{equation*}
$$

If we choose the Green's function such that it has the corresponding boundary value $G\left(x, t ; x_{s^{\prime}}^{\prime}, t^{\prime}\right)=0$ or $\partial G\left(x, t ; x_{s^{\prime}}^{\prime}, t^{\prime}\right) / \partial n^{\prime}=0$ then $\phi_{2}$ vanishes. The proviso with this convenient method is that one can find a Green's function which say vanishes on the boundary. It is usually only possible to find this image Green's function for simple geometries involving flat planes, cylinders, and spheres for example. For an arbitrary boundary this in effect would amount to fully solving the problem. In any case we get one term to drop by our choice of boundary condition on $\phi$ or $\partial \phi / \partial n^{\prime}$. To solve the problem we must write an integral equation on the remaining boundary value (either $\phi$ or $\partial \phi / \partial n^{\prime}$ ). We do this in the next section for the Helmholtz equation. We also treat, in the next section, an example for a flat interface where we can find an image Green's function. The result will be the Rayleigh-Sommerfeld diffraction formulae.

```
The third term in (1.15) is
```

$$
\begin{align*}
\phi_{3}(\underset{\sim}{x}, t)=\frac{1}{c^{2}} \iiint d_{\sim} & {\left[G\left(\underset{\sim}{x}, t ;{\underset{\sim}{x}}^{\prime}, 0\right) \frac{\partial}{\partial t}, \phi(\underset{\sim}{x}, 0)\right.} \\
& \left.-\phi(\underset{\sim}{x}, 0) \frac{\partial}{\partial t^{\prime}}, G\left(\underset{\sim}{x}, t,{\underset{\sim}{x}}^{\prime}, 0\right)\right], \tag{1.49}
\end{align*}
$$

in terms of initial conditions on the field, i.e. $\varphi\left(x^{\prime}, 0\right)$ and $\partial \phi\left(x^{\prime}, 0\right) / \partial t^{\prime}$. Again if we are able to choose the Green's function such that its initial values matched those of $p$, the term would vanish. Regardless of the choice of $G$, both initial values on $\phi$ mast be known, so once we specify the choice of $G$, this term is a known function. We specify the initial conditions on the retarded Green's fanction 1ater in this section.

To summarize, we have that

$$
\begin{equation*}
\phi(x, t)=\phi_{1}(x, t)+\phi_{2}(x, t)+\phi_{3}(x, t), \tag{1.50}
\end{equation*}
$$

with $\phi_{1}$ given by (1.17), $\phi_{2}$ by (1.46) and $\phi_{3}$ by (1.49). $\phi_{1}$ is known if $G$ and $S$ are known. $\phi_{2}$ is only partially known since either $\phi\left(X_{s}^{\prime}, t^{\prime}\right)$ or $\partial \rho\left(x_{s}^{\prime}, t^{\prime}\right) / \partial n^{\prime}$ is known but not both. $\phi_{3}$ is known if $G$ and both initial conditions $\phi\left(x^{\prime}, 0\right)$ and $\partial \phi\left(x^{\prime}, 0\right) / \partial t^{\prime}$ are known.

### 2.1.1 FOLL (BODNDARY) GRERN'S PONCTION

Rather than find the field fanction $\phi$ due to a general source term $S$, we instead find the Green's function $F$ due to a point source, i.e. we want to solve for the function $P$, the full Green's function, which satisfies the equation

$$
\begin{equation*}
\left[\forall^{\prime 2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime}}\right] I^{\prime}\left(x^{\prime}, t^{\prime}, x^{\prime \prime}, t^{\prime \prime}\right)=-\delta\left(x^{\prime}-x^{\prime \prime}\right) \delta\left(t-t^{\prime \prime}\right) \tag{1.51}
\end{equation*}
$$

This is the same as (1.10) if we replace $\phi$ by $F$ and $S$ by the point source on the rhs of (1.51). Equation (1.15) then becomes

$$
\begin{align*}
& P\left(x, t, x_{\sim}^{\prime \prime}, t^{\prime \prime}\right)=G\left(x, t ; x^{\prime \prime}, t^{\prime \prime}\right) \\
& +\int_{\delta}^{t+} d t^{\prime} \iint S^{\prime}\left[G\left(\underset{\sim}{x}, t ; X_{s}^{\prime}, t^{\prime}\right) \frac{\partial}{\partial n^{\prime}} P\left(\underset{\sim}{x^{\prime}}, t^{\prime} ; x_{\sim}^{\prime \prime}, t^{\prime \prime}\right)\right. \\
& \left.-\mathbb{P}\left(\underset{\sim}{x} s^{\prime}, t^{\prime},{\underset{\sim}{x}}^{\prime \prime}, t^{\prime \prime}\right) \frac{\partial}{\partial n^{\prime}} G\left(\underset{\sim}{x}, t,{\underset{\sim}{x}}_{s^{\prime}}, t^{\prime}\right)\right] \\
& +\frac{1}{c^{2}} \iiint d_{x^{\prime}}\left[G\left(\underset{\sim}{x}, t,{\underset{\sim}{x}}^{\prime}, 0\right) \frac{\partial}{\partial t} P\left({\underset{\sim}{x}}^{\prime}, 0, \underset{\sim}{x^{\prime \prime}}, t^{n}\right)\right. \\
& \left.-T\left(x_{\sim}^{\prime}, 0, x^{\prime \prime}, t^{\prime \prime}\right) \frac{\partial}{\partial t^{\prime}} G\left(\underset{\sim}{x}, t,{\underset{\sim}{x}}^{\prime}, 0\right)\right] \quad . \tag{1.52}
\end{align*}
$$

which is the integral representation for the full Green's function of the
problem. Once we know its solution we can find the solution for any source $S\left(x^{\prime}, t^{\prime \prime}\right)$ by maltiplying (1.52) by $S\left(x^{\prime \prime}, t^{\prime \prime}\right)$ and integrating over $x^{\prime \prime}$ and $t^{\prime \prime}$. For example, the result due to a source $S$ is just

$$
\begin{equation*}
\phi(x, t)=\int d t^{n} \iiint \int_{\sim}^{n} \underset{\sim}{x}\left(\underset{\sim}{x}, t_{i}{\underset{\sim}{x}}^{n}, t^{n}\right) S\left({\underset{\sim}{x}}^{n}, t^{n}\right) \tag{1.53}
\end{equation*}
$$

Our specification of the initial-boundary value problem is analogous to before. We assume $G$ is known, and specify one boundary condition on either I or its normal derivative, as well as both initial conditions on $\mathrm{F}_{\mathrm{t}}$. The full solution of the problem requires us to solve an integral equation on the remaining value evalnated on the surface. Note that if $I$ and $G$ satisfy the same initial and boundary conditions, both integral terms vanish, and $\mathbf{I}=G$ which is the complete solution of the problem.

### 2.1.2 INITIAL CONDITIONS ON $\underbrace{(3,1)}$

From Ch. 1, Eq. (1.12), we have a representation valid for any Green's function solving the wave equation. It is written in terms of the pole shifts and is given by $\left(x=\left(x, x_{0}\right), \omega_{k}=|\underline{k}|\right)$

$$
\begin{align*}
G^{(3,1)}(x)= & \frac{-1}{(2 \pi)^{4}}\left[P \iiint \int \frac{e^{i k \cdot \frac{x}{\sim}} e^{-i k_{0} x_{0}}}{k_{0}^{2}-\omega_{k}^{2}} d k d k_{0}\right. \\
& \left.+\frac{\pi i}{2} \iiint \frac{e^{i k \cdot \frac{\Sigma}{\sim}}}{\omega_{k}}\left[a e^{-i \omega_{k} x_{0}}-\beta e^{i \omega_{k} x_{0}}\right] d k\right] \tag{1.54}
\end{align*}
$$

The initial condition on $G^{(3,1)}$ is specified at $x_{0}=0$ so that

$$
\begin{align*}
G^{(3,1)}(\underset{\sim}{x}, 0)=-\frac{1}{(2 \pi)^{4}}[ & P \iiint \int \frac{e^{i k \cdot \frac{x}{\sim}}}{k_{0}^{2}-\omega_{k}^{2}} d k \underset{\sim}{d k_{0}} \\
& \left.+\frac{\pi i}{2}(\alpha-\beta) \iiint \frac{e^{i k \cdot \frac{I}{\sim}}}{\omega_{k}} d \underset{\sim}{k}\right] \tag{1.55}
\end{align*}
$$

We can evaluate the $k_{0}$ integral in the first term directly

$$
\begin{equation*}
P \int \frac{d k_{0}}{k_{0}^{2}-\omega_{k}^{2}}=\pi i\left[\frac{1}{2 \omega_{k}}+\frac{1}{\left(-2 \omega_{k}\right)}\right]=0 \tag{1.56}
\end{equation*}
$$

so that it always vanishes. The second term vanishes if $\alpha=\beta$, $i$. e. for the retarded (R), advanced (A), and principal value (P) Green's functions. So we have the initial conditions

$$
\begin{equation*}
G^{(3,1)}(\underset{\sim}{x}, 0)=0 \quad(R, A, P) \tag{1.57}
\end{equation*}
$$

The time derivative of $\mathbf{G}^{(3,1)}$ is from (1.54)

$$
\begin{align*}
\partial_{0} G^{(3,1)}(x)= & -\frac{1}{(2 \pi)^{4}}\left[P \iiint \int \frac{e^{i k \cdot \frac{x}{\sim}\left(-i k_{0}\right)} e^{-i k_{0} x_{0}}}{k_{0}^{2}-\omega_{k}^{2}} d k d k_{0}\right. \\
& \left.+\frac{\pi i}{2} \iiint \frac{e^{i k \cdot \frac{x}{\sim}}}{\omega_{k}}\left[-i \omega_{k}\right]\left[\alpha e^{i \omega_{k} x_{0}}+\beta e^{i \omega_{k} x_{0}}\right] d k\right] . \tag{1.58}
\end{align*}
$$

Next set $x_{0}=0$. The first term vanishes identically because the integrand of the $k_{0}$-integral is an odd function. The second integral becomes

$$
\begin{equation*}
\frac{\pi}{2}(\alpha+\beta) \iiint e^{i k \cdot x} d k=\frac{\pi}{2}(2 \pi)^{3}(\alpha+\beta) \delta(x) \tag{1.59}
\end{equation*}
$$

so that

$$
\begin{equation*}
\partial_{0} G^{(3,1)}(\underset{\sim}{x}, 0)=-\frac{1}{4}(\alpha+\beta) \delta(\underset{\sim}{x}) \tag{1.60}
\end{equation*}
$$

For the retarded Green's function $\alpha=\beta=-1$, for the advanced Green's function $\alpha=\beta=1$, and for the principal value Green's function $\alpha=\beta=0$, so we have

$$
\partial_{0} G^{(3,1)}(\underset{\sim}{x}, 0)=\left[\begin{array}{l}
1 / 2 \delta(\underset{\sim}{x})  \tag{R}\\
-1 / 2 \\
\\
0
\end{array}\right.
$$

By (1.16) we thus have that for $t=t^{\prime}=0$

$$
\begin{equation*}
G\left(\underset{\sim}{x}, 0 ;{\underset{\sim}{x}}^{\prime}, 0\right)=c G_{R}^{(3,1)}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}, 0\right)=0 \tag{1.62}
\end{equation*}
$$

and that

$$
\begin{align*}
\left.\frac{\partial}{\partial t^{\prime}} G\left(\underset{\sim}{x}, t,{\underset{\sim}{x}}^{\prime}, t^{\prime}\right)\right|_{t=t^{\prime}=0} & =\partial_{0} G_{R}^{(3,1)}\left(\underset{\sim}{\left.x-x^{\prime}, 0\right)}\right. \\
& =1 / 2 \delta\left(\underset{\sim}{x}-x_{\sim}^{\prime}\right) \tag{1.63}
\end{align*}
$$

Note that in our integral representation (1.15) the initial condition is set at $t^{\prime}=0$ for the source. This was a matter of choice and led us to integrate the representation from 0 to $\infty$. The integration was reduced by cansality. The initial condition on the field was thus at $t=0$, expressed under the integral by $\phi\left(x^{\prime}, 0\right)$ bat on the field fanction as $\phi(x, 0)$ for example.

We can derive these initial conditions another way. We have specific functional forms for the Green's functions which enable us to derive the se results directly. For example from Ch. 1 we have

$$
\begin{equation*}
G_{R}^{(3,1)}\left(x, x^{\prime}\right)=\frac{\delta(r-t)}{4 \pi r}, \tag{1.64}
\end{equation*}
$$

where

$$
x=\left(x, x_{0}\right), \quad r=\left|x-x_{N}^{\prime}\right|, \quad \text { and } \quad \tau=c\left(t-t^{\prime}\right)
$$

Evaluating this function directly we have

$$
\begin{equation*}
\left.G_{R}^{(3,1)}\left(x, x^{\prime}\right)\right|_{\tau=0}=\frac{\delta(r)}{4 \pi r}, \tag{1.65}
\end{equation*}
$$

and

$$
\begin{equation*}
\left.\partial_{0} G_{R}^{(3,1)}\left(x, x^{\prime}\right)\right|_{\tau=0}=-\frac{\delta^{\prime}(r)}{4 \pi r} \tag{1.66}
\end{equation*}
$$

If we evaluate these terms as distributions and integrate over all space we get $f$ or the rhs of (1.65)

$$
\begin{equation*}
\int_{0}^{\infty} r^{2} d r \int_{0}^{\pi} \sin \theta d \theta \int_{0}^{2 \pi} d \phi \frac{\delta(r)}{4 \pi r}=0, \tag{1.67}
\end{equation*}
$$

and for the rhs of (1.66)

$$
\begin{align*}
& \int_{0}^{\infty} r^{2} d r \int_{0}^{\pi} \sin \theta d \theta \int_{0}^{2 \pi} d \phi\left[-\frac{\delta^{\prime}(r)}{4 \pi r}\right] \\
& =-\int_{0}^{\infty} r \delta^{\prime}(r) d r=\int_{0}^{\infty} \delta(r) d r=1 / 2 . \tag{1.68}
\end{align*}
$$

Note that if we wrote $G_{R}^{(3,1)}$ with a step function $\theta(\tau)$, the derivative contains $\delta(r) / 4 \pi r$ which vanishes as a distribution in three dimensions.

Hence as distributions
(a) $\delta(r) / 4 \pi r$ is equivalent to zero in three dimensions and
(b) $-\delta^{\prime}(x) / 4 \pi r$ is equivalent to $1 / 2 \delta(x)$ in three dimensions
so that our result is for G

$$
\begin{aligned}
& G\left(\underset{\sim}{x}, 0 ; x^{\prime}, 0\right)=0 \\
& \partial_{t}, G\left(\underset{\sim}{x}, 0 ; x^{\prime} 0\right)=1 / 2 \delta\left(x-x^{\prime}\right)
\end{aligned}
$$

as before.

### 2.2 BRI,MHOLTZ EQUATION

In this section we construct the integral representation for the solution of the Helmholtz equation as well as the integral equations on a surface field value necessary to solve the boundary value problem. We do it in a different way from the wave equation, by using index notation. We assume the field $\phi$ satisfies a Helmholtz equation with a source $S$

$$
\begin{equation*}
\left(\partial_{j}^{j} \partial_{j}^{0}+k_{0}^{2}\right) \phi\left(x_{\sim}^{\prime}\right)=-S\left({\underset{\sim}{x}}^{\prime}\right), \tag{2.1}
\end{equation*}
$$

and the equation for the Green's function satisfies the same equation but with a delta function source

$$
\begin{equation*}
\left(\partial_{j}^{\prime} \partial j+k_{0}^{2}\right) G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=-\delta\left(\underset{\sim}{x}-x^{\prime}\right) . \tag{2.2}
\end{equation*}
$$

Note that the differential operators in (2.2) operate on the source coordinate, which is explicitly permitted by reciprocity.

We assume that $G^{(3)}$ is a known function and that $\rho$ satisfies certain boundary conditions which we specify later. Also here we assume that the differential equation (2.1) and source $S$ are valid and exist only in a halfspace. We assume perfectly reflecting boundary conditions on this halfspace illustrated below.


Fig. 2.2
Next cross multiply the equations to form

$$
\begin{align*}
& G^{(3)}\left(\underset{d}{ }, x^{\prime}\right)\left(\partial_{j}^{\prime} \partial_{j}^{\prime}+k_{0}^{2}\right) \phi\left(x^{\prime}\right) \\
& -\left[\left(\partial_{j}^{\prime} \partial_{j}^{\prime}+k_{0}^{2}\right) G^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)\right] \phi\left(x^{\prime}\right)=G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right) S\left(x_{\sim}^{\prime}\right) \\
& +\phi\left(x^{\prime}\right) \delta\left(\underset{\sim}{x}-x^{\prime}\right) . \tag{2.3}
\end{align*}
$$

The $k_{0}^{2}$ terms cancel and the lefthand side can be witten as a divergence. The result is

$$
\begin{equation*}
\phi\left(x_{\sim}^{\prime}\right) \delta\left(\underset{\sim}{x}-x_{N}\right)=G^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right) S\left(x_{\sim}^{\prime}\right)+\partial_{j} F_{j}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right) \tag{2.4}
\end{equation*}
$$

where

$$
\begin{equation*}
F_{j}\left(x, x^{\prime}\right)=G^{(3)}\left(\underset{\sim}{x}, x^{\prime}\right) \partial_{j}^{\prime} \phi\left(x^{\prime}\right)-\left[\partial_{j} G\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right)\right] \phi\left(x^{\prime}\right) \quad . \tag{2.5}
\end{equation*}
$$

If we multiply (2.4) by the step function

$$
\theta\left(z^{\prime}-h\left(\underset{\sim}{x_{t}^{\prime}}\right)\right),
$$

and integrate the result over $x^{\prime}$, the result is restricted to $V_{1}$ in the figure by the step function, and we get

$$
\begin{align*}
\phi(x) \theta\left(z-h\left({\underset{\sim}{x}}_{t}\right)\right)= & \iiint_{\sim} G^{(3)}\left(\underset{\sim}{x}, x_{\sim}^{\prime}\right) S\left(x_{N}^{\prime}\right) d x_{N}^{\prime} \\
& +\iiint_{j}\left[\partial{ }_{j}^{\prime} F_{j}\left(x, x_{\sim}^{\prime}\right)\right] \theta\left(z^{\prime}-h\left(x_{i}^{\prime}\right)\right) d x_{\sim}^{\prime} . \tag{2,6}
\end{align*}
$$

where the source integral is restricted even further by its support (which must be in $V_{1}$ ). The latter integral in (2.6) can be integrated by parts to give

$$
\begin{align*}
& \iiint\left[\partial_{j}^{\prime} F_{j}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)\right] \theta\left(z^{\prime}-h\left(\underset{\sim}{x_{t}^{\prime}}\right)\right) d{\underset{\sim}{x}}^{\prime} \\
& \quad=-\iiint F_{j}\left(\underset{\sim}{x}, x^{\prime}\right) \partial_{j}^{\prime} \theta\left(z^{\prime}-h\left({\underset{\sim}{x}}_{t}^{\prime}\right)\right) d x^{\prime} \tag{2.7}
\end{align*}
$$

It can easily be seen that since the integral is over all space, each integrated term vanishes. For example the $j=1$ integrated term is

$$
\left.F_{1}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right) \theta\left(z^{\prime}-h\left({\underset{\sim}{x}}_{t}^{\prime}\right)\right)\right|_{x^{\prime}=-\infty} ^{x=+\infty} \quad=0
$$

The step function derivative in (2.7) is

$$
\begin{equation*}
\partial_{j}^{\prime} \theta\left(z^{\prime}-h\left(\underset{\sim}{x^{\prime}}\right)\right)=\delta\left(z^{\prime}-h(\underset{\sim}{x})\right) n_{j}\left(x_{t}^{\prime}\right) \tag{2.8}
\end{equation*}
$$

Where the delta function is the characteristic function of the surface and $\mathbf{n}_{\mathbf{j}}$ is

$$
\begin{equation*}
n_{j}\left(x_{\sim}^{\prime}\right)=\delta_{j 3}-\partial_{j t} h\left({\underset{\sim}{x}}_{t}^{\prime}\right) \tag{2.9}
\end{equation*}
$$

which is a vector in the direction of the surface normal (but not a unit vector) and the derivative terms for $j=1$ and $j=2$ are

$$
\begin{equation*}
h_{x^{\prime}}=\frac{\partial}{\partial x^{\prime}} h\left({\underset{\sim}{x}}^{\prime}\right) \quad, \quad h_{y^{\prime}}=\frac{\partial}{\partial y^{\prime}} h\left({\underset{\sim}{x}}_{t}^{\prime}\right) \tag{2.10}
\end{equation*}
$$

which are the surface slopes. We assume the surface is differentiable. The result yields for (2.6)

$$
\begin{equation*}
\phi(x) \theta\left(z-h\left(x_{i}^{\prime}\right)\right)=\phi^{i n}(x)+\phi^{s c}(\underset{\sim}{x}) \tag{2.11}
\end{equation*}
$$

with the total field (on the lhs) given in terms of the incident field

$$
\begin{equation*}
\phi^{\text {in }}(x)=\iiint_{\nabla_{1}} G^{(3)}\left(x, x^{\prime}\right) S\left(x^{\prime}\right) d x^{\prime}, \tag{2.12}
\end{equation*}
$$

and the scattered field

$$
\begin{equation*}
\phi^{s c}(\underset{\sim}{x})=-\iiint F_{j}\left(\underset{\sim}{x}, x^{\prime}\right) n_{j}\left({\underset{\sim}{x}}^{\prime}\right) \delta\left(z^{\prime}-h\left(x_{i}^{\prime}\right)\right) d x^{\prime} . \tag{2.13}
\end{equation*}
$$

We can evaluate the delta function in (2.13) by setting the vector $\underset{\sim}{x}$ on the surface, $\underset{\sim}{x}{ }_{s}^{\prime}=\left({\underset{\sim}{x}}^{t}, h\left({\underset{\sim}{x}}_{t}\right)\right)$, to yield

$$
\begin{equation*}
\phi^{s c}(x)=-\iint F_{j}\left(\underset{\sim}{x} \cdot \frac{x_{s}^{\prime}}{d} n_{j}\left({\underset{\sim}{x}}_{\prime}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime} .\right. \tag{2.14}
\end{equation*}
$$

As an example, if $S$ is a point source

$$
S\left(x_{\sim}^{\prime}\right)=\delta\left(x^{\prime}-{\underset{\sim}{n}}^{\prime \prime}\right) .
$$

the incident field is that field due to the point source evaluated at the field point $\underset{\sim}{\text { x. }}$ Since $G^{(3)}$ satisfies our outgoing radiation condition it is

$$
\phi^{i n}(\underset{\sim}{x})=G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n \prime}\right)
$$

In general we can write the integral representation for $\phi$ with $z e V_{1}$ as (from (2.11), (2.13) and (2.5))

$$
\begin{equation*}
\phi(x)=\phi^{i n}(\underset{\sim}{x})+\iint\left[N^{(3)}\left(x, x^{\prime}\right) \phi\left(x_{s}^{\prime}\right)-G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{0}^{\prime}\right) N\left(x_{\sim}^{\prime}\right)\right] d x_{t}^{\prime} \tag{2.15}
\end{equation*}
$$

in terms of the normal derivative of the field evaluated on the surface

$$
\begin{equation*}
N\left(\underset{\sim}{x}{ }_{s}^{\prime}\right)=n_{j}\left({\underset{\sim}{x}}_{t}^{\prime}\right) \partial_{j}^{\prime} \phi\left(\underset{\sim}{x_{s}^{\prime}}\right) \tag{2.16}
\end{equation*}
$$

and the normal derivative of the Green's function

$$
\begin{equation*}
N^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}^{\prime}\right)=n_{j}\left(x_{i}^{\prime}\right) \partial_{j}^{\prime} G^{(3)}\left(\underset{\sim}{x}, x_{N}^{\prime}\right) \tag{2.17}
\end{equation*}
$$

The representation (2.15) is called the Helmholtz-Kirchhoff Representation of the field. Our boundary value problem consists in specifying either $N$ or $\phi$ on the surface and then constructing an integral equation on the remaining boundary value. There are several ways to do this which we now describe.

## Eg. 1. FIRST EIND EQUATION FOR N

We assume that $\rho$ satisfies a Dirichlet boundary condition on the surface, i.e.

$$
\begin{equation*}
\phi\left(\mathrm{x}_{\mathrm{s}}\right)=0 \tag{2.18}
\end{equation*}
$$

From (2.15) we thus have that in the 1 imit as $z \rightarrow h(\underset{\sim}{x})$ (where $z \varepsilon V_{1}$ ) so that $x \rightarrow x^{\prime}$, the 1 hs vanishes and we get

$$
\begin{equation*}
\phi^{i n}\left(z_{s}\right)=\iint G^{(3)}\left(x_{s}, x_{s}^{\prime}\right) N\left(x_{s}^{\prime}\right) d x_{t}^{\prime} \tag{2.19}
\end{equation*}
$$

which is an integral eguation of first kind for N. Both $\phi^{i n}$ and $G^{(3)}$ are known, and, as we noted in Ch. 1, $G^{(3)}$ is continuous at the boundary. The square root term in its denominator is an integrable singularity. Once we solve (2.19) for $N$, we substitute the result into (2.15) using (2.18) to yield the field representation

$$
\begin{equation*}
\phi(x)=\phi^{i n}(x)-\iint G^{(3)}\left(x, x_{s}^{\prime}\right) N\left(x_{s}^{\prime}\right) d x_{t}^{\prime} . \tag{2.20}
\end{equation*}
$$

## Bg. 2. SBCOND EIND BQUATION ROR \&

\$satisfies the Neumann boundary condition on the surface, i.e.

$$
\begin{equation*}
N(\underset{\sim}{x})=0 . \tag{2.21}
\end{equation*}
$$

In the 1 imit as $z \rightarrow h\left(x_{t}\right), x \rightarrow \underset{\sim}{x}$ the 1 hs of (2.15) goes to $\phi\left(x_{\sim}\right)$. To find the 1 imit of the normal derivative of the Green's function in (2.15) we recall from Ch. 1 Sec. 5 that we can write

$$
\begin{equation*}
\partial_{j}^{\prime} G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}^{\prime}\right)=-\frac{1}{2} R_{j}\left(\underset{\sim}{x}-{\underset{\sim}{x}}_{\prime}^{\prime}\right)+\frac{1}{2} \delta_{j 3} \operatorname{sgn}\left(z^{\prime}-h\left({\underset{\sim}{x}}_{\prime}^{\prime}\right)\right) \delta\left({\underset{\sim}{x}}-\frac{x_{i}^{\prime}}{\sim}\right) . \tag{2.22}
\end{equation*}
$$

which is analogous to Eq. (5.20) in Ch. 1 except that here we are differentiating on the source coordinate in $G^{(3)}$ and we thas have an overall minus sign. The 1 imit of the integral resulting from (2.15) thas yields

$$
\frac{1}{2} \phi\left(x_{s}\right)=\phi^{i n}\left({\underset{\sim}{x}}_{s}\right)-\frac{1}{2} \iint P\left({\underset{\sim}{x}}_{s} x_{\sim}^{\prime}\right) \phi\left({\underset{\sim}{x}}_{s}\right) d{\underset{\sim}{x}}_{\prime}^{\prime} .
$$

or

$$
\begin{equation*}
\phi\left(x_{s}\right)=2 \phi^{i n}\left(x_{s}\right)-\iint P\left(x_{s} \cdot x_{s}^{\prime}\right) \phi\left(x_{s}^{\prime}\right) d x_{t}^{\prime} . \tag{2.23}
\end{equation*}
$$

where the function $P$ is defined as
in terms of the regular part $\mathbf{R}_{\mathbf{j}}$. The latter is defined in Eq. (5.21) in Ch. 1. The result, (2.23), is an integral equation of second kind for $\phi$. The result when solved is substituted back into (2.15) to give with (2.21) the resulting field expression for $\phi$

$$
\begin{equation*}
\phi(x)=\phi^{i n}(\underset{\sim}{x})+\iint N^{(3)}\left(x, x_{N}^{\prime}\right) \phi\left(x_{N}^{\prime}\right) d x_{N}^{\prime} . \tag{2.25}
\end{equation*}
$$

A Born approximation to (2.23) (i.e. neglecting the integral term) illustrates that for a vanishing normal derivative on the surface, the field on the surface is twice the incident field. This is also equivalent to a Kirchhoff approximation for a reflection coefficient equal to one. There is a well developed theory for solving integral equations of second kind. First kind equations are more difficult to solve in general. (Ref s. 2.2, 2.3 and 2.4.)

We have found an integral equation of first kind for $N$ for the Dirichlet problem, (2.19), and a second kind equation for $p$ for the Nemann problem, (2.23). We can also find a second kind equation for $N$ and a first kind equation for $\rho$. To do this, differentiate (2.15) and maltiply by the normal to get

$$
\begin{align*}
n_{m}\left(x_{t}\right) \partial_{m} \phi(x)= & n_{m}\left(x_{N}\right) \partial_{m} \phi^{i n}\left(x_{\sim}\right) \\
& +\iint\left[\eta_{m}\left(x_{t}\right) \partial_{m} N^{(3)}\left(\underset{\sim}{x}, x_{N}^{\prime}\right) \phi\left(x_{s}^{\prime}\right)\right. \\
& \left.-n_{m}\left(x_{t}\right) \partial_{m} G^{(3)}\left(x, x_{N}^{\prime}\right) N\left(x_{N}^{\prime}\right)\right] d x_{\sim}^{\prime} \quad . \tag{2.26}
\end{align*}
$$

## Eg. 3. SBCOND EIND BQDATION ON N

Let $p$ satisfy the Dirichlet boundary condition (2.18). Substitute the result in (2.26) and take the surface 1 imit as $\underset{\sim}{\boldsymbol{x}} \underset{\sim}{x} s^{\circ}$. The normal derivative of $G^{(3)}$ produces a regular part plus a jump discontinuity. Here the differentiation is on the field variable of $G^{(3)}$ and the result (5.20) applies. The resuting 1 imit of (2.26) is

$$
\begin{equation*}
N\left(x_{s}\right)=2 N^{i n}\left(x_{s}\right)-\iint P\left({\underset{\sim}{x}}_{s},{\underset{\sim}{x}}_{\prime}^{\prime}\right) N\left({\underset{\sim}{x}}_{\prime}^{\prime}\right) d{\underset{\sim}{t}}_{\prime}^{\prime} . \tag{2.27}
\end{equation*}
$$

where $N\left({\underset{\sim}{x}}_{s}\right)$ is defined by (2.16), $N^{i n}$ is defined by

$$
\begin{equation*}
N^{i n}\left({\underset{\sim}{x}}_{s}\right)=n_{m}\left(x_{t}\right) \partial_{m} q^{i n}\left(x_{s}\right) \tag{2.28}
\end{equation*}
$$

and $\bar{P}$ differs from $P$ in (2.24) in that the normal is a function of the exterior variable, i.e.

$$
\begin{equation*}
\bar{P}\left(x_{s} \cdot x_{s}^{\prime}\right)=n_{m}\left({\underset{\sim}{x}}_{t}\right) R_{m}\left({\underset{\sim}{x}}_{s}-x_{s}^{0}\right), \tag{2.29}
\end{equation*}
$$

defined in terms of the regalar part $R_{m}$ in (5.20) in Ch. 1. The result (2.27) is an integral eguation of second kind for N. Its Born approximation is that $N$ on the surface is just twice the normal derivative of the incident field, and this is the same as the Kirchhoff approximation for a reflection coefficient equal to -1.

## Eg. 4. FIRST KIND EQDATION FOR

For the Nemann boundary condition (2.21) the 1 imit as $\boldsymbol{x}_{\mathrm{s}} \rightarrow \mathrm{z}_{\mathrm{s}}$ of the 1 hs of (2.26) vanishes as does the second term in the integral. Recall from Ch. 1 that we can regularize the second derivative of $G^{(3)}$ as

$$
\begin{align*}
& \partial_{m} \partial_{j} G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=R_{m j}\left(\underset{\sim}{x}-x_{N}{ }^{\prime}\right) \\
& +\frac{1}{2} \operatorname{sgn}\left(z-z^{\prime}\right)\left[\delta_{m 3} \partial_{j t}+\delta_{j 3} \partial_{m t}\right] \delta\left({\underset{\sim}{x}}_{t}-x_{i}^{\prime}\right), \tag{2.30}
\end{align*}
$$

where here we are differentiating once with respect to the source (primed) variable and once with respect to the field (unprimed) variable (see (2.16)). The regular part $R_{m j}$ is defined in (5.55) of Ch. 1. The resulting

1imit of (2.26) becomes

$$
\begin{align*}
N^{i n}\left({\underset{\sim}{x}}_{s}\right)= & -\iint Q\left({\underset{\sim}{x}}_{s}{\underset{\sim}{x}}_{s}^{\prime}\right) \phi\left({\underset{\sim}{x}}_{\prime}^{\prime}\right) d{\underset{\sim}{x}}_{\prime}^{\prime} \\
& -\frac{1}{2} \iint\left[\left(n_{j t}\left({\underset{\sim}{x}}_{\prime}^{\prime}\right) \partial_{j t}+n_{m}\left(x_{n}\right) \partial_{m t}\right) \delta\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{\prime}^{\prime}\right)\right] \phi\left({\underset{\sim}{x}}_{s}^{\prime}\right) d x_{t}^{\prime} ; \tag{2.31}
\end{align*}
$$

where

$$
\begin{equation*}
Q\left({\underset{\sim}{x}}_{s} \cdot \frac{x_{s}^{\prime}}{\sim}\right)=n_{m}(\underset{\sim}{x} t) R_{m j}\left({\underset{\sim}{x}}_{s}-{\underset{\sim}{x}}_{s}^{\prime}\right) n_{j}\left({\underset{\sim}{x}}_{t}^{\prime}\right) \tag{2.32}
\end{equation*}
$$

The delta fanction terms in (2.31) all vanish because they integrate to the normal derivative of the field evaluated on the surface and this is assumed to vanish. The final result is an integral eguation of first kind on $\phi$,

$$
\begin{equation*}
-N^{i n}\left({\underset{\sim}{x}}_{s}\right)=\iint Q\left({\underset{\sim}{x}}_{s} x_{s}^{\prime}\right) \phi\left(x_{s}^{\prime}\right) d x_{t}^{\prime} \tag{2.33}
\end{equation*}
$$

### 2.3 BAYLBIGR-SOMMERPRLD DIFFRACTION FORMOLAE

From (2.15) if we assume that $S=0$ so $\phi^{i n}=0$ (zero source condition), and that we have a problem with a flat geometry, it is possible to find a Green's function using an image source which either vanishes on the (flat) boundary or whose normal (i.e.z) derivative vanishes. Our free-space retarded Green's function is

$$
G_{R}^{(3)}\left(x, x^{\prime}\right)=\frac{e^{i k_{0}\left|x^{-} x^{\prime}\right|}}{4 \pi\left|x^{\prime}-x^{\prime}\right|} .
$$

Boundary Green's functions can be witten as

$$
\begin{equation*}
G_{ \pm}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right) \pm G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{x}{i}}_{0}^{\prime}\right) \tag{3.2}
\end{equation*}
$$

where $\underset{\sim}{x}=(x, y, z), x^{\prime}=\left(x^{\prime}, y^{\prime}, z^{\prime}\right)$, and $\dot{X}_{i}^{\prime}=\left(x^{\prime}, y^{\prime},-z^{\prime}\right)$. It is easily seen that on the $z^{\prime}=0$ plane

$$
\begin{equation*}
G_{-}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{t}^{\prime}\right)=0 \tag{3.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial}{\partial z^{\prime}} G_{+}\left(x, x_{i}^{\prime}\right)=0 \tag{3.4}
\end{equation*}
$$

Similarly it can be shown that for $z^{\prime}=0$

$$
\begin{equation*}
G_{+}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{t}^{\prime}\right)=2 G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{t}^{\prime}\right), \tag{3.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\partial_{j}^{\prime} G_{-}\left(x_{\sim}, x_{t}^{\prime}\right) n_{j}\left(x_{\sim}^{\prime}\right)=\frac{\partial}{\partial z^{\prime}} G_{-}\left(x_{\sim}, x_{t}^{\prime}\right)=-2 \frac{\partial}{\partial z} G_{R}^{(3)}\left(x_{\sim}, x_{t}^{\prime}\right), \tag{3.6}
\end{equation*}
$$

where the latter relation in (3.6) is written in terms of differentiation on the field coordinate $z$.

If we choose $G^{(3)}=G_{+}$in (2.15) we get the result that, for the flat
boundary, $\left(\partial G_{+} / \partial z^{\prime}=N^{(3)}=0\right)$ the fieldrepresentation is

$$
\begin{equation*}
\phi(\underset{\sim}{x})=-2 \iint_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{t}^{\prime}\right) N\left({\underset{\sim}{x}}_{t}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime} \tag{3.7}
\end{equation*}
$$

where

$$
\begin{equation*}
N\left({\underset{\sim}{x}}_{t}^{\prime}\right)=\frac{\partial}{\partial z}, \phi\left({\underset{\sim}{t}}_{t}^{\prime}\right) \tag{3.8}
\end{equation*}
$$

If we choose $G^{\left(3^{\prime}\right)}=G_{\text {_ }}$ in (2.15) so that on the f1at surface $G^{(3)}=G_{-}=0$ and $N^{(3)}=-2 \partial G_{R}^{(3)} / \partial z$ we get the result

$$
\begin{equation*}
\phi(\underset{\sim}{x})=-2 \frac{\partial}{\partial z} \iint G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{t}^{\prime}\right) \phi\left({\underset{\sim}{x}}_{t}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime} \tag{3.9}
\end{equation*}
$$

Equations (3.7) and (3.9) are the Rayleigh-Sommerfeld diffraction formulae which gields the field value in terms of either boundary conditions on $N$ or p. Note that the formalae are not usefnl for Dirichlet or Nemann type problems. (Refs. 2.5 and 2.6.)

One advantage these formulas have is that they are self-consistent, i. e. as $\underset{\sim}{x}{\underset{Z}{Z}}_{t}(z \rightarrow 0)$ the 1 imit of the $f$ unction $\phi(\underset{\sim}{x})$ in the field is equal to whatever is assumed for $\phi$ on the surface, and similarly for N. For example, the 1 imit of the 1 hs of (3.9) is $\phi\left(x_{t}\right)$. The 1 imit of the rhs can be found from our regularization of the derivative of $G_{Q^{(3)}}$ from (5.20). For $j=3$ it is

$$
\frac{\partial}{\partial z} G_{R}^{(3)}\left(\underset{\sim}{x-z_{\sim}^{0}}\right)=\frac{1}{2} R_{3}\left(\underset{\sim}{x-x_{t}}\right)-\frac{1}{2} \operatorname{sgn}(z) \delta\left({\underset{\sim}{x}}_{t}-x_{i}^{\prime}\right)
$$

where from (5.20) and (5.18)

In the 1 imit of $z=0, R_{3}=0$ since it is the integral of an odd function of $k_{z}$ $\tilde{G}_{\mathbf{R}}{ }^{(3)}$ is an even function of $\mathbf{k}_{\mathrm{z}}$, and the exponential is not a function of $\mathbf{k}_{\mathbf{z}}$ 。 The result is

$$
\lim _{z \rightarrow 0} \frac{\partial}{\partial z} G_{R}^{(3)}\left(\underset{\sim}{x}-x_{t}^{\prime}\right)=-\frac{1}{2} \delta\left(x_{t}-x_{t}^{\prime}\right)
$$

which when substituted into (3.9) yields the self-consistent result that $\phi({\underset{\sim}{x}})=\phi({\underset{\sim}{t}})$. Similarly, if we differentiate (3.7) with respect to $z$ and take the 1 imit as $z \rightarrow 0$ we get the self-consistent result $N\left(X_{\sim}\right)=N\left(x_{t}\right)$.

The Rayleigh-Sommerfeld formulae can be used for other geometries where it is possible to find image-type Green's functions, i.e. geometries containing canonical shapes such as cylinders, spheres, etc. They are al so a useful starting point for geometries where the shape is nearly canonical, i.e. where the shape can be defined in a perturbation sense as canonical plus a small correction.

It can al so be shown that the Rayleigh-Sommerfeld diffraction formulae are consistent with the Kirchhoff boundary conditions on $\boldsymbol{\rho}$ or $N$ (see Sec. 6) except at the edge of an aperture. (Ref. 2.7.)

### 2.4 EXTBNDBD BODNDARY CONDITION

In (2.11) we established the integral relation for the field

$$
\begin{equation*}
\phi(x) \theta\left(z-h\left({\underset{\sim}{x}}_{t}\right)\right)=\phi^{i n}(x)+\phi^{s c}(\underset{\sim}{x}) \tag{4.1}
\end{equation*}
$$

where

$$
\begin{equation*}
\phi^{i n}(x)=\iint_{\nabla_{1}} \int_{1} G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}^{\prime}\right) S\left(x_{N}^{\prime}\right) d x_{\sim}^{\prime} \tag{4.2}
\end{equation*}
$$

and

$$
\begin{equation*}
\phi^{s c}(\underset{\sim}{x})=\iint\left[N^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{S}^{\prime}\right) \phi\left({\underset{\sim}{x}}_{s}^{\prime}\right)-G^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}^{\prime}\right) N\left(\underset{\sim}{x_{s}^{\prime}}\right)\right] d{\underset{\sim}{x}}_{t}^{\prime} \tag{4.3}
\end{equation*}
$$

The total field in $V_{1}$ is thas due to volume sources $S$, and to a 1 ayer of point and dipole sources on the surface $h$ with source densities $N\left(x_{s}^{\prime}\right)$ and $\phi\left({\underset{\sim}{x}}_{s}^{\prime}\right)$ respectively.

If we assume that $z e V_{2}$, the 1 hs of (4.1) is zero and the result is a volume boundary condition

$$
\begin{equation*}
\phi^{i n}(x)+\phi^{s c}(x)=0 \quad \underset{\sim}{x} \varepsilon V_{2} \tag{4.4}
\end{equation*}
$$

called either the extended boundary condition (Ref. 2.8), the extinction coefficient (Ref. 2.9) or the null field equation (Ref. 2.10). It is just that boundary condition on the scattered field, hence on the point and dipole sources, necessary to extinguish the incident field everywhere in $V_{2}$. Hence it is a volume and not a surface boundary condition. The induced surface fields extinguish the incident field everywhere below the surface, and directly incorporates into the solution of the problem the fact that the field in region $V_{2}$ mast vanish identically.

### 2.5 T-MATRIX

Onr field representation from (2.11) is

$$
\begin{equation*}
\phi(x)=\phi^{i n}(x)+\phi^{s c}(\underset{\sim}{x}) \quad{\underset{\sim}{x}}^{x} \varepsilon V_{1} . \tag{5.1}
\end{equation*}
$$

Me choose the Dirichlet boundary condition (2.18) so that

$$
\begin{equation*}
\phi^{s c}(\underset{\sim}{x})=-\iint G^{(3)}\left(\underset{\sim}{x}, \underset{\sim}{x}{ }_{s}^{\prime}\right) N(\underset{\sim}{x} \underset{s}{\prime}) d{\underset{\sim}{x}}_{t}^{\prime} \tag{5.2}
\end{equation*}
$$

Which follows from (2.15). Dse the Peyl representation for $G^{\left({ }^{(3)}\right.}=G_{R}^{(3)}$ from (4.11) in Ch. 1 where $z^{\prime}$ is evaluated on the surface

$$
\begin{equation*}
G_{R}^{(3)}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}^{0}\right)=\frac{\pi i}{(2 \pi)^{3}} \iint \frac{\exp \left[i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{t}-{\underset{\sim}{x}}_{t}^{\prime}\right)+i K_{+}\left|z-h\left({\underset{\sim}{x}}_{t}^{\prime}\right)\right|\right]}{\mathbb{K}_{+}} d_{\underset{\sim}{k}} \tag{5.3}
\end{equation*}
$$

where $K_{+}=\left(k_{0+}^{2}-k_{t}^{2}\right)^{1 / 2}$. Assume $z$ is greater than the highest surface excursion so that the absolnte value in (5.3) can be dropped. Substitute the result in (5.2) so that we can write

$$
\begin{equation*}
\phi^{s c}(\underline{x})=\iint \exp \left[i\left({\underset{\sim}{c}}_{t} \cdot{\underset{\sim}{x}}_{t}+\mathbb{K}_{t} z\right)\right] T\left({\underset{\sim}{t}}_{t}\right) d k_{t} \tag{5.4}
\end{equation*}
$$

where

Thas $\phi s c$ can be expressed as a sum (integral) of upgoing (propagating) or decaying (evanescent) waves in the positive z-direction. $T$ is derived from the surface sonrce density $N$ and is called the $T$-matrix. It is directly related to the scattering cross section.

In order to solve for $N$ and hence $T$ (by (5.5)) and hence the scattered fieldin $V_{1}$ (by (5.4)) we use the extended boundary condition (4.4). Then for $\underset{\sim}{x} \varepsilon V_{2},(5.2)$ becomes

$$
\begin{equation*}
p^{i n}(x)=\iint G^{(3)}\left(x, x_{s}^{\prime}\right) N\left(x_{s}^{\prime}\right) d x_{t}^{\prime} . \tag{5.6}
\end{equation*}
$$

Assume $z$ is less than the lowest surface excursion so that the absolute value in (5.3) can be dropped. The result inserted in (5.6) yields

$$
\begin{equation*}
(2 \pi)^{3}(\pi i)^{-1} K \tilde{\phi}^{i n}\left(k_{t}\right)=\iint e \operatorname{xp}\left[-i{\underset{\sim}{t}}_{t}{\underset{\sim}{x}}_{t}^{\prime}+i K_{t} h\left({\underset{\sim}{x}}_{t}^{\prime}\right)\right] N\left({\underset{\sim}{x}}_{s}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime} \tag{5.7}
\end{equation*}
$$

which is used to determine the surface density $N$ in terms of the Fourier transform of the incident field.

This method has been extended to scattering from multi-1ayered single body and to multi-bodies. It is also possible to develop a way to find T directly in terms of quantities which do not directly involve the surface fields. (Ref. 2.11.)

### 2.6 EIRCHHOFF APPROXIMATION

The Helmholtz-Kirchhoff integral representation for the field is given by (2.15) as

$$
\begin{equation*}
\phi(x)=\phi^{i n}(x)+\iint\left[N^{(3)}\left(x, x_{\sim}^{x}\right) \phi\left(x_{N}^{\prime}\right)-G^{(3)}\left(\underset{\sim}{x},{\underset{N}{s}}_{\prime}^{x}\right) N\left(x_{N}^{\prime}\right)\right] d x_{t}^{\prime} . \tag{6.1}
\end{equation*}
$$

The Kirchhoff approximation consists in assuming both the surface and normal surface derivative values of the fields, i.e. both $\phi\left(x_{s}^{\prime}\right)$ and $N\left(\underset{\sim}{x} f^{\prime}\right)$. To motivate the choice of boundary conditions we consider plane wave scattering from a flat interface. The total field is

$$
\begin{equation*}
\phi(x)=\phi^{i n}(\underset{\sim}{x})+R \phi^{s c}(\underset{\sim}{x}) \tag{6.2}
\end{equation*}
$$

where $R$ is the reflection coefficient and $\phi^{i n}$ and $\phi^{s c}$ are incident and scattered plane waves, the latter of which is specular so that

$$
\begin{equation*}
\phi^{i n}(x)=\exp \left[i k_{1}\left[\alpha_{0} x+\beta_{0} y-\gamma_{0} z\right]\right] \tag{6.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\phi^{s c}(x)=\exp \left[i \underline{k}_{1}\left[\alpha_{0} x+\beta_{0} y+\gamma_{0} z\right]\right] \tag{6.4}
\end{equation*}
$$

where $\alpha_{0}, \beta_{0}, \gamma_{0}$ are the direction cosines of the waves. $O_{n}$ the surface $z=0$

$$
\begin{equation*}
\phi\left({\underset{\sim}{x}}_{t}, 0\right)=(1+R) \phi^{i n}\left(x_{t}, 0\right) \tag{6.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial}{\partial z} \phi\left({\underset{\sim}{z}}_{t}, 0\right)=-i k_{1} \gamma_{0}(1-R) \phi^{i n}\left({\underset{\sim}{x}}_{t}, 0\right) \tag{6.6}
\end{equation*}
$$

We assume the true interface is gently undulating so that we can replace the $z$-derivative by the normal derivative. Also we evaluate the terms on the true surface, not $z=0$. We thas have the approximate boundary conditions

$$
\begin{equation*}
\phi\left({\underset{\sim}{x}}_{s}^{\prime}\right)=(1+R) \phi^{i n}\left({\underset{\sim}{x}}_{s}^{0}\right)=\phi^{0}\left({\underset{\sim}{x}}_{0}^{\prime}\right) \tag{6.7}
\end{equation*}
$$

and

$$
\begin{equation*}
N\left(\underset{\sim}{x_{s}^{\prime}}\right)=-i k_{1} \gamma_{0}(1-R) \phi^{i n}\left(x_{s}^{\prime}\right)-N^{0}\left({\underset{\sim}{s}}_{0}^{\prime}\right) \tag{6.8}
\end{equation*}
$$

Then the fi 1 d is known from (6.1).
 using (6.7) and (6.8)

$$
\begin{align*}
& \phi\left({\underset{\sim}{x}}_{s}\right)=\phi^{i n}\left({\underset{\sim}{x}}_{s}\right)-\iint G^{(3)}\left({\underset{\sim}{x}}_{s} x_{s}^{\prime}\right) N^{0}\left({\underset{\sim}{x}}_{s}^{\prime}\right) d x_{t}^{\prime} \\
& -\frac{1}{2} \iint P\left({\underset{\sim}{x}}_{s}, x_{s}^{\prime}\right) \phi^{0}\left({\underset{\sim}{x}}_{s}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime}+\frac{1}{2} \phi^{0}\left({\underset{\sim}{x}}_{s}\right) \quad . \tag{6.9}
\end{align*}
$$

so that we do not in general recover the assumed surface value i.e. $\phi\left(\mathcal{X}_{\mathrm{s}}\right)$ is not necessarily $\phi^{0}\left({\underset{\sim}{s}}^{\prime}\right)$. Differentiating (6.1) using the nomal derivative merely yields another equation (which is linearly dependent) and no recovery of surface field values which have been assumed.

If we do take the normal derivative of (6.1), i.e. multiply by $n_{m}\left(x_{t}\right) \partial_{m}$ and pass to the surface limit we get

$$
\begin{aligned}
& N\left(x_{s}\right)=N^{i n}\left(x_{s}\right)+\iint Q\left(x_{s^{\prime}}{\underset{\sim}{x}}_{s}^{\prime}\right) \phi^{0}(\underset{\sim}{x} s) d x_{t}^{\prime} \\
& +\frac{1}{2} \iint\left[\left(n_{j t}(\underset{\sim}{x}) \partial_{j t}+n_{m t}\left({\underset{\sim}{x}}_{t}\right) \partial_{m t}\right) \delta\left({\underset{\sim}{x}}_{t}-\underset{\sim}{x} t\right)\right] \phi^{0}(\underset{\sim}{x}) d{\underset{\sim}{x}}
\end{aligned}
$$

$$
\begin{aligned}
& +\frac{1}{2} N^{0}(\underset{\sim}{x}) .
\end{aligned}
$$

Integrating the delta function terms gives an additional term $N^{0}$ so that

$$
\begin{aligned}
N\left({\underset{\sim}{x}}_{s}\right)= & N^{i n}\left({\underset{\sim}{x}}_{s}\right)+\frac{3}{2} N^{0}\left({\underset{\sim}{x}}_{s}\right) \\
& +\iint\left[Q\left(\underset{\sim}{x} s^{\prime}{\underset{\sim}{x}}^{\prime}\right) \phi^{0}(\underset{\sim}{x})-\frac{1}{2} \bar{P}\left(\underset{\sim}{x} s^{\prime}{\underset{\sim}{x}}^{\prime}\right) N^{0}\left(\underset{\sim}{x}{ }_{s}^{\prime}\right)\right] d{\underset{\sim}{x}}^{\prime} .
\end{aligned}
$$

where $Q$ and $\overline{\mathbf{P}}$ are defined in (2.32) and (2.29) respectively. Again we do not in general recover the assumed boundary condition.

If however we do assume that we recover the boundary condition so that (6.9) becomes

$$
\begin{aligned}
\phi^{0}\left(x_{s}\right)=2 \phi^{i n}\left(x_{s}\right)- & 2 \iint G^{(3)}\left({\underset{\sim}{s}}^{\prime} \cdot{\underset{\sim}{x}}^{\prime}\right) N^{0}\left({\underset{\sim}{x}}_{s}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime}, \\
& -\iint P\left({\underset{\sim}{x}}_{s} \cdot{\underset{\sim}{x}}_{s}^{\prime}\right) \phi^{0}\left({\underset{\sim}{x}}_{s}^{\prime}\right) d{\underset{\sim}{x}}_{t}^{\prime},
\end{aligned}
$$

and we substitute in (6.7) and (6.8) we can solve for the reflection coefficient to get

$$
R=\frac{\rho^{i n}\left(x_{s}\right)-\iint P\left(x_{s} x_{s}^{\prime}\right) \phi_{s}^{i n}\left(x_{s}^{\prime}\right) d x_{t}^{\prime}+2 i k_{1} \gamma_{0} \iint G^{(3)}\left(x_{s} \cdot x_{s}^{\prime}\right) \phi^{i n}\left(x_{s}^{\prime}\right) d x_{t}^{\prime}}{\phi^{i n}\left(x_{s}\right)+\iint P\left(x_{s^{\prime}} x_{s}^{\prime}\right) \phi^{i n}\left(x_{s}^{\prime}\right) d x_{t}^{\prime}+2 i k_{1} \gamma_{0} \iint G^{(3)}\left({\underset{\sim}{x}}_{s} x_{s}^{\prime}\right) \phi^{i n}\left(x_{s}^{\prime}\right) d x_{t}^{\prime}},
$$

which in theory shonld be independent of $x, y$ and $h$.

## 3. RLASTICITY

In this chapter we study the propagation and scattering of waves in elastic media. To do this we derive equations satisfied by the longitudinal and transverse displacement components, discuss the free-space elastic Green's function, and use it to construct integral representations for the full Green's function (or displacement) in terms of values of displacement and traction (stress) on the surrounding surfaces, and other sources which may be present. We illustrate how to find the integral equations for these surface values in terms of a regularized kernel. We further discuss the possible boundary conditions, the plane wave states convenient for layered media problems, and the representation of the displacement in terms of potentials. We also treat the scattering problem at a plane interface for various compressional and shear wave combinations.

### 3.1 PRRLIMINARIES AND BQUATIONS

Define the orthogonal Cartesian coordinate system $x_{j}, j=1,2,3$, which is sometimes written in vector notation $\underset{\sim}{x}=\left(x_{1}, x_{2}, x_{3}\right)$. We use the symbol $u_{j}(\underset{\sim}{x}, t)$ for the three components of elastic displacement. The (symmetric) strain tensor is given by

$$
\begin{equation*}
e_{j k}=\frac{1}{2}\left(\partial_{j} u_{k}+\partial_{k} u_{j}\right) \tag{1.1}
\end{equation*}
$$

The stress tensor $\tau_{j k}$ is related to the strain by Hooke's 1 aw

$$
\begin{equation*}
\tau_{j k}=C_{j k p m} e_{p m} \tag{1.2}
\end{equation*}
$$

where repeated indices are summed over (from 1 to 3 ) and where we have defined the elastic constants $C_{j k p m}$. Since each* subscript runsfrom to 1 , there are in the most general case 81 independent elastic constants. They are really only constant in a homogeneous elastic medium, and we assume this here. For an inhomogene ous elastic medium, the elasticities are in general material functions of position. (See Appendix 3C.)

We reduce the number of independent elastic constants further by using the following symmetry restrictions derived from infinitesimal stress-strain theory:
(a) stress-strain symmetry given by

$$
\begin{equation*}
\tau_{j k}=\tau_{k j} \quad \text { and } \quad e_{p m}=e_{m p} \tag{1.3}
\end{equation*}
$$

This implies

$$
\begin{equation*}
C_{j k p m}=C_{k j p m}=C_{j k m p} \tag{1.4}
\end{equation*}
$$

If the indices are thus taken in pairs, eq. $C_{(j k)(p m)}$ then since each pair has 6 independent values $(11,22,33,12,13,23)$ there exist a total of $6 \cdot 6=36$
independent elasticities for a body without material symmetry.
(b) There is a second definition of classical linear elasticity resting on the Postulate: The work done by the stress in a deformation depends only on the strain and is recoverable work. This implies an additional symmetry

$$
\begin{equation*}
C_{j k p m}=C_{p m j k} \tag{1.5}
\end{equation*}
$$

Again thinking in terms of pairs of indices, the above imply a number of constraints given by the combinations of 6 things (independent pair values) taken 2 at a time or $\binom{6}{2}=15$ constraints. There are thas $36-15=21$ independent elasticities remaining.
(c) A final large number of constraints is introduced by isotropy. For sufficient material symmetry of the body so that the body is an isotropic elastic material, the number of independent elasticities reduces to 2. We can write the remaining non-zero tems as

$$
\begin{align*}
& C_{j \mathbf{j} \mathbf{j}}=\lambda+2 \mu  \tag{1.6}\\
& \mathbf{C}_{\mathbf{j} \mathbf{j k k}}=\lambda  \tag{1.7}\\
& \mathbf{C}_{\mathbf{j k j k}}=\mu=C_{\mathbf{j k k} \mathbf{j}}, \tag{1.8}
\end{align*}
$$

in terms of the Lame modulus $\lambda$ and the shear modulus $\mu$. The latter are al so often written in terms of Poisson's ratio $\sigma$ and Young's modulus $E$ as $\lambda=2 \mu \sigma /(1-2 \sigma)$ and $\mu=E / 2(1+\sigma)$. The resulting elastic constants can be summarized as

$$
\begin{equation*}
C_{j k p m}=\lambda \delta_{j k} \delta_{p m}+\mu\left[\delta_{j p} \delta_{k m}+\delta_{j m} \delta_{k p}\right] \tag{1.9}
\end{equation*}
$$

which is the most general fourth-rank isotropic tensor with the above
symmetries, and Hooke's law witten as

$$
\begin{equation*}
\tau_{j k}=\lambda \theta \delta_{j k}+\mu\left[\partial_{j} u_{k}+\partial_{k} \mathbf{u}_{j}\right] \tag{1.10}
\end{equation*}
$$

in terms of the dilatation $\theta=\partial_{j} u_{j}$ 。

## BQUATIONS OR HOTION

The basic equations of motion of the vector displacement are just Newton's law, $F=m a$. The force is the spatial divergence of the stress tensor. Dsing mass density $\rho$ we can write the equations of motion as

$$
\begin{equation*}
\partial_{k} \tau_{j k}=\rho \frac{\partial^{2}}{\partial t^{2}} u_{j} \tag{1.11}
\end{equation*}
$$

If we let $u_{j}(\underset{\sim}{x}, t)=\exp (-i \omega t) u_{j}(\underset{\sim}{x})$, i.e. we factor out a harmonic time dependence (so we essentially work in frequency space) we get

$$
\begin{equation*}
\partial_{k} \tau_{j k}+k^{2} \mathbf{u}_{j}=0 \quad, \quad k^{2}=\omega^{2} \rho \tag{1.12}
\end{equation*}
$$

Substituting for the stress using Hooke's 1 aw

$$
\begin{equation*}
\frac{1}{2} \partial_{k} C_{j k p m}\left[\partial_{p} u_{m}+\partial_{m} u_{p}\right]+k^{2} u_{j}=0 \tag{1.13}
\end{equation*}
$$

For an inhomogeneous medium, the $C^{\prime} s$ would be differentiated. For a homogeneous and anisotropic medium (eq. a crystal) we use the fact that the C's are constant and that $C_{j k p m}=C_{j k m p}$ to write

$$
\begin{equation*}
C_{j k p m} \partial_{k} \partial_{p} u_{m}+K^{2} u_{j}=0 \tag{1.14}
\end{equation*}
$$

which is the set of equations we work with. If we assume

$$
u_{j}(\underset{\sim}{x})=u_{j}^{0} \exp (i \underset{\sim}{k} \cdot \underset{\sim}{x})
$$

then the set of equations can be written as

$$
\left(k^{2} \delta_{j m}-c_{j k p m} k_{k} k_{p}\right) \mathbf{n}_{m}^{0}=0
$$

This is a set of three homogeneous equations of first degree for $u_{m}^{0}$. Solutions exist if

$$
\left|k^{2} \delta_{j m}-c_{j k p m} k_{k} k_{p}\right|=0
$$

i.e. if the determinant of coefficients vanishes. This is a cubic equation in $\omega^{2}$ (or $K^{2}$ ) and has three roots, $\omega_{i}(k)$. Now $\omega$ is 1 inear in $k$ so that the wave velocities (group velocities) $\partial \omega / \partial k_{j}$ are independent of $k_{j}$. Velocity of the wave is a function of its direction, not of its frequency. In general in anisotropic bodies we have three different velocities of propagation.

For an isotropic body we will find only two different velocities of propagation. Substitute (1.9) into (1.14) to get

$$
\begin{equation*}
\mu \partial_{m} \partial_{m} u_{j}+(\lambda+\mu) \partial_{j} \partial_{m} u_{m}+K^{2} u_{j}=0, \tag{1.15}
\end{equation*}
$$

or in the notation of vector analysis

$$
\begin{equation*}
\mu \Delta \underset{\sim}{u}+(\lambda+\mu) \operatorname{grad} \underset{\sim}{\nabla} \cdot \underset{\sim}{u}+\mathbb{K}_{\sim}^{\mathbf{u}}=\underset{\sim}{0} . \tag{1.16}
\end{equation*}
$$

Equivalently we could define the operator

$$
\begin{equation*}
\Delta^{*}=\mu \Delta+(\lambda+\mu) \operatorname{grad} \operatorname{div}=(\lambda+2 \mu) \operatorname{grad} \operatorname{div}-\mu \operatorname{curl} \text { curl } \tag{1.17}
\end{equation*}
$$

which plays the same role in elastic theory that the Laplacian $\Delta$ plays in harmonic function theory (e.g. for $\mu=1=-\lambda, \Delta^{*}=\Delta$ ). Our equation is thus

$$
\begin{equation*}
\left(\Delta^{*} u\right)_{j}+K_{j}^{2}=0_{j} \tag{1.18}
\end{equation*}
$$

If we decompose the displacement into longitudinal (L) and transverse (T) parts

$$
\begin{equation*}
\mathbf{u}_{\mathbf{j}}=\mathbf{u}_{\mathbf{j}}^{\mathbf{L}}+\mathbf{u}_{\mathbf{j}}^{\mathbf{T}} \tag{1.19}
\end{equation*}
$$

where the transverse part is divergenceless (solenoidal) and the longitudinal part cur11ess (irrotational)

$$
\begin{equation*}
\partial_{j} u_{j}^{T}=0 \quad ; \quad \varepsilon_{i m j} \partial_{m} u_{j}^{L}=0_{i} \tag{1.20}
\end{equation*}
$$

then we can write the longitudinal displacement as the divergence of a scalar potential $\phi$

$$
\begin{equation*}
\mathbf{u}_{\mathbf{j}}^{\mathbf{L}}=\partial_{\mathrm{j}} \phi \tag{1.21}
\end{equation*}
$$

and the transverse part as the curl of a vector potential $A_{p}$

$$
\begin{equation*}
u_{j}^{T}=\varepsilon_{j m p} \partial_{m} A_{p} \tag{1.22}
\end{equation*}
$$

We discuss these potentials later in this chapter.
Substituting these results in (1.18) using (1.17) we get

$$
\begin{equation*}
(\lambda+2 \mu) \underset{\sim}{\nabla}\left(\underset{\sim}{\nabla} \cdot{\underset{\sim}{u}}^{\mathbf{L}}\right)-\mu \underset{\sim}{\nabla} \times \underset{\sim}{\nabla} \times \underset{\sim}{u}{ }^{T}+K^{2}\left({\underset{\sim}{u}}_{\mathbf{u}}^{\mathbf{L}}+{\underset{\sim}{\underset{\sim}{u}}}^{T}\right)=\underset{\sim}{0} . \tag{1.23}
\end{equation*}
$$

If we take the curl of this equation we get an equation in only the transverse displacement

$$
\begin{equation*}
\underset{\sim}{x}\left[K_{\underset{\sim}{2}}{ }^{T}-\mu \underset{\sim}{\nabla} \underset{\sim}{x} \underset{\sim}{u} T\right]=0 \tag{1.24}
\end{equation*}
$$

The divergence of the bracket in (1.24) al so vanishes. Dp to an additional scalar potential (the Laplacian of which vanishes) we can set the bracket to zero. Osing (1.20) this is

$$
\begin{equation*}
\Delta \mathbf{u}_{j}^{T}+k_{T}^{2} \mathbf{u}_{j}^{T}=0_{j} \tag{1.25}
\end{equation*}
$$

where $k_{T}^{2}=K^{2} / \mu=\omega^{2} \rho / \mu=\omega^{2} / c_{T}^{2}$ is the square of the transverse wave number, and $c_{T}=(\mu / \rho)^{1 / 2}$ the transverse wave speed. Similarly, taking the divergence of (1.23) we wind up setting another solenoidal and irrotational bracket to zero

$$
\begin{equation*}
\mathbf{K}^{2}{\underset{\sim}{u}}_{\mathbf{L}}+(\lambda+2 \mu) \underset{\sim}{\nabla}(\underset{\sim}{\nabla} \cdot \underset{\sim}{\mathbf{u}})=\underset{\sim}{0}, \tag{1.26}
\end{equation*}
$$

and if we use the relation $\Delta=$ grad div - curlcurl we get

$$
\begin{equation*}
\Delta u_{j}^{L}+k_{L}^{2} u_{j}^{2}=0_{j} \tag{1.27}
\end{equation*}
$$

where $k_{L}^{2}=k^{2} /(\lambda+2 \mu)=\omega^{2} \rho /(\lambda+2 \mu)=\omega^{2} / c_{L}$ is the square of the longitudinal wave number, and $c_{L}=[(\lambda+2 \mu) / \rho]^{1 / 2}$ is the 1 ongitudinal wave speed.

### 3.2 FREE-SPACE ELASTIC GREEN'S FUNCTION

The free-space elastic Green's function is the tensor solution to the point source generalization of (1.18) which is

$$
\begin{equation*}
\left[\Delta^{*} G^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)\right]_{i j}+K^{2} G_{i j}^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=-\delta_{i j} \delta\left(\underset{\sim}{x}-x^{\prime}\right) \tag{2.1}
\end{equation*}
$$

and is explicity given by (see Appendix 3B)

$$
\begin{equation*}
G_{i j}^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\frac{1}{\mu} \delta_{i j} G^{T}\left(\underset{\sim}{x}, x^{\prime}\right)+\frac{1}{K^{2}} \partial_{i} \partial_{j}\left[G^{T}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)-G^{L}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)\right] \tag{2.2}
\end{equation*}
$$

Where $G^{T}$ and $G^{L}$ are the scalar free space Green's fanction with wave numbers $k_{T}$ and $k_{L}$ respectively. That is they are

$$
\begin{equation*}
G^{T, L}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\frac{\exp \left(i k_{T, L}\left|x_{\sim}^{x-x^{\prime}}\right|\right)}{4 \pi\left|\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right|} \tag{2.3}
\end{equation*}
$$

We choose the solution satisfying the outgoing radiation condition, i.e. the retarded solution. Note that $G_{i j}^{0}$ is regular in the sense of the singularities we discussed in Ch. 1. $G^{T}$ is regular, and the second derivative of $G^{T}$ contains the same singular terms as the second derivative of $G^{L}$. i.e. the singularities are independent of wavenumber and thus cancel. The derivative of (2.2) will occur in later integral equations we derive, and this is regularized in App. 3A.

To prove that $G^{0}$ is a solution of (2.1) substitute it and use the definition of $\Delta^{*}$ in (1.17) to get

$$
\begin{aligned}
{\left[\Delta^{*} G(\underset{\sim}{x}, \underset{\sim}{x})\right]_{i j}=} & \mu \partial_{m} \partial_{m} G_{i j}^{0}+(\lambda+\mu) \partial_{i} \partial_{m} G_{m j}^{0} \\
= & \mu \partial_{m} \partial_{m}\left[\mu^{-1} \delta_{i j} G^{T}+\mathbb{K}^{-2} \partial_{i} \partial_{j}\left(G^{T}-G^{L}\right)\right] \\
& +(\lambda+\mu) \partial_{i}\left[\mu^{-1} \partial_{j} G^{T}+\mathbb{K}^{-2} \partial_{j} \partial_{m} \partial_{m}\left(G^{T}-G^{L}\right)\right] .
\end{aligned}
$$

Ose the differential equations satisfied by $G^{T}$ and $\mathbf{G}^{\text {L }}$, viz.

$$
\begin{equation*}
\left[\partial_{m} \partial_{m}+k_{T, L}^{2}\right]^{G^{T}, L_{(x, ~}^{x}}(\underset{\sim}{\prime})=-\delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right) \tag{2.4}
\end{equation*}
$$

and the wave number values $k_{T}^{2}=K^{2} / \mu$ and $k_{L}^{2}=K^{2} /(\lambda+2 \mu)$ to get

$$
\left[\Delta^{*} G^{0}\left(\underset{\sim}{x}{\underset{\sim}{x}}^{\prime}\right)\right]_{i j}=-\partial_{i} \partial_{j}\left(G^{T}-G^{L}\right)-\delta_{i j} k_{T}^{2} G^{T}-\delta_{i j} \delta\left(\underset{\sim}{x}-x^{\prime}\right),
$$

which can be written as (2.1), i.e.

$$
\left[\Delta^{*} G^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)\right]_{i \mathbf{j}}=-K^{2} G_{i j}^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)-\delta_{i j} \delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right) .
$$

He can al so derive the divergence and curl of (2.2) as

$$
\begin{equation*}
\partial_{m} G_{m p}^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=(\lambda+2 \mu)^{-1} \partial_{p} G^{L}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right), \tag{2.5}
\end{equation*}
$$

and

$$
\begin{equation*}
e_{i j m}{ }_{j} G_{m p}^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)=\mu^{-1} e_{i j p} \partial_{j} G^{T}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right) \tag{2.6}
\end{equation*}
$$

Note that j ust as in the scalar case $\mathrm{G}_{\mathrm{ij}}^{0}\left(\underset{\sim}{x}, \mathrm{x}^{\prime}\right)$ is a homogeneous function of its arguments and can be written as $G_{i j}^{0}\left(\underline{x}-x^{\prime}\right)$ 。

### 3.3 SURFACE INTEGRAL EQUATIONS

We derive the surface integral equation satisfied by the full elastic tensor Green's function $G_{i j}$ which satisfies the same differential equation (2.1) as $G_{i j}^{0}$ and in addition certain boundary conditions which we specify later. The development is analogous to that in the scalar case. $G_{i j}$ satisfies the equation

$$
\begin{equation*}
\left[\Delta^{*} G\left(\underset{\sim}{x}, \underline{x}^{n}\right)\right]_{i n}+K^{2} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)=-\delta_{i n} \delta\left({\underset{\sim}{x}}^{-x^{\prime \prime}}\right) \tag{3.1}
\end{equation*}
$$

and we write the equation (2.1) on $G_{i j}^{0}$ but here differentiate on the source variable

$$
\begin{equation*}
\left[\Delta^{*} G^{0}(\underset{\sim}{x} \cdot, \underset{\sim}{x})\right]_{i j}+K^{2} G_{i j}^{0}\left(x^{\prime} \cdot \underset{\sim}{x}\right)=-\delta_{i j} \delta\left(\underset{\sim}{x}{ }^{\prime}-\underset{\sim}{x}\right) \tag{3.2}
\end{equation*}
$$

By cross multiplication we form the quantity

$$
\begin{equation*}
G_{i j}^{0}(\underset{\sim}{x}, \underline{\sim})\left[\Delta^{*} G\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)\right]_{i n}-\left[\Delta^{*} G^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)\right]_{i j} G_{i n}\left(\underset{\sim}{x} \cdot{\underset{\sim}{x}}^{n}\right) \tag{3.3}
\end{equation*}
$$

Nert write (3.3) in two ways which we then equate. The first way is

$$
\begin{align*}
(a)= & G_{i j}^{0}(\underset{\sim}{x}, \underset{\sim}{x})\left[-K^{2} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)-\delta_{i n} \delta\left({\underset{\sim}{x}}^{-x^{n}}\right)\right] \\
& -\left[-K^{2} G_{i j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)-\delta_{i j} \delta\left({\underset{\sim}{x}}^{\prime}-\underset{\sim}{x}\right)\right] G_{i n}\left(\underset{\sim}{x}, x^{n}\right) \\
= & G_{j n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right) \delta\left(\underset{\sim}{x^{\prime}}-\underset{\sim}{x}\right)-G_{j n}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right) \delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{n}\right) \tag{3.4}
\end{align*}
$$

where we have used the symmetry of $G^{0}$ as

$$
\begin{equation*}
\mathbf{G}_{\mathbf{n j}}^{0}=\mathbf{G}_{\mathrm{jn}}^{0}, \tag{3.5}
\end{equation*}
$$

to write the last term in (3.4). The second way is to use the specific definition of the differential operator given by (1.17) to wite

$$
\begin{align*}
(b)= & G_{i j}^{0}(\underset{\sim}{x}, \underline{x})\left[\mu \partial_{m} \partial_{m} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)+(\lambda+\mu) \partial_{i} \partial_{m} G_{m n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)\right] \\
& -\left[\mu \partial_{m} \partial_{m} G_{i j}^{0}(\underset{\sim}{x}, \underset{\sim}{x})+(\lambda+\mu) \partial_{i} \partial_{m} G_{m j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)\right] G_{i n}\left(x, x^{n}\right) \tag{3.6}
\end{align*}
$$

Next factor a divergence term out of this as

$$
\begin{aligned}
& (b)=G_{i j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right) \partial_{p}\left[\mu \delta_{p m} \partial_{m} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)+\mu \delta_{p m} \partial_{i} G_{m n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)\right. \\
& \left.+\lambda \delta_{i p} \partial_{m} G_{m n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)\right]
\end{aligned}
$$

$$
\begin{align*}
& \left.+\lambda \delta_{i p}{ }^{\partial} \mathrm{G}_{\mathrm{mj}}{ }^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)\right] \mathrm{G}_{\mathrm{in}}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right) \quad . \tag{3.7}
\end{align*}
$$

and finally note that we can factor the divergence out of the full term

$$
\begin{align*}
& \left.+\lambda \delta_{i p} \partial_{m} G_{m n}\left(x, x^{\prime \prime}\right)\right] \\
& -[\mu \delta_{p m} \partial_{m} G_{i j}^{0}(\underset{\sim}{x}, \underbrace{x}_{\sim})+\mu \delta_{p m}{ }_{i} G_{m j}^{0}\left(\underset{\sim}{x}{ }^{\prime}, \underset{\sim}{x}\right) \\
& \left.\left.\left.+\lambda \delta_{i p}{ }^{\partial_{m}} G_{m j}^{0}\left(x^{\prime}, x\right)\right] G_{i n}\left(x, x^{\prime \prime}\right)\right]\right] \\
& -\left[\partial_{p} G_{i j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)\right]\left[\mu \delta_{p m} \partial_{m} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)+\mu \delta_{p m} \partial_{i} G_{m n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)\right. \\
& \left.+\lambda \delta_{i p}{ }^{\partial_{m}} G_{m n}\left(\underset{\sim}{x}, x^{\prime \prime}\right)\right] \\
& +\left[\mu \delta_{p m} \partial_{m} G_{i j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)+\mu \delta_{p m}{ }^{\partial} \mathbf{G}_{m j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)\right. \\
& \left.+\lambda \delta_{i p} \partial_{m} G_{m j}^{0}\left({\underset{\sim}{x}}^{\prime}, \underset{\sim}{x}\right)\right] \partial_{p} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime \prime}\right) \quad . \tag{3.8}
\end{align*}
$$

It can easily be seen that the last six tems all cancel so that (b) is the divergence of a triple index object. We write it as

$$
\begin{equation*}
(b)=\partial_{p}\left[\underline{z}^{\prime} \cdot \underset{\sim}{x} \cdot{\underset{\sim}{x}}^{n}\right]_{\mathrm{pjn}} \tag{3.9}
\end{equation*}
$$

where we define the symbol
with the operator $T$ defined as

$$
\begin{equation*}
\left[T G\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)\right]_{p i n}=\mu \partial_{p} G_{i n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right)+\mu \partial_{i} G_{p n}\left(\underset{\sim}{x}, x^{n}\right)+\lambda \delta_{i p} \partial_{m} G_{m n}\left(x, x^{n}\right), \tag{3.11}
\end{equation*}
$$

which will be related to the traction of $G$ across the surface. The definition (3.11) follows from the first set of terms in (3.8). Analogously, for the free space Green's function we have

$$
\begin{equation*}
\left[T G^{0}(\underset{\sim}{x}, x)\right]_{p i j}=\mu \partial_{p} G_{i j}^{0}\left(x_{\sim}^{\prime}, \underset{\sim}{x}\right)+\mu \partial_{i} G_{p j}^{0}\left(x^{\prime}, x\right)+\lambda \delta_{i p} \partial_{m} G_{m j}^{0}(\underset{\sim}{x}, \underline{x}) \tag{3.12}
\end{equation*}
$$

Note that the differential operators in (3.11) act on the "field" coordinate $\underset{\sim}{x}$, and in (3.12) act on the "source" coordinate $X$ of the function. Al so note that (3.11) and (3.12) are symmetric in $p$ and i. We could al so use (2.5) to simplify the 1 ast term in (3.12). Equating (3.4) and (3.9) we get finally

$$
\begin{equation*}
G_{j n}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right) \delta\left({\underset{\sim}{x}}^{\prime}-\underset{\sim}{x}\right)=G_{j n}^{0}\left(x_{\sim}^{\prime}, \underset{\sim}{x}\right) \delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{n}\right)+\partial_{p}\left[{\underset{\sim}{x}}^{\prime}, \underset{\sim}{x},{\underset{\sim}{x}}^{n}\right]_{p j n} . \tag{3.13}
\end{equation*}
$$

From this basic identity we are able to start forming integral equations for many different problems.

## Eg. 1. TRACTION FREE SURFACE

We specify the surface $z=h(\underset{\sim}{x})$ and we want to find the field above this surface. Multiply (3.13) by the step function $\theta\left(z-h\left(X_{t}\right)\right)$ and integrate over all space $\iiint_{\text {dx }}$. The result is

$$
\begin{align*}
& G_{j n}\left(x_{\sim}^{\prime} \cdot{\underset{\sim}{x}}^{\prime \prime}\right) \theta\left(z^{\prime}-h(\underset{\sim}{x}\right. \\
&=) \\
&= G_{j n}^{0}\left(x^{\prime} \cdot{\underset{\sim}{n}}^{\prime \prime}\right) \theta\left(z^{\prime \prime}-h\left({\underset{\sim}{x}}_{t}^{\prime \prime}\right)\right)  \tag{3.14}\\
&+\iiint \partial_{p}\left[{\underset{\sim}{x}}^{\prime}, \underset{\sim}{x},{\underset{\sim}{x}}^{\prime \prime}\right]_{p j n} \theta\left(z-h\left({\underset{\sim}{x}}_{t}\right)\right) d x .
\end{align*}
$$

The field term on the 1 hs of (3.14) thus exists provided $z^{\prime}>h(x t$ ), i.e. assuming the vector field position $\underset{\sim}{x}$ is above the surface. The source term exists provided $z^{\prime \prime}>h\left({\underset{\sim}{x}}^{\prime \prime}\right)$, i.e. provided the source point ${\underset{\sim}{n}}^{\prime \prime}$ is above the surface. For the integral term we integrate by parts. Surface terms at infinity vanish either because of the step function or because of the radiation condition, and if we use the fact that

$$
\begin{equation*}
\partial_{p} \theta(z-h(\underset{\sim}{x}))=\delta\left(z-h\left({\underset{\sim}{x}}_{t}\right)\right) n_{p}(\underset{\sim}{x}), \tag{3.15}
\end{equation*}
$$

where

$$
\begin{equation*}
n_{p}\left({\underset{\sim}{x}}_{t}\right)=\delta_{p 3}-\partial_{p t} h\left({\underset{\sim}{x}}_{t}\right), \tag{3.16}
\end{equation*}
$$

is a vector in the direction of the surface normal (not a unit vector) we get that for ${\underset{\sim}{x}}^{\prime}$ and ${\underset{\sim}{x}}^{\prime \prime}$ above the surface

$$
G_{j n}\left({\underset{\sim}{x}}^{\prime} \cdot{\underset{\sim}{x}}^{n}\right)=G_{j n}^{0}\left({\underset{\sim}{x}}^{n} \cdot{\underset{\sim}{x}}^{n}\right)-\iint n_{p}\left({\underset{\sim}{x}}_{t}\right)\left[{\underset{\sim}{x}}^{\prime} \cdot{\underset{\sim}{x}}^{s^{\prime}}{\underset{\sim}{n}}^{n}\right]_{p j n}{ }^{d x}{\underset{\sim}{x}},
$$

where we have evaluated the delta function and $\underset{\sim}{x} s=(\underset{\sim}{x} t h(\underset{\sim}{x} t)$ ) is a threevector evaluated on the surface. Explicitly the symbol term in (3.17) is from (3.10)

$$
\begin{align*}
& n_{p}\left({\underset{\sim}{x}}_{t}\right)\left[{\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s},{\underset{\sim}{x}}^{\prime \prime}\right] \operatorname{pjn} \\
& =G_{i j}^{0}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s}\right) n_{p}\left({\underset{\sim}{t}}_{t}\right)\left[T G\left({\underset{\sim}{s}}_{s},{\underset{\sim}{x}}^{n}\right)\right] \text { pin } \\
& -n_{p}\left(x_{t}\right)\left[T G^{0}\left({\underset{\sim}{x}}^{\prime} \cdot{\underset{\sim}{x}}_{s}\right)\right]{ }_{p i j} G_{i n}\left({\underset{\sim}{x}}_{s} \cdot{\underset{\sim}{x}}^{m}\right) \tag{3.18}
\end{align*}
$$

If we assume a zero traction (or traction-free or just free) boundary condition given by

$$
\begin{equation*}
n_{p}\left({\underset{\sim}{x}}_{t}\right)\left[T G\left({\underset{\sim}{s}}_{s},{\underset{\sim}{x}}^{n}\right)\right]_{p i n}=0_{i n} \tag{3.19}
\end{equation*}
$$

then (3.17) becomes

$$
\begin{align*}
G_{j n}\left(x^{\prime},{\underset{\sim}{x}}^{n}\right)= & G_{j n}^{0}\left(x_{\sim}^{\prime},{\underset{\sim}{x}}^{n}\right) \\
& +\iint n_{p}\left({\underset{\sim}{x}}_{t}\right)\left[T G^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}\right)\right]_{p i j} G_{i n}\left({\underset{\sim}{x}}_{s} \cdot{\underset{\sim}{x}}^{n}\right) d \underset{\sim}{x} \tag{3.20}
\end{align*}
$$

which is our first result, the integral representation of the Green's function at the field point $x^{\prime}$ 'above the surface in terms of its "value" on the surface $G_{i n}\left({\underset{\sim}{x}}^{s}, \mathbb{x}^{\prime \prime}\right)$ which however, is unnown. To find an integral equation for the surface value take the surface limit, i.e. let ${\underset{\sim}{x}}^{\prime} \rightarrow x_{\sim}^{\prime} s^{\prime}$ a point on the surface. As in the scalar cases we treated, we must regularize the kernel of the transform in (3.20). We do this in Appendix 3A, where from Eq. (A.31) we find that we can write

$$
\begin{align*}
n_{p}\left({\underset{x}{t}}^{t}\right) & {\left[T G^{0}\left(\underset{\sim}{x} \cdot,{\underset{\sim}{x}}_{s}\right)\right]_{p i j} } \\
= & K_{j i}\left({\underset{\sim}{x}}^{\prime} \cdot{\underset{\sim}{x}}_{s}\right) \\
& +1 / 2 \operatorname{sgn}\left(z^{\prime}-h\left({\underset{\sim}{x}}_{t}^{\prime}\right)\right) \delta\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right) \\
& \cdot\left[\delta_{i j}-\delta_{i 3} \delta_{j t} h\left({\underset{\sim}{x}}_{t}\right)-\Lambda \delta_{j 3}{ }_{i}{ }_{i t} h\left({\underset{\sim}{x}}_{t}\right)\right], \tag{3.21}
\end{align*}
$$

Where $K_{j i}$ is the regular part and is given explicitly by (A.32) and $\Lambda=\lambda /(\lambda+2 \mu)$. Substituting this result in (3.20) and taking the surface limit we get
where the matrix $Q_{j i}$ is definedas

$$
Q_{j i}\left({\underset{\sim}{x}}^{\prime}\right)=1 / 2\left[\delta_{i j}+\delta_{i 3}{ }_{j}{ }_{j}^{\prime} t{ }^{h}\left(\underset{\sim}{x}{ }_{t}^{\prime}\right)+\Lambda \delta_{j 3}{ }_{i}{ }_{i t} h\left({\underset{\sim}{x}}_{t}\right)\right] .
$$

Equation (3.22) is the integral equation for the value of the Green's function on the surface, $G_{i n}\left(\underset{\sim}{x}, x^{\prime \prime}\right)$. The procedure is to solve the se coupled integral equations (in general computationally, and in general difficult) for the surface values and then substitute them into the integral representation (3.20). An alternative integral equation can be formed if we define the surface Green's function as the lhs of (3.22)

$$
\begin{equation*}
G_{j n}^{s}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}^{n}\right)=Q_{j i}\left({\underset{\sim}{x}}_{t}\right) G_{i n}\left({\underset{\sim}{x}}_{s}^{\prime},{\underset{\sim}{x}}^{n}\right) . \tag{3.24}
\end{equation*}
$$

Osing the inverse matrix to $Q$ given by

$$
\begin{equation*}
Q_{j m}\left({\underset{\sim}{t}}_{t}^{\prime}\right) 0_{m p}\left({\underset{\sim}{x}}_{t}^{\prime}\right)=\delta_{j p} \tag{3.25}
\end{equation*}
$$

we can write

$$
\begin{equation*}
G_{m n}\left({\underset{\sim}{x}}_{s}^{\prime},{\underset{\sim}{x}}^{n}\right)=0_{m p}\left({\underset{\sim}{x}}_{t}^{\prime}\right) G_{p n}^{s}\left({\underset{\sim}{x}}_{s^{\prime}},{\underset{\sim}{x}}^{n}\right), \tag{3.26}
\end{equation*}
$$

so that (3.22) becomes

$$
\begin{equation*}
G_{j n}^{s}\left({\underset{\sim}{x}}_{s}^{\prime} \cdot x^{n}\right)=G_{j n}^{0}\left(x_{s}^{\prime}, x_{\sim}^{n}\right)+\iint X_{j i}\left({\underset{\sim}{x}}_{s}^{\prime}{\underset{\sim}{x}}_{s}\right) 0_{i p}\left({\underset{\sim}{x}}_{t}\right) G_{p n}^{s}\left({\underset{\sim}{x}}_{s} \cdot{\underset{\sim}{x}}^{n}\right) d x_{t} \tag{3.27}
\end{equation*}
$$

Explicitly we have, in matrix notion

$$
Q\left({\underset{\sim}{x}}_{t}\right)=\frac{1}{2}\left[\begin{array}{lll}
1 & 0 & \partial_{1} h(\underset{\sim}{x} t  \tag{3.28}\\
0 & 1 & \partial_{2} h\left({\underset{\sim}{x}}_{t}\right) \\
\lambda_{1} h\left({\underset{\sim}{x}}_{t}\right) & \Lambda_{2} h(\underset{\sim}{x} t
\end{array}\right]
$$

and

$$
D\left({\underset{\sim}{x}}_{t}\right)=\frac{2}{1-\Lambda\left[\partial_{t} h\left({\underset{\sim}{x}}_{t}\right)\right]^{2}}\left[\begin{array}{lll}
1-\Lambda\left[\partial_{2} h\left({\underset{\sim}{x}}_{t}\right)\right]^{2} & \Lambda \partial_{1}\left({\underset{\sim}{x}}_{t}\right) \partial_{2} h\left({\underset{\sim}{x}}_{t}\right) & -\partial \partial_{1} h(\underset{\sim}{x} t \\
\Lambda \partial_{1} h\left({\underset{\sim}{x}}_{t}\right) \partial_{2} h\left({\underset{\sim}{x}}_{t}\right) & 1-\Lambda\left[\partial_{1} h(\underset{\sim}{x} t)\right]^{2} & -\partial_{2} h\left({\underset{\sim}{x}}_{t}\right) \\
-\Lambda \partial_{1} h\left({\underset{\sim}{x}}_{t}\right) & -\Lambda \partial_{2} h\left({\underset{\sim}{x}}_{t}\right) & 1
\end{array}\right]
$$

and the field representation (3.20) can be written as

$$
\begin{aligned}
G_{j n}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}^{n}\right)= & G_{j n}^{0}\left(x^{\prime},{\underset{\sim}{x}}^{n}\right) \\
& +\iint_{m}\left({\underset{\sim}{x}}_{t}\right)\left[T G^{0}\left(x_{\sim}^{\prime},{\underset{\sim}{x}}_{s}\right)\right]_{m i j} v_{i p}\left({\underset{\sim}{x}}_{t}\right) G_{p n}^{s}\left({\underset{\sim}{x}}^{\prime}{\underset{\sim}{x}}^{\prime \prime}\right) d{\underset{\sim}{x}}_{t} \cdot(3,30)
\end{aligned}
$$

## Eg. 2. VOLOMR SOURCBS - TRACTION FREE SORFACE

The source terms in (3.20), (3.22), (3.27) or (3.30) are all due to the free space Green's function $\mathbf{G}^{0}$. If instead we want to solve for an arbitrary tensor field $u_{i n}(\underset{\sim}{x})$ due to the tensor sources $S_{i n}(\underset{\sim}{x})$, i.e. when we have the differential equation

$$
\begin{equation*}
\left[\Delta^{*} u(\underset{\sim}{x})\right]_{i n}+K^{2} u_{i n}(x)=-S_{i n}(x) \tag{3.31}
\end{equation*}
$$

we multiply (3.20), (3.22), (3.27), or (3.30) from the right by $S_{n q}\left(x_{\sim}^{\prime \prime}\right)$ and integrate over ${\underset{\sim}{x}}^{n}$. From (3.27) for example we have that

$$
\begin{equation*}
u_{j n}^{s}\left({\underset{\sim}{x}}_{s}^{\prime}\right)=u_{j n}^{i n}\left({\underset{\sim}{x}}_{s}^{\prime}\right)+\iint X_{j i}\left({\underset{\sim}{x}}_{\prime}^{\prime}, x_{s}\right) 0_{i p}(\underset{\sim}{x}) u_{p n}^{s}(\underset{\sim}{x}) d{\underset{\sim}{x}}_{t}, \tag{3.32}
\end{equation*}
$$

where

$$
\begin{equation*}
u_{j n}^{s}\left(x_{\sim}^{\prime}\right)=\iiint G_{j m}^{s}\left(x_{s}^{\prime}, x^{n}\right) S_{m n}\left({\underset{\sim}{x}}^{n}\right) d x^{\prime \prime}, \tag{3.33}
\end{equation*}
$$

and the incident field is

$$
\begin{equation*}
u_{j n}^{i n}\left(x_{s}^{\prime}\right)=\iiint G_{j m}^{0}\left({\underset{\sim}{x}}_{s}^{\prime},{\underset{\sim}{x}}^{n}\right) S_{m n}\left({\underset{\sim}{x}}^{n}\right) d{\underset{\sim}{x}}^{\prime \prime} \tag{3.34}
\end{equation*}
$$

The function evaluated at a point in the field, $\underset{\sim}{x}$, i.e. off the surface, is defined via

$$
\begin{equation*}
u_{j n}\left(x^{\prime}\right)=\iiint G_{j m}\left(x^{\prime} \cdot{\underset{\sim}{x}}^{n}\right) S_{m n}\left(x^{n}\right) d{\underset{\sim}{x}}^{n} \tag{3.35}
\end{equation*}
$$

For a vector equation, for example, for the displacement $u_{i}$ we have
that including vector source terms

$$
\begin{equation*}
\left(\Delta^{*} u(\underset{\sim}{x})\right)_{i}+K^{2} u_{i}(\underset{\sim}{x})=-S_{i}(\underset{\sim}{x}) \tag{3.36}
\end{equation*}
$$

the field representation can be found from (3.20) by multiplying from the right by $S_{n}\left({\underset{\sim}{x}}^{\prime \prime}\right)$ and integrating over ${\underset{\sim}{n}}^{n}$. Osing the definition

$$
\begin{equation*}
u_{j}\left(\underline{x}^{\prime}\right)=\iiint G_{j n}\left(x^{\prime} \cdot \underline{x}^{n}\right) S_{n}\left(x^{n}\right) d x^{n}, \tag{3.37}
\end{equation*}
$$

we can wite this field displacement value in terms of its surface valne as

$$
\begin{equation*}
\left.u_{j}\left({\underset{\sim}{x}}^{\prime}\right)=u_{j}^{i n}\left({\underset{\sim}{x}}^{\prime}\right)+\iint_{p}\left({\underset{\sim}{x}}_{t}\right)\left[T G^{0}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s}\right)\right]\right]_{p i j} u_{i}\left({\underset{\sim}{x}}_{s}\right) d{\underset{\sim}{x}}_{t}, \tag{3.38}
\end{equation*}
$$

where the incident field is analogous to (3.34) if we simply drop the n-index. The surface integral equation can be found from (3.22) for example by the same procedure. It is

$$
\begin{equation*}
Q_{j i}\left({\underset{\sim}{x}}_{t}^{\prime}\right) u_{i}\left({\underset{\sim}{x}}_{s}^{\prime}\right)=u_{j}^{i n}\left(x_{s}^{\prime}\right)+\iint K_{j i}\left({\underset{\sim}{x}}_{s}^{\prime} \cdot{\underset{\sim}{x}}_{s}\right) u_{i}\left({\underset{\sim}{x}}_{s}\right) d{\underset{\sim}{x}}_{t} \tag{3.39}
\end{equation*}
$$

## Eg. 3. THO ELASTIC MEDIA

The traction free boundary in examples 1 and 2 was essentially a perfectly reflecting boundary condition. If instead we have two different elastic media joined at our rough interface $z=h\left(\underline{x}_{t}\right)$ as in Fig. 3.1


Fig. 3.1
we can still use our identity (3.13) but now must apply it twice. First,
multiply by $\theta\left(z-h\left({\underset{\sim}{x}}_{t}\right)\right)$ and integrate over all space $\underset{\sim}{x}$. We get essentially the results of Eg. 1 except that the traction is not zero and all the fields and parameters have superscript (1). We assume, of course, that the source is in region (1). The field representation in region (1) is thas for the displacement vector in analogy to (3.38)

$$
\begin{align*}
u_{j}^{(1)}\left(x^{\prime}\right)= & u_{j}^{i n}\left(x_{\sim}^{\prime}\right)+\iint n_{p}\left({\underset{\sim}{x}}_{t}\right)\left[T G^{0}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s}\right)\right]_{p i j}^{(1)} u_{i}^{(1)}\left({\underset{\sim}{x}}_{s}\right) d x_{t} \\
& -\iint_{j i}^{0(1)}\left(x^{\prime} \cdot{\underset{\sim}{x}}_{s}\right) T_{i}^{(1)}\left({\underset{\sim}{x}}_{s}\right) d_{\sim}^{x} x_{t} \tag{3.40}
\end{align*}
$$

where we have abbreviated the traction as

$$
\begin{equation*}
T_{i}^{(1)}\left(X_{s}\right)=n_{p}\left(x_{t}\right)\left[T u\left(x_{s}\right)\right]_{p i}^{(1)} \tag{3.41}
\end{equation*}
$$

The field in region (1) is thas given once its surface values and the surface values of its traction are known. An integral equation for the surface values can be found by letting $\underset{\sim}{x^{\prime}} \rightarrow{\underset{\sim}{x}}_{\prime}^{\prime}{ }_{s}$ from above the surface. This is exactly the result we had before with the addition of the surface traction term. No regularization of $G^{0(1)}$ is required since it is not singular. The result is

$$
\begin{align*}
& Q_{j i}^{(1)}\left({\underset{\sim}{x}}_{t}^{\prime}\right) u_{i}^{(1)}\left(\underset{\sim}{x_{s}^{\prime}}\right)=u_{j}^{i n}\left(\underset{\sim}{x}{ }_{s}^{\prime}\right)+\iint X_{j i}^{(1)}\left({\underset{\sim}{x}}_{s}^{\prime},{\underset{\sim}{x}}_{s}\right) u_{i}^{(1)}(\underset{\sim}{x}) d{\underset{\sim}{x}}_{t} \\
& -\iint G_{j i}^{0(1)}\left(\underset{\sim}{x}{ }_{s}^{\prime},{\underset{\sim}{x}}_{s}\right) T_{i}^{(1)}\left(\underset{\sim}{x}{ }_{s}\right) d{\underset{\sim}{x}} \text {, } \tag{3.42}
\end{align*}
$$

where the $Q$-matrix follows from (3.23) if we replace the elastic parameters by $\lambda^{(1)}$ and $\mu^{(1)}$.

In region (2) we multiply (3.13) by $\theta(h(\underset{\sim}{x})-z)$ and integrate over all
space. There is no source term now since we assumed it to be in region (1) and the field representation in region (2) is thus

$$
\begin{align*}
& u_{j}^{(2)}\left(\underset{\sim}{x^{\prime}}\right)=-\iint_{p}(\underset{\sim}{x})\left[T G^{0}\left(\underset{\sim}{x}{ }^{\prime},{\underset{\sim}{x}}_{s}\right)\right]_{p i j}^{(2)}{ }_{i}^{(2)}(\underset{\sim}{x}) d{\underset{\sim}{x}}_{t} \\
& +\iint G_{j i}^{o(2)}\left(\underset{\sim}{x}{ }^{\prime} \cdot \underset{\sim}{x}\right) T_{i}^{(2)}\left(\underset{\sim}{x}{ }_{s}\right) d{\underset{\sim}{x}}_{t} \quad . \tag{3.43}
\end{align*}
$$

in terms of surface displacements and tractions which result as we take the limit from region (2). The additional minus sign on the rhs of (3.43) results from the fact that when we integrate the divergence term by parts and differentiate the step function, it has the negative of the argument it had in region (1). The surface integral equation follows analogously by taking the limit $\underset{\sim}{x^{\prime}} \rightarrow \underset{\sim}{x}{ }_{s}^{\prime}$ as $z^{\prime} \rightarrow h\left(\underset{\sim}{x} t^{\prime}\right)$ from below. We thus get an additional minus sign from the singular terms since the signum function is ne gative. The resulting surface integral equation is

$$
\begin{align*}
& +\iint_{j i}^{0(2)}\left({\underset{\sim}{x}}^{\prime} \cdot{\underset{\sim}{x}}_{s}\right) T_{i}^{(2)}\left({\underset{\sim}{x}}_{s}\right) d{\underset{\sim}{x}}_{t} \quad, \tag{3.44}
\end{align*}
$$

where the Q-matrix here follows from (3.23) by replacing the elastic paramters by $\lambda^{(2)}$ and $\mu^{(2)}$.

We thus have two vector integral equation (3.42) and (3.44) with four unknown vector quantities on the surface. We require two additional vector boundary conditions, actually continuity conditions, at the boundary interface. They are the continuity of vector displacement and traction (stress) given by

$$
\begin{align*}
& u_{m}^{(1)}(\underset{\sim}{x})=u_{m}^{(2)}\left({\underset{\sim}{x}}_{s}\right)  \tag{3.45}\\
& T_{i}^{(1)}\left({\underset{\sim}{x}}_{s}\right)=T_{i}^{(2)}\left({\underset{\sim}{x}}_{s}\right) \tag{3.46}
\end{align*}
$$

The resulting equations (3.42) and (3.44) are thus two coupled equations for surface displacements and tractions. These are solved and the results used to find the fields in the upper region using (3.40) and in the 1 ower region using (3.43).

## Eg. 4. BLASTIC LAYER

For an elastic layer sandwiched between two rough surfaces $h_{1}\left(x_{t}\right)$ and $h_{2}\left(x_{t}\right)$ as in Fig. 3.2

z

$$
h_{2}({\underset{\sim}{x}} t)
$$

Fig. 3.2
we can again use the identity (3.13). Now however, we multiply it with the product of two step functions

$$
\begin{equation*}
\theta\left(h_{1}\left({\underset{\sim}{x}}_{t}\right)-z\right) \theta\left(z-h_{2}(\underset{\sim}{x})\right) \tag{3.47}
\end{equation*}
$$

and integrate on $\underset{\sim}{x}$ over all space. The result is for $\underset{\sim}{x}$ and $\underset{\sim}{x}$ in the 1 ayer

$$
\begin{align*}
G_{j n}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}^{n}\right)= & G_{j n}^{0}\left(x_{\sim}^{\prime} \cdot{\underset{\sim}{x}}^{n}\right) \\
& +\iiint_{p} \partial_{p}\left[{\underset{\sim}{x}}^{\prime} \cdot \underset{\sim}{x},{\underset{\sim}{x}}^{\prime \prime}\right]_{p j n} \theta\left(h_{1}(\underset{\sim}{x})-z\right) \theta\left(z-h_{2}(\underset{\sim}{x})\right){\underset{\sim}{x}}^{t} \tag{3.48}
\end{align*}
$$

which is written in terms of the symbol defined in (3.10). Integration by parts of this integral produces the derivative of the product of step
functions which is given by

$$
\begin{align*}
& \partial_{p} \theta\left(h_{1}\left({\underset{\sim}{x}}_{t}\right)-z\right) \theta\left(z-h_{2}(\underset{\sim}{x})\right) \\
& =-n_{p}^{(1)}(\underset{\sim}{x}) \delta\left(x-h_{1}(\underset{\sim}{x})\right) \theta\left(z-h_{2}(\underset{\sim}{x})\right) \\
& +n_{p}^{(2)}(\underset{\sim}{x}) \delta\left(z-h_{2}(\underset{\sim}{x})\right) \theta\left(h_{z}\left({\underset{\sim}{x}}_{t}\right)-z\right) \quad . \tag{3.49}
\end{align*}
$$

where the normal vectors are defined as in (3.16) with the superscript indicating either $h_{2}$ or $h_{2}$. Evaluating the step functions in (3.49) at the value of the delta function we see that each is equal to 1 since $h_{1}>h_{2}$. The result is that the Green's tensor in (3.48) in the 1 ayer is given by contibutions from two surface integrals

$$
\begin{align*}
& G_{j n}\left(x_{\sim}^{n} \cdot x_{\sim}^{n}\right)=G_{j n}^{0}\left(x^{\prime} \cdot{\underset{\sim}{x}}^{n}\right)-\iint_{m}^{(1)}\left({\underset{\sim}{x}}_{t}\right)\left[{\underset{\sim}{x}}^{\prime} \cdot{\underset{\sim}{x}}_{s}^{(2)} \cdot x^{n}\right]_{m j n} d{\underset{\sim}{x}}_{t} \\
& +\iint_{m}^{(1)}\left({\underset{\sim}{x}}_{t}\right)\left[{\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s}^{(1)},{\underset{\sim}{x}}^{\prime \prime}\right]_{m j n}{ }^{d}{\underset{\sim}{x}}_{t}, \tag{3.50}
\end{align*}
$$

where the symbol terms are

$$
\begin{align*}
& n_{m}^{(p)}\left(x_{t}\right)\left[x^{\prime} \cdot{\underset{\sim}{x}}_{(p)}^{\left(x^{n}\right.}\right]_{m j n} \\
& =n_{m}^{(p)}(\underset{\sim}{x})\left[\left[G_{i j}^{0}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s}^{(p)}\right)\left[\operatorname{TG}\left(\underset{\sim}{x} \underset{s}{(p)},{\underset{\sim}{x}}^{\prime \prime}\right)\right]_{m i n}\right.\right. \\
& \left.\left.-\left[T G^{0}\left(\underset{\sim}{x},{\underset{\sim}{x}}_{s}^{(p)}\right)\right]_{\operatorname{mij}} G_{i n}{ }_{(\underset{\sim}{x}}^{(p)},{\underset{\sim}{x}}^{n}\right)\right], \tag{3.51}
\end{align*}
$$

for the $p^{\text {th }}$ surface.
If for simplicity we again choose both surfaces to be traction free then using (3.19) the symbol terms reduce to

$$
\begin{aligned}
& n_{m}^{(p)}\left({\underset{\sim}{t}}_{t}\right)\left[{\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{s}^{(p)},{\underset{\sim}{x}}^{\prime \prime}\right]_{\text {mj } n}
\end{aligned}
$$

so that (3.50) reduces to

$$
\begin{aligned}
& G_{j n}\left(x_{\sim}^{\prime} \cdot{\underset{\sim}{x}}^{n}\right)=G_{j n}^{0}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}^{n}\right)
\end{aligned}
$$

which expresses the value of the field in the waveguide in terms of its values on the two surfaces.

To generate the two coupled integral equations for these surface values, first let ${\underset{\sim}{\prime}}^{\prime} \rightarrow{\underset{X}{s}}^{\prime}(2)$. Since this approach is from above the surface ( $z^{\prime}>h_{z}\left({\underset{\sim}{x}}^{\prime}{ }^{\prime}\right)$ ) the signum function in the regularized term (A.31) is positive, and the result is

$$
\begin{aligned}
& =G_{j n}^{0}\left({\underset{\sim}{x}}^{(2)}, x^{n \prime}\right) \\
& +\iint \mathbb{K}_{j i}{\underset{\sim}{x}}_{s}^{(2)} \cdot{\underset{\sim}{x}}_{(2)}^{(2)} G_{i n}\left({\underset{\sim}{x}}_{(2)}^{(2)} \cdot{\underset{\sim}{x}}^{n}\right) d \underset{\sim}{x}
\end{aligned}
$$

which is a surface integral equation for the two surface values. Note that the second integral in (3.54) does not have a singular kernel since $h_{1} \neq h_{2}$.

To form the second integral equation, let $\underset{\sim}{x^{\prime} \rightarrow}{ }_{s}^{\prime}(1)$. This approach is
from below the surface $\left(z^{\prime}\left\langle h_{1}\left(x_{n}^{\prime}\right)\right)\right.$ so the signum function in the regularized term (A.31) is negative, and the result is

$$
\begin{aligned}
& Q_{j p}^{(1)}(\underset{\sim}{x}) G_{p n}\left({\underset{\sim}{x}}_{\prime}^{(1)},{\underset{\sim}{x}}^{n}\right) \\
& =G_{j n}^{0}\left({\underset{\sim}{x}}_{s}^{\prime}(1),{\underset{\sim}{x}}^{n}\right) \\
& -\iint K_{j i}\left({\underset{\sim}{x}}_{s}^{(1)},{\underset{\sim}{x}}_{s}^{(1)}\right) G_{i n}\left({\underset{\sim}{x}}_{s}^{(1)} \cdot{\underset{x}{ }}^{n}\right) d{\underset{\sim}{x}}_{t}
\end{aligned}
$$

which is our second integral equation for the two surface values. The procedure is thus to solve (3.54) and (3.55) for the surface values, and then to substitute the results into (3.53) to find the value of the fieldin the waveguide. In addition, if desired, the matrices $Q$ can be inverted, and the results witten in terms of what we called the surface values in Eg. 1 .

## Bg. 5. FLAT SORFACE AT ZBRO TRACTION

From example 2 we have from (3.39) the integral equation for the vector displacement values on the arbitrary surface $h(\underset{\sim}{x})$. For $h=0$ we have that

$$
\begin{align*}
& Q_{j m}(\underset{\sim}{x})=\frac{1}{2} \delta_{j m} \quad,  \tag{3.56}\\
& \mathbf{n}_{\mathbf{p}}({\underset{\sim}{x}})=\delta_{p 3} \quad . \tag{3.57}
\end{align*}
$$

and

$$
\begin{equation*}
K_{j i}\left({\underset{\sim}{x}}_{s}^{\prime},{\underset{\sim}{x}}_{s}\right) \rightarrow K_{j i}^{0}\left({\underset{\sim}{x}}_{t}^{\prime} \cdot{\underset{\sim}{x}}_{t}\right) \tag{3.58}
\end{equation*}
$$

where from (A.32) and (A.29) we get

$$
\begin{align*}
& =K_{3 i j}(\underset{\sim}{x}-\underset{\sim}{x}) \\
& =k_{T}^{-2}\left[R_{3 i j}^{T}\left(\underset{\sim}{x} t_{i}^{\prime}-{\underset{\sim}{x}}_{t}\right)-R_{3 i j}^{L}(\underset{\sim}{x}-\underset{\sim}{x})\right] \\
& -\frac{1}{2}\left[\delta_{i j} R_{3}^{T}\left({\underset{\sim}{x}}_{\prime}^{\prime}-{\underset{\sim}{x}}^{\prime}\right)+\delta_{3}{ }_{j}{ }_{i}^{T}\left(\underset{\sim}{x}{ }_{t}^{\prime}-\underset{\sim}{x}\right)\right. \\
& \left.+\delta_{3}[\lambda /(\lambda+2 \mu)] R_{j}^{L}\left({\underset{\sim}{x}}_{t}^{\prime}-x_{t}\right)\right] \quad . \tag{3.59}
\end{align*}
$$

and from (A.23) and (A.21) we have that
and

$$
\begin{align*}
& P_{3 i j}^{T, L}(\underset{\sim}{k})=2 i\left[\left[\left(k_{i t} k_{j t} \underline{K}_{T, L}^{2}+\delta_{i 3} \delta_{j 3} K_{T, L}^{4}\right] P\left[\frac{1}{k_{z}}\right]\right.\right. \\
& \left.\left.\left.+\left[\delta_{i 3} k_{j t}+\delta_{j 3} k_{i t}\right]\right]_{T, L}^{2}\right]\right] \text {. } \tag{3.61}
\end{align*}
$$

The exponential in (3.60) is independent of $k_{z}$ because of the flat surface limit and $G^{T} L_{(k)}$ are even functions of $k_{z}$, so the principal value term in (3.61) vanishes because it is an odd function of $\mathbf{k}_{\mathbf{z}}$. We have that

$$
\begin{align*}
R_{3 i j}^{T, L}\left(\underset{\sim}{x} \cdot-{\underset{\sim}{x}}^{\prime}\right)= & \frac{2 i}{(2 \pi)^{3}} \iiint \int_{\underset{\sim}{k}} e^{i \underset{\sim}{k} \cdot\left({\underset{\sim}{x}}^{\prime}-{\underset{\sim}{x}}_{t}\right)} K_{T, L}^{2} \\
& \cdot \widetilde{G}^{T, L}(k)\left[\delta_{i 3} k_{j t}+\delta_{j 3} k_{i t}\right] . \tag{3.62}
\end{align*}
$$

The $k_{z}$-integral can be evaluated

$$
\begin{equation*}
\int \mathrm{d} k_{z} \tilde{\mathrm{G}}^{T, L}(\mathrm{k})=\pi \mathrm{i} / \mathrm{K}_{\mathrm{T}, \mathrm{~L}}, \tag{3.63}
\end{equation*}
$$

so that we have

From (A.7) and (A.8) we have that

$$
\begin{equation*}
R_{j}^{T, L}\left(\underset{\sim}{x}{ }_{t}^{\prime}-\underset{\sim}{x}\right)=\frac{1}{(2 \pi)^{3}} \iiint d \underset{\sim}{k} e^{i \underset{\sim}{k}} \cdot{ }_{t}^{\cdot\left(\underset{\sim}{x} t^{\prime}-{\underset{\sim}{x}}_{t}\right)} G^{T, L}(k) P_{j}^{T, L}(\underset{\sim}{k}), \tag{3.65}
\end{equation*}
$$

where

$$
\begin{equation*}
P_{j}^{T, L}(\underset{\sim}{k})=2 i\left[\underline{k}_{j t}+\delta_{j 3}{ }^{K_{T, L}^{2}} P\left[\frac{1}{k_{z}}\right]\right] \tag{3.66}
\end{equation*}
$$

Again we can evaluate the $k_{z}$-integral since terms odd in $k_{z}$ vanish. The result is

$$
\begin{equation*}
R_{j}^{T, L}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right)=\frac{-1}{(2 \pi)^{2}} \iint d{\underset{\sim}{k}}_{t} e^{i{\underset{\sim}{k}}_{t} \cdot\left({\underset{\sim}{x}}_{\prime}^{\prime}-{\underset{\sim}{x}}_{t}\right)}{ }_{k_{j t}} / K_{T, L} . \tag{3.67}
\end{equation*}
$$

Osing (3.64) and (3.67) in (3.59) we get the result
where

$$
\begin{align*}
M_{j i}^{(0)}\left(k_{\sim_{t}}\right)= & \delta_{j 3} k_{i t}\left[k_{T}^{-2}\left(K_{T}-K_{L}\right)-\frac{1}{2} K_{T}^{-1}\right] \\
& +\delta_{i 3}{ }_{j, t}\left[k_{T}^{-2}\left(K_{T}-K_{L}\right)-\frac{1}{2} K_{L}^{-1}(\lambda /(\lambda+2 \mu))\right] \tag{3.69}
\end{align*}
$$

This form appears somewhat unsymmetric. It can be rewritten. The coefficient in the first term is

$$
\begin{equation*}
\mathbf{k}_{T}^{-2}\left(K_{T}-K_{L}\right)-\frac{1}{2} K_{T}=\left(K_{T}^{2}-2 K_{T} K_{L}-\mathbf{k}_{t}^{2}\right) / 2 \mathbf{K}_{T}^{2} K_{T} \tag{3.70}
\end{equation*}
$$

and, using the fact that

$$
\begin{equation*}
\lambda /(\lambda+2 \mu)=\left(\mathbf{k}_{\mathrm{T}}^{2}-2 \mathbf{k}_{\mathrm{L}}^{2}\right) / \mathbf{k}_{\mathrm{T}}^{2} \tag{3.71}
\end{equation*}
$$

the coefficient of the second term is

$$
\begin{equation*}
\mathbf{k}_{T}^{-2}\left(K_{T}-K_{L}\right)-\frac{1}{2} K_{L}^{-1}(\lambda /(\lambda+2 \mu))=\left(2 K_{L} K_{T}-K_{T}^{2}+\mathbf{k}_{t}^{2}\right) / 2 \mathbf{k}_{T}^{2} K_{L} \tag{3.72}
\end{equation*}
$$

and (3.69) becomes

$$
\begin{equation*}
M_{j i}^{(0)}(\underset{\sim}{k})=\left(\delta_{j 3} k_{i t} K_{T}^{-1}-\delta_{i 3} k_{j t} K_{L}^{-1}\right)\left(K_{T}^{2}-2 K_{T} K_{L}-k_{t}^{2}\right) / 2 k_{T}^{2} \tag{3.73}
\end{equation*}
$$

The resulting integral equations for the displacement on the surface is from (3.39)

$$
\begin{equation*}
\frac{1}{2} u_{j}\left({\underset{\sim}{x}}_{t}^{\prime}\right)=u_{j}^{i n}\left({\underset{\sim}{x}}_{t}^{\prime}\right)+\iint d{\underset{\sim}{x}}_{t} k_{j i}^{(0)}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right) u_{i}\left({\underset{\sim}{x}}_{t}\right) \tag{3.74}
\end{equation*}
$$

which are coupled convolution equations. They can be solved exactly. This is done in Sec. 13, after we first discuss the flat surface examples by more conventional means in Secs. 4-10. Once they are solved, the displacements in the field (i.e. off the surface) follow from (3.38)

$$
\begin{equation*}
u_{j}\left(x_{t}^{0}\right)=u_{j}^{i n}\left({\underset{\sim}{x}}_{t}^{\prime}\right)+\iint d{\underset{\sim}{x}}_{t}\left[T G^{0}\left({\underset{\sim}{x}}_{t}^{\prime},{\underset{\sim}{x}}_{t}\right)\right]_{3 i j} u_{i}\left(x_{t}\right) \tag{3.75}
\end{equation*}
$$

where the traction operator on the free-space Green's tensor follows from (3.12)

$$
\begin{align*}
& {\left[T G^{0}\left({\underset{\sim}{x}}^{\prime},{\underset{\sim}{x}}_{t}\right)\right]_{3 i j}=\left[\mu \frac{\partial}{\partial z} G_{i j}^{0}(\underset{\sim}{x}, \underline{\sim})+\mu \partial_{i} G_{3}^{0}(\underset{\sim}{x}, \underline{x})\right.} \\
& \left.+\lambda \delta_{i} 3 \partial_{m} G_{m j}^{0}\left(x^{\prime}, x\right)\right]_{z=0} \quad . \tag{3.76}
\end{align*}
$$

### 3.4 POTENTIALS AND PLANB VAVBS

We have been solving for either the Green's tensor $G_{i j}$ or the vector displacement

$$
\begin{equation*}
u_{i}(\underset{\sim}{x})=\iiint_{i j}\left(\underset{\sim}{x-x^{\prime}}\right) S_{j}\left(x_{\sim}^{\prime}\right) d x_{\sim}^{\prime} \tag{4.1}
\end{equation*}
$$

We now return to expressing this vector displacement in terms of potentials proviously mentioned in (1.19)-(1.22). We have that

$$
\begin{align*}
u_{i}(\underset{\sim}{x}) & =u_{i}^{L}(\underset{\sim}{x})+u_{i}^{T}(\underset{\sim}{x})  \tag{4.2}\\
& =\partial_{i} \phi+\varepsilon_{i j m} \partial_{j} A_{m} . \tag{4.3}
\end{align*}
$$

in terms of the scalar potential $\phi$ for the longitudinal waves (or P-waves) and the vector potential $A_{m}$ for the transverse waves (shear, or S-waves). The vector displacement has three independent components, and the scalar and vector potentials have four. We thus require a constraint on these potentials and it is usually written as

$$
\begin{equation*}
\partial_{\mathrm{m}} A_{\mathrm{m}}=0 \tag{4.4}
\end{equation*}
$$

in analogy to the gauge condition in electromagnetic theory. As we noted in (1.25) and (1.27) we have that

$$
\begin{equation*}
\Delta u_{i}^{L, T}(\underset{\sim}{x})+k_{L, T^{2}}^{\mathbf{u}_{i}^{L}, T}(\underset{\sim}{x})=0_{i} \tag{4.5}
\end{equation*}
$$

and we can thus choose equations on the potentials as

$$
\begin{equation*}
\Delta \phi(\underset{\sim}{x})+\mathrm{k}_{\mathrm{L}}^{2} \phi(\underset{\sim}{x})=0 \tag{4.6}
\end{equation*}
$$

and

$$
\begin{equation*}
\Delta A_{m}(\underset{\sim}{x})+k_{T}^{2} A_{m}(\underset{\sim}{x})=0_{m} \tag{4.7}
\end{equation*}
$$

Both the potentials thus satisfy lelmholtz equations with the corresponding longitudinal and transverse wavenumbers and wave speeds civen by

$$
\begin{equation*}
k_{L}=\omega / c_{L} \quad, \quad c_{L}=[(\lambda+2 \mu) / \rho]^{1 / 2} \tag{4.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathbf{k}_{T}=\omega / c_{T} \quad, \quad c_{T}=[\mu / \rho]^{1 / 2} \tag{4.9}
\end{equation*}
$$

Far enough away from any source, a wave front from that source can be (at least locally) approximated as a plane wave. We thus study plane wave sollutions of (4.6) and (4.7) given by

$$
\begin{equation*}
\phi(\underset{\sim}{x})=\exp \left[i \underset{\sim}{k}{ }^{L} \cdot \underset{\sim}{x}\right]=\exp \left[i{\underset{m}{k}}_{\mathrm{x}_{\mathrm{m}}}^{\mathrm{x}}\right] \tag{4.10}
\end{equation*}
$$

and
where

$$
\begin{equation*}
\left|\mathbf{k}_{\mathbf{k}}^{\mathbf{L}}\right|=\mathbf{k}_{\mathrm{L}} \quad \text { and } \quad|\underset{\sim}{\mathbf{k}}|=\mathbf{k}_{T} \tag{4.12}
\end{equation*}
$$

In each case the direction of propagation of the wave is $\mathfrak{k}^{L}$ and $\mathfrak{k}^{T}$ respectively. The direction of the longitudinal displacement is given by the gradient of (4.10), i.e.

$$
\begin{equation*}
\mathbf{u}_{i}^{L}(\underset{\sim}{x})=i{\underset{i}{L}}_{\mathbf{L}} \phi(\underset{\sim}{x}) \tag{4.13}
\end{equation*}
$$

and thus the longitudinal displacement is in the same direction as the
propagation of the wave. The displacement of the shear potential is in the direction given by the vector $a_{m}$. Because of the gauge condition (4.4) we thus have

$$
\begin{equation*}
k_{m}^{T} a_{m}=0 \tag{4.14}
\end{equation*}
$$

so that for the transverse wave the direction of propagation of the wave is orthogonal to the direction of its displacement. We thus have three independent potentials, $\phi$ for $P$-waves, and $A_{m}$ (with $\partial_{m} A_{m}=0$ ) for two independent components of shear waves, called SV and SH waves, where the V and $H$ notations are referred to as polarizations of the shear waves, again in analogy with electromagnetic theory. We thus have to choose these potentials to yield three wave shapes. It will turn out that we can choose two of them in a plane (but not orthogonal). These are the $P$ and $S V$ waves. The third one, the SH-wave, will be orthogonal to this plane. The wave number vectors ${\underset{\sim}{\mid}}^{\mathrm{L}}$ and ${\underset{\sim}{k}}^{\mathrm{T}}$ indicate the direction of propagation of the waves. Their components are the direction cosines. An alternative teminology of ten used in the seismic literature are the slowness vectors, related to the wavenumbers vectors as

$$
{\underset{\sim}{s}}^{\mathrm{L}, \mathrm{~T}}=\hat{\mathbf{k}}^{\mathrm{L}, \mathrm{~T}} / \mathrm{c}_{\mathrm{L}, \mathrm{~T}} .
$$

The $x$-component of either of the vectors is the same. It is $j u s t$ the ray parameter

$$
\begin{equation*}
p=\sin \theta_{L} / c_{L}=\sin \theta_{T} / c_{T} \tag{4.15}
\end{equation*}
$$

and the equality on the rhs is j ust an expression of Snell's Law.
For $\underline{\text { S-waves }}$ we choose the propagation direction by choosing the vector
${\underset{\sim}{k}}^{T}$ in the $x-z$ plane

$$
\begin{equation*}
\underset{\sim}{k}{ }^{T}=k_{z}^{T} \underset{z}{A}+k_{z}^{T} \underset{k}{a}, \tag{4.16}
\end{equation*}
$$

where $i$ and $k$ are unit vectors along the $x$ and $z$ directions. The vector potential is thus

$$
\begin{equation*}
A_{m}(x)=a_{m} \exp \left(i k_{x}^{T}+i k_{z}^{T}\right)=A_{m}(x, z) \tag{4.17}
\end{equation*}
$$

which is only a function of $x$ and $z$. The gage condition (4.4) yields

$$
\begin{equation*}
\frac{\partial}{\partial z} A_{1}+\frac{\partial}{\partial z} A_{3}=0 \tag{4.18}
\end{equation*}
$$

For the SV-polarization we assume in addition that the y-component of displacement vanishes, i.e. that

$$
\begin{equation*}
u_{2}^{T}=\varepsilon_{2 j m} \partial_{j} A_{m}=\frac{\partial A_{1}}{\partial z}-\frac{\partial A_{3}}{\partial z}=0 \tag{4.19}
\end{equation*}
$$

The result of (4.18) and (4.19) is that $A_{1}$ andf $A_{3}$ satisfy Cauchy-Riemann equations, and thus that $A_{3}+i A_{1}$ is an analytic function of $x+i z$ which is analytic everywhere. In addition, for $A_{m}$ expressible as a plane wave, the complex function $A_{3}+i A_{1}$ is bounded. Liouville's theorem states that a function which is evexywhere analytic and bounded is a constant. Since only gradients of $A_{3}$ and $A_{1}$ are used to calculate a physically measurable quantity such as displacement, the constant doesn't matter, and we can choose it to be zero. Thus $A_{3}=A_{1}=0$ and the vector potential has only one component

$$
\begin{equation*}
\underset{\sim}{A}=(0, A, 0) \quad, \tag{4.20}
\end{equation*}
$$

whose divergence vanishes by (4.4). This latter is true if A is not a function of $y$. Thus the displacement for $S V$-waves can be written as

$$
\begin{align*}
u_{i}^{T, S V}(x, z) & =\varepsilon_{i j m} \partial_{j} A_{m}(x, z) \\
& =\left(-\frac{\partial}{\partial z} A(x, z), 0, \frac{\partial}{\partial x} A(x, z)\right), \tag{4.21}
\end{align*}
$$

Which has components in and $z$ directions only. By (4.7) A satisfies the differential equation

$$
\begin{equation*}
\Delta A(x, z)+\mathbf{k}_{T}^{2} A(x, z)=0 \tag{4.22}
\end{equation*}
$$

where the Laplacian is only in $x$ and $z$. Note that we not only have $\mathbb{k}^{T} \cdot \underset{\sim}{A}=0$ but also that

$$
\begin{equation*}
{\underset{\sim}{\mathbf{k}}}^{T} \cdot{\underset{\sim}{\mathbf{u}}}^{\mathrm{T}, \mathrm{SV}}=0 \tag{4.23}
\end{equation*}
$$

explicitly illustrating that $S V$ shear displacements are orthogonal to the propagation direction.

For the SH-polarization, we assume that the displacement is orthogonal to $S V$, i.e. that it has no components in the $x-z$ plane. We can write this displacement as

$$
\begin{equation*}
\mathbf{u}_{i}^{T, S H}(x, z)=(0, v(x, z), 0) \tag{4.24}
\end{equation*}
$$

which thus only has a $y$-component. The reason that $v$ is not a function of $y$ is that we must have

$$
\begin{equation*}
\partial_{i} u_{i}^{T, S H}=0 \tag{4.25}
\end{equation*}
$$

as we previously noted. This implies $\partial v / \partial y=0$ and again we drop any constant, or absorb it into the other functional dependence. In addition from (4.5) we have that

$$
\begin{equation*}
\left(\Delta+k_{T}^{2}\right) v(x, z)=0 \tag{4.26}
\end{equation*}
$$

For the $P$-wave, in the $x-z$ plane, there is no displacement in the $y-$ direction so that

$$
\begin{equation*}
u_{2}^{L, P}=\frac{\partial \phi}{\partial y}=0 \text { and } \phi=\phi(x, z) \tag{4.27}
\end{equation*}
$$

We thus have that

$$
\begin{equation*}
\mathbf{u}_{j}^{L}(x, z)=\left(\frac{\partial}{\partial x} \phi(x, z), 0, \frac{\partial}{\partial z} \rho(x, z)\right) \tag{4.28}
\end{equation*}
$$

Notice from (4.21), (4.24) and (4.28), $P$ and $S V$ waves decouple from $S H$ waves.

For the Green's tensor we defined the operator $T$ as in (3.11). Dotting this with the normal we get the traction

$$
\begin{align*}
n_{p}\left(z_{t}\right)\left[\operatorname{TG}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime \prime}\right)\right]_{p i j}= & \mu n_{m}(\underset{\sim}{x}) \partial_{m} G_{i j}(\underset{\sim}{x}, \underset{\sim}{x}) \\
& +\mu n_{m}\left({\underset{\sim}{x}}_{t}\right) \partial_{i} G_{m j}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right) \\
& +\lambda n_{i}\left({\underset{\sim}{x}}_{t}\right) \partial_{m} G_{m j}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{n}\right) \tag{4.29}
\end{align*}
$$

If we integrate this equation over $\underset{\sim}{x}$ " and a vector source $S_{j}\left({\underset{\sim}{x}}^{\prime \prime}\right)$, we get that the Green's function becomes the displacement and we can define the
(vector) stress as

$$
\begin{align*}
\tau_{z i} & =n_{p}\left({\underset{\sim}{x}}_{t}\right)[T u(\underset{\sim}{x})]_{p i} \\
& =\mu n_{m}\left({\underset{\sim}{x}}_{t}\right) \partial_{m} u_{i}(\underset{\sim}{x})+\mu n_{m}\left({\underset{\sim}{x}}_{t}\right) \partial_{i} u_{m}(\underset{\sim}{x})+\lambda_{i} n_{i}\left({\underset{\sim}{t}}_{t}\right) \partial_{m} u{ }_{m}\left(x_{\sim}\right), \tag{4.30}
\end{align*}
$$

the first term of which is a normal derivative, and the third term is a divergence. Its three components correspond to normal ( $i=3=z$ ) and tangential ( $i=1,2=x, y$ ) stresses. (The additional stress components which arise by replacing $z$ on the 1 hs of (4.30) by $x$ or $y$ only act in the plane, and do not act across a boundary either normally or tangentially.) We can explicitly write the displacement as

$$
\begin{align*}
u_{i}(x, z) & =u_{i}^{L}(x, z)+u_{i}^{T}(x, z) \\
& =\delta_{i 1} u_{i}(x, z)+\delta_{i 2} v(x, z)+\delta_{i 3} u_{3}(x, z), \tag{4.31}
\end{align*}
$$

where, from (4.21) and (4.28) we have that

$$
\begin{equation*}
u_{1}(x, z)=\frac{\partial \phi}{\partial x}-\frac{\partial A}{\partial z} \tag{4.32}
\end{equation*}
$$

and

$$
\begin{equation*}
u_{3}(x, z)=\frac{\partial \phi}{\partial z}+\frac{\partial A}{\partial x} \tag{4.33}
\end{equation*}
$$

From (4.30) the stress becomes

$$
\begin{align*}
\tau_{z i}= & \mu n_{m}\left({\underset{\sim}{x}}_{t}\right) \partial_{m}\left[\delta_{i 1} u_{1}(x, z)+\delta_{i 2} v(x, z)+\delta_{i 3} u_{3}(x, z)\right] \\
& +\mu n_{m}\left({\underset{\sim}{x}}_{t}\right) \partial_{i}\left[\delta_{m 1} u_{1}(x, z)+\delta_{m 2} v(x, z)+\delta_{m 3} u_{3}(x, z)\right] \\
& +\lambda n_{i}\left({\underset{\sim}{x}}_{t}\right) \partial_{m}\left[\delta_{m 1} u_{1}(x, z)+\delta_{m 2} v(x, z)+\delta_{m 3} u_{3}(x, z)\right] \tag{4.34}
\end{align*}
$$

Since we have discussed plane waves by themselves (i.e. no superposition of plane wavs) and these are appropriate for planar iterface problems, we write the stress components for a flat interface (0superscript) where $n_{m}\left({\underset{\sim}{f}}_{t}\right)=\delta_{m 3}$. These are easily seen to be from (4.34)

$$
\begin{align*}
& \tau_{z 1}^{0}=\mu\left[\frac{\partial^{2} A}{\partial x^{2}}-\frac{\partial^{2} A}{\partial z^{2}}\right]+2 \mu \frac{\partial^{2} \phi}{\partial x^{2} z},  \tag{4.35}\\
& \tau_{z 2}^{0}=\mu \frac{\partial v}{\partial z}, \tag{4.36}
\end{align*}
$$

and

$$
\begin{equation*}
\tau_{z 3}^{0}=2 \mu\left[\frac{\partial^{2} \phi}{\partial z^{2}}+\frac{\partial^{2} A}{\partial z \partial x}\right]-\lambda{k_{L}^{2} \phi}_{L} \tag{4.37}
\end{equation*}
$$

We can further reduce these stresses by quoting the results for the various polarizations separately as

P-wave

$$
\begin{equation*}
\tau_{z 1}=2 \mu \frac{\partial^{2} \phi}{\partial x \partial z}, \tau_{z 2}=0, \tau_{z 3}=2 \mu \frac{\partial^{2} \phi}{\partial z^{2}}-\lambda k_{L}^{2} \phi, \tag{4.38}
\end{equation*}
$$

SV-wave

$$
\begin{equation*}
\tau_{z 1}=\mu\left[\frac{\partial^{2} A}{\partial x^{2}}-\frac{\partial^{2} A}{\partial z^{2}}\right], \tau_{z 2}=0, \tau_{z 3}=2 \partial^{\partial^{2} A} \frac{\partial z \partial x}{\partial z} \text {, } \tag{4.39}
\end{equation*}
$$

and
SH-wave

$$
\begin{equation*}
\tau_{z 1}=\tau_{z 3}=0, \quad \tau_{z 2}=\mu \frac{\partial v}{\partial z} . \tag{4.40}
\end{equation*}
$$

These latter three equations are only useful provided the waves do not mix.

### 3.5 BODNDARY CONDITIONS

At the interface between two elastic media (illustrated in Fig. 3.3 be 1 ow)

(2) $\lambda_{2}, \mu_{2}, \rho_{2}$

Fig. 3.3

We can have several types of boundary conditions.
(a) Rigid Contact

For this case both displacements $\mathbf{u}_{j}$ and stresses $\tau_{z j}$ are continuous. ?
That is we have six conditions ( $\mathrm{j}=1,2,3$ )

$$
\begin{equation*}
u_{j}^{(1)}(\underset{\sim}{x})=u_{j}^{(2)}\left({\underset{\sim}{x}}_{s}\right), \tag{5.1}
\end{equation*}
$$

and

$$
\begin{equation*}
\tau_{z j}^{(1)}\left({\underset{\sim}{x}}_{s}\right)=\tau_{z j}^{(2)}\left({\underset{\sim}{x}}^{x}\right) . \tag{5.2}
\end{equation*}
$$

where $\underset{\sim}{x}{ }_{s}=(x, y, h)$ is a position vector on the surface, and the superscript symbols (1) and (2) indicate the displacement or stress in the particular region as evaluated on the boundary.
(b) Free Contact

For this case the stresses are continuous, but the displacements are discontinuous.
(c) Free Surface

For this case we have that the stresses vanish on the surface

$$
\begin{equation*}
\tau_{z j}(\underset{\sim}{x})=0_{j} . \tag{5.3}
\end{equation*}
$$

and the displacements are not specified. The latter can be easily computed for a flat interface and we do so in the next several sections. This boundary condition is also called the zero stress or zero traction or traction free condition.
(d) Infinitely Rigid Surface

For this case, the displacements vanish on the surface, i.e.

$$
\begin{equation*}
u_{j}\left({\underset{\sim}{x}}_{s}\right)=0_{\cdot j} \tag{5.4}
\end{equation*}
$$

The stresses are not specified a priori, but can be computed.

## RLOID-RLASTIC BOUNDARY

At a fluid-fluid interface we know that we have the continuity of pressure and normal velocity. The question is how do we pass from the continuity conditions at an elastic-elastic interface to those at a fluidelastic interface.
(a) Continuity of Normal Velocity

For an elastic-elastic interface we have continuity of displacements as in (5.1). In general the displacement is time dependent, and the time derivative of displacement is velocity. Since we have time-harmonic problems we can work directly with displacements. In a fluid, the tangential displacements are zero, i.e.

$$
\begin{equation*}
u_{1}\left({\underset{\sim}{x}}_{s}\right)=u_{2}\left({\underset{\sim}{x}}_{s}\right)=0 \tag{5.5}
\end{equation*}
$$

and the normal displacements (equivalently normal velocities) are continuous

$$
\begin{equation*}
u_{3}^{(1)}\left(\underset{\sim}{x_{s}}\right)=u_{3}^{(2)}\left(\underset{\sim}{x_{s}}\right) \tag{5.6}
\end{equation*}
$$

(b) Continuity of Pressure

In a fluid the shear modulus vanishes, $\mu=0$. The shear wave sped al so vanishes, $c_{T}=0$, and the Lame' modulus is expressible in terms of the longitudinal (compressional) speed $c_{L}$ as $\lambda=\rho c_{L}^{2}$ where $\rho$ is density. We have continuity of stress. Eq. (5.2), and in a fluid the stress is using (4.30) and $\mu=0$

$$
\begin{equation*}
\tau_{z j}=\lambda n_{j} \partial_{m}^{u}{ }_{m} \tag{5.7}
\end{equation*}
$$

In a fluid, the time-derivative of the time-dependent displacement is just the velocity, which can be represented as the gradient of the scalar velocity potential \&. i.e.

$$
\begin{equation*}
\frac{\partial}{\partial t} u_{j}(x, t)=\partial_{j} \Phi \tag{5.8}
\end{equation*}
$$

For harmonic time-dependence $\exp (-i \omega t)$ we have

$$
\begin{equation*}
u_{j}(x)=(i / \omega) \partial_{j} \Phi(\underset{\sim}{x}) \tag{5.9}
\end{equation*}
$$

The divergence of this can be written as

$$
\begin{equation*}
\partial_{j} \mathbf{u} j(\underset{\sim}{x})=(i / \omega) \partial_{j} \partial_{j} \Phi(\underset{\sim}{x})=-(i / \omega) k_{0}^{2} \Phi \text {. } \tag{5.10}
\end{equation*}
$$

where $k_{0}=\omega / c_{L}$, and where the latter tem on the rhs of (5.10) follows since satisfies the acoustic Relmholtz equation. Substituting (5.10) into (5.7) using $k_{0}$ and $\lambda$ definitions we get

$$
\begin{equation*}
\tau_{z j}=n_{j}(-i \omega \rho \Phi) \tag{5.11}
\end{equation*}
$$

The pressure can be written in terms of the velocity potentials

$$
\begin{equation*}
p=-\rho \frac{\partial}{\partial t} \Phi=i \omega \rho \Phi \tag{5.12}
\end{equation*}
$$

so that, in a f1uid,

$$
\begin{equation*}
\tau_{\mathbf{z j}}=-\mathbf{n}_{\mathbf{j}} \mathbf{p} \tag{5.13}
\end{equation*}
$$

For a flat interface, $\mathbf{n}_{\mathbf{j}}=\delta_{j}$ and the stress results are

$$
\begin{equation*}
p=-\tau_{z z} \quad, \quad 0=\tau_{z x}, \quad 0=\tau_{z y} \tag{5.14}
\end{equation*}
$$

In general for an arbitrary surface the boundary conditions are expressed as the continuity of normal velocity

$$
\begin{equation*}
n_{j} v_{j}=\frac{\partial}{\partial t}\left(n_{j} u_{j}\right) \tag{5.15}
\end{equation*}
$$

the continuity of normal stress

$$
\begin{equation*}
-\mathbf{p}=\tau_{\mathbf{i} \mathbf{j}} \mathbf{n}_{\mathbf{i}} \mathbf{n}_{\mathbf{j}} \tag{5.16}
\end{equation*}
$$

and the vanishing of the elastic shear stresses

$$
\begin{equation*}
0=\tau_{i j} n_{i}{ }^{L}{ }_{j}=\tau_{i j}{ }^{n}{ }_{i}{ }_{j} \tag{5.17}
\end{equation*}
$$

where $L_{j}$ and $v_{j}$ are orthogonal vectors in the local tangent plane.

## ELECTROMAGNETIC THEORI

As an aside we note that we can recover the results of electromagnetic theory from those of elasticity. The equation for dispalcement is (using vector analysis notation) from (1.16)

$$
\begin{equation*}
\mu \Delta \underset{\sim}{\mathbf{u}}+(\lambda+\mu) \operatorname{grad}(\underset{\sim}{\nabla} \cdot \underset{\sim}{u})+\mathbb{K}^{2} \underset{\sim}{\mathbf{u}}=\underset{\sim}{0}, \tag{1.16}
\end{equation*}
$$

where $\Delta=$ grad div-curlcur1. From the Maxwell equations on the electric $\underset{\sim}{E}$ and magnetic H fields

$$
\begin{equation*}
\underset{\sim}{\nabla} \times \underset{\sim}{E}=i k \underset{\sim}{\mathbb{H}} \quad, \underset{\sim}{\nabla} \times \underset{\sim}{\mathbb{H}}=-i \underset{\sim}{E}, \tag{5.18}
\end{equation*}
$$

we derive, by taking the curl of the first equation, an equation on $\underset{\sim}{\mathrm{E}}$

$$
\begin{equation*}
\underset{\sim}{\nabla} \underset{\sim}{\nabla} \times \underset{\sim}{E}-\mathbf{k}^{2} \underset{\sim}{E}=\underset{\sim}{0} \text {, } \tag{5.19}
\end{equation*}
$$

or

$$
\begin{equation*}
\Delta \underset{\sim}{\mathrm{E}}+\mathbf{k}^{2} \underset{\sim}{E}-\operatorname{grad}(\underset{\sim}{\nabla} \cdot \underset{\sim}{\mathrm{E}})=\underset{\sim}{0}, \tag{5.20}
\end{equation*}
$$

If we make the make the formal interchanges

$$
\begin{equation*}
\underset{\sim}{\mathbf{E}} \rightarrow \underset{\sim}{\mathbf{u}} ; \quad \mathbf{k}^{2}=\mathbf{K}^{2} / \mu=\mathbf{k}_{\mathbf{T}}^{2} \text {. } \tag{5.21}
\end{equation*}
$$

and set $\lambda+2 \mu=0 \quad\left(c_{L}=0\right)$, then ( 5.20 is $j u s t$ (1.16). Note that the electromagnetic wave number $k$ becomes the transverse (shear) wavenumber $\mathbb{k}_{\mathrm{T}}$, whereas in the fluid the acoustic wavenumber $k_{0}$ became $k_{L}$, the longitudinal (compressiona1) wavenumber.

### 3.6 P- WAVE INCIDENCE ON $\triangle$ FREB SURFACE

We noted in Sec. 4 that $P$ and $S V$ waves couple and $S H$ waves decouple from these. A flat interface does not alter this property, and here we consider $p$-wave incidence on a free (zero-traction) flat boundary located at $z=0$. From (5.3) we thas have three boundary conditions

$$
\begin{equation*}
\tau_{z j}=0 \quad j=1,2,3 \tag{6.1}
\end{equation*}
$$

but from (4.36) the $j=2$ condition is antomatically satisfied since we have no SH waves. From (4.35) and (4.37) we can thus write the two boundary conditions as

$$
\begin{equation*}
\frac{\partial^{2} A}{\partial x^{2}}-\frac{\partial^{2} A}{\partial z^{2}}+2 \frac{\partial^{2} \phi}{\partial x^{\partial z}}=0 \tag{6.2}
\end{equation*}
$$

and

$$
\begin{equation*}
2 \mu\left[\frac{\partial^{2} \phi}{\partial z^{2}}+\frac{\partial^{2} A}{\partial x \partial z}\right]-\lambda k_{L \phi}^{2} \phi=0 \tag{6.3}
\end{equation*}
$$

where the potentials $\phi$ and $A$ are evaluated at $z=0$.
Assume the $p$-wave has incident ( $A_{0}, i$ ) and reflected (B,r) plane wave components
and the $S V$ wave only a reflected component

$$
\begin{equation*}
A(x, z)=C \exp \left[i\left(k_{x}^{T, r} r_{x}+k_{z}^{T, r} z\right)\right] . \tag{6.5}
\end{equation*}
$$

where the wavenumber components satisfy the dispersion relations from (4.6), (4.7) and (4.20)

$$
\begin{equation*}
\left(k_{x}^{L, i}\right)^{2}+\left(k_{z}^{L, i}\right)^{2}=\left(k_{x}^{L, r}\right)^{2}+\left(k_{z}^{L, r}\right)^{2}=k_{L}^{2}, \tag{6.6}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(k_{x}^{T}, \mathbf{r}\right)^{2}+\left(k_{z}^{T}, \mathbf{r}\right)^{2}=k_{T}^{2} \tag{6.7}
\end{equation*}
$$

We have two boundary conditions but three unknown constants $A_{0}, B$, and $C$ and we solve for the ratios $R_{1}=B / A_{0}=R_{P \rightarrow P}$, the reflection coefficient (amplitude ratio) for scattering to $P$-waves with a P-wave incident field, and $R_{2}=C / A_{0}=R_{P \rightarrow S V}$, the reflection coefficient for scattering from $P$ to SV waves. Note these are reflection coefficients for potentials. Substituting (6.4) and (6.5) into (6.2) and (6.3) we get the equations

$$
\begin{equation*}
a_{12} R_{2} e^{i k_{x}^{T, r} x}-a_{12} R_{1} e^{i k_{x}^{L, r} x}=-a_{12} e^{i k_{x}^{L, i} x} . \tag{6.8}
\end{equation*}
$$

and

$$
\begin{equation*}
a_{22} R_{2} e^{i k_{x}^{T, r}}+a_{22} R_{1} e^{i k_{x}^{L, T} x}=-a_{22} e^{i k_{x}^{L, i} x} \tag{6.9}
\end{equation*}
$$

where

$$
\begin{align*}
& a_{11}=\left(k_{z}^{T, r}\right)^{2}-\left(k_{X}^{T}\right)^{2} \quad,  \tag{6.10}\\
& a_{12}=2{\underset{k}{L}}_{L, r}^{k_{z}^{L, r}},  \tag{6.11}\\
& a_{21}=2 \mu k_{x}^{T, r} k_{z}^{T, r} \quad, \tag{6.12}
\end{align*}
$$

and

$$
\begin{equation*}
a_{22}=2 \mu\left(k_{z}^{L}, \mathbf{r}\right)^{2}+\lambda \mathbf{k}_{L}^{2} \tag{6.13}
\end{equation*}
$$

Since we have plane wave incidence, the solution of (6.8) and (6.9) should be independent of $x$. Any point of incidence will do, no preferred value of $x$ is possible, or equivalently, the equations must be translationally invariant in $x$. This is true in (6.8) and (6.9) provided

$$
\begin{equation*}
k_{x}^{L, i}=k_{x}^{L, T}=k_{x}^{T, r} \tag{6.14}
\end{equation*}
$$

The lhs of this equation is true if

$$
\begin{equation*}
k_{L} \sin \theta_{L i}=k_{L} \sin \theta_{L r} \tag{6.15}
\end{equation*}
$$

where the angles are defined in Fig. 3.4 later in this section.. The 1 atter is true if the angle of incidence and the angle of reflection (of the P-wave) are equal, $\theta_{L_{i}}=\theta_{L_{r}}$. The right hand equality in (6.14) is true provided

$$
\begin{equation*}
k_{L} \sin \theta_{L i}=k_{T} \sin \theta_{T r} \tag{6.16}
\end{equation*}
$$

so that the sin of the angle of the reflected shear wave is

$$
\begin{equation*}
\sin \theta_{T r}=\left(k_{L} / k_{T}\right) \sin \theta_{L i}=\left(c_{T} / c_{L}\right) \sin \theta_{L i} . \tag{6.17}
\end{equation*}
$$

This is just Snell's Law (for reflection of $P$ - to SV-waves) and can be expressed in terms of the ray parameter (4.15). Dsing these results the equations (6.8) and (6.9) reduce to

$$
\begin{equation*}
a_{11} R_{2}-a_{13} R_{1}=-a_{12} \text {, } \tag{6.18}
\end{equation*}
$$

and

$$
\begin{equation*}
a_{21} R_{2}+a_{22} R_{1}=-a_{22} \text {, } \tag{6.19}
\end{equation*}
$$

which have the solution

$$
\begin{equation*}
R_{1}=R_{P \rightarrow P}=\left(a_{12} a_{21}-a_{111_{22}}\right) / \Delta \text {. } \tag{6.20}
\end{equation*}
$$

and

$$
\begin{equation*}
R_{2}=R_{P \rightarrow S V}=-2 a_{12} a_{22} / \Delta \tag{6.21}
\end{equation*}
$$

where

$$
\begin{equation*}
\Delta=a_{12} a_{21}+a_{11} a_{22} \tag{6.22}
\end{equation*}
$$

Noting that

$$
\begin{equation*}
\mathbf{k}_{\mathrm{z}}^{\mathrm{T}, \mathbf{r}}=\mathbf{k}_{\mathrm{T}} \cos \theta_{T_{r}} \quad, \quad \mathbf{k}_{\mathrm{x}}^{\mathrm{T}, \mathrm{r}}=\mathbf{k}_{\mathrm{T}} \sin \theta_{\mathrm{T}_{r}} \text {. } \tag{6.23}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathbf{k}_{\mathrm{X}}^{\mathrm{L}, \mathrm{r}}=\mathbf{k}_{\mathrm{L}} \cos \theta_{\mathrm{Li}} ; \quad \mathbf{k}_{\mathrm{z}}^{\mathrm{L}, \mathrm{r}}=\mathbf{k}_{\mathrm{L}} \sin \theta_{L i} \text {. } \tag{6.24}
\end{equation*}
$$

we can write the $a_{i j}$ coefficients in many ways, viz.

$$
\begin{align*}
a_{11} & =\mathbf{k}_{T}^{2}\left(\cos ^{2} \theta_{T_{r}}-\sin ^{2} \theta_{T_{r}}\right) \\
& =\mathbf{k}_{T}^{2}\left(1-2 \sin ^{2} \theta_{T_{r}}\right) \\
& =\mathbf{k}_{T}^{2}\left(1-2\left(c_{T} / c_{L}\right)^{2} \sin ^{2} \theta_{L i}\right) \\
& =\mathbf{k}_{T}^{2}\left(1-2 p^{2} c_{T}^{2}\right) \tag{6.25}
\end{align*}
$$

where $p$ is the ray parameter from (4.15) and

$$
\begin{align*}
a_{12} & =2 k_{L}^{2} \sin \theta_{L_{i}} \cos \theta_{L_{i}} \\
& =2 k_{L}^{2} p c_{L}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}  \tag{6.26}\\
a_{21} & =2 \mu k_{T}^{2} \sin \theta_{T r} \cos \theta_{T r} \\
& =2 \mu k_{T}^{2}\left(c_{T} / c_{L}\right) \sin \theta_{L i}\left[1-\left(c_{T} / c_{L}\right)^{2} \sin ^{2} \theta_{L i}\right]^{1 / 2} \\
& =2 \mu k_{T}^{2} p c_{T}\left(1-p^{2} c_{T}^{2}\right)^{1 / 2} \\
& =2 \rho k_{T}^{2} p c_{T}^{3}\left(1-p^{2} c_{T}^{2}\right)^{1 / 2}, \tag{6.27}
\end{align*}
$$

and

$$
\begin{align*}
a_{22} & =k_{L}^{2}\left(2 \mu \cos ^{2} \theta_{L i}+\lambda\right) \\
& =k_{L}^{2}\left[(2 \mu+\lambda)-2 \mu \sin ^{2} \theta_{L i}\right] \\
& =k_{L}^{2} \rho\left(c_{L}^{2}-2 c_{T}^{2} \sin ^{2} \theta_{L i}\right) \\
& =\rho k_{L}^{2} c_{L}^{2}\left(1-2 p^{2} c_{T}^{2}\right) \tag{6.28}
\end{align*}
$$

Osing the fe forms we can write the reflection coefficients in various ways. We 1ist two for each

$$
\begin{equation*}
\frac{B}{A_{0}}=R_{P \rightarrow P}=\frac{4 p^{2} c_{t}^{3}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-p^{2} c_{t}^{2}\right)^{1 / 2}-c_{L}\left(1-2 p^{2} c_{t}^{2}\right)^{2}}{4 p^{2} c_{t}^{3}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-p^{2} c_{t}^{2}\right)^{1 / 2}+c_{L}\left(1-2 p^{2} c_{t}^{2}\right)^{2}} \tag{6.29}
\end{equation*}
$$

or, using the fact that

$$
\begin{equation*}
\cos \theta_{L i}=\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}, \cos \theta_{t r}=\left(1-p^{2} c_{L}^{2}\right)^{1 / 2} . \tag{6.30}
\end{equation*}
$$

we get

$$
\begin{equation*}
R_{P \rightarrow P}=\frac{4 p^{2}\left(\cos \theta_{L i} / c_{L}\right)\left(\cos \theta_{T r} / c_{T}\right)-\left(c_{T}^{-2}-2 p^{2}\right)^{2}}{4 p^{2}\left(\cos \theta_{L i} / c_{L}\right)\left(\cos \theta_{T r} / c_{T}\right)+\left(c_{T}^{-2}-2 p^{2}\right)^{2}} \tag{6.31}
\end{equation*}
$$

Similarly we can derive

$$
\begin{equation*}
\frac{C}{A_{0}}=R_{P \rightarrow S V}=\frac{-4 p c_{T}^{2}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-2 p_{p}^{2} c_{T}^{2}\right)}{4 p^{2} c_{T}^{3}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-p^{2} c_{T}^{2}\right)+c_{L}\left(1-2 p^{2} c_{T}^{2}\right)^{2}} \tag{6.32}
\end{equation*}
$$

or

$$
\begin{equation*}
R_{P \rightarrow S V}=\frac{-4 p\left(\cos \theta_{L i} / c_{L}\right)\left(c_{T}^{-2}-2 p^{2}\right)}{4 p^{2}\left(\cos \theta_{L i} / c_{L}\right)\left(\cos \theta_{t r} / c_{T}\right)+\left(c_{T}^{-2}-2 p^{2}\right)^{2}} \tag{6.33}
\end{equation*}
$$

The conventional representation of this scattering process is illustrated in Fig. 3.4.


The straight 1 ines indicate the direction of propagation and the arrows indicate the direction of displacement of the wave. All the $x$-components of the wave vectors are equal from translational invariance in the x -direction. Note that the $P$-wave displacements

$$
\begin{equation*}
\mathbf{u}_{\mathbf{j}}^{\mathrm{L}}=\boldsymbol{\partial}_{\mathbf{j}} \phi \tag{6.34}
\end{equation*}
$$

are along the direction of propagation, whereas the SV-wave displacements

$$
\begin{equation*}
u_{j}^{T}=\varepsilon_{j m p} \partial_{m} A_{p}=-\delta_{j i} \frac{\partial A}{\partial z}+\delta_{j 3} \frac{\partial A}{\partial x}, \tag{6.35}
\end{equation*}
$$

are orthogonal to the propagation direction.
(proof) Find the displacements of the SV-wave

$$
A(x, z)=C \exp \left[i\left(k_{z}^{T, r_{x}}+k_{z}^{\left.T, r_{z}\right)}\right]\right.
$$

The $z$-component of displacement on the $z=0$ surface is

$$
\frac{\partial}{\partial x} A(x, 0)=i k_{x}^{T, r} C \exp \left(i k_{x}^{T, r} x\right)
$$

Since we are using phasors we have to take the real part of this to get the physical displacement

$$
\operatorname{Re} \frac{\partial}{\partial x} A(x, 0)=-k_{x}^{T, r} \sin \left(k_{x}^{T, r} x\right) C
$$

The $x$-component of displacement is
$\operatorname{Re}\left[-\frac{\partial}{\partial z} A(x, 0)\right]=k_{x}^{t, r} \sin \left(k_{x}^{t, r} x\right) C \quad$,
which is positive. Thus the z-component of displacement is negative, as we have illustrated in the figure.

## CONVRNTIONS

As we remarked, the terms $R_{P \rightarrow P}$ and $R_{P \rightarrow S V}$ we have calculated are the reflection coefficients for the potentials. What are often quoted are reflection coefficients for displacements. There are several posible definitions for these. We have that the displacement for the incident P-wave is $\boldsymbol{\partial}_{\mathbf{j}} \phi^{i n}$, for the scattered $P_{\text {-wave }} \boldsymbol{\partial}_{\mathbf{j}} \boldsymbol{\phi}^{S C}$, and for the $S V$ waves given by (6.35). The reflection coefficient for the z-component of $p$-wave displacement is

$$
R_{P \rightarrow P}^{z}=\left.\frac{\partial \phi^{S C} / \partial z}{\partial \phi^{i n} / \partial z}\right|_{z=0}=-\frac{B}{A_{0}},
$$

and for the $x$-component of $p$-wave displacement

$$
\mathbb{R}_{\mathrm{P} \rightarrow \mathrm{P}}^{\mathrm{I}}=\left.\frac{\partial \phi^{S C} / \partial x}{\partial \phi^{i n} / \partial x}\right|_{z=0}=\frac{B}{A_{0}}
$$

The reflection coefficient for the $z$-componemt of SV-waves due to P-wave incidence is

$$
R_{P \rightarrow S V}^{Z}=\left.\frac{\partial A / \partial x}{\partial \phi^{i n} / \partial z}\right|_{z=0}=\frac{k_{z}^{T, r}}{k_{z, i}^{L, i}} \frac{C}{A_{0}}=\frac{k_{z}^{L, i}}{k_{z}^{L, i}} \frac{C}{A_{0}}=\tan \theta_{L i} \frac{C}{A_{0}},
$$

and for the $x$-component

$$
R_{P \rightarrow S V}^{X}=\left.\frac{-\partial A / \partial z}{\partial \varphi^{i n} / \partial x}\right|_{z=0}=-\frac{k_{z}^{T, r}}{k_{z}^{L, i}} \frac{C}{A_{0}}=\tan \theta_{\operatorname{tr}} \frac{C}{A_{0}} .
$$

In addition, if one deals with displacement amplitudes which are angle independent, the way to find the reflection coefficients is to replace $A_{0}$. B, and C by

$$
A_{0} \rightarrow k_{L} A_{0} \quad, \quad B \rightarrow k_{L} B \quad, \quad C \rightarrow k_{T} C
$$

which define the reflection coefficients

$$
\bar{R}_{P \rightarrow S V}=\frac{\mathbf{k}_{L} B}{\bar{k}_{L} A_{0}}=\frac{B}{A_{0}} .
$$

and

$$
\bar{R}_{P \rightarrow S V}=\frac{k_{L} c}{k_{L} A_{0}}=\frac{c_{L}}{c_{T}} \frac{c}{A_{0}} .
$$

In addition Aki and Richards (Ref. 3.3) use a convention whereby

$$
\bar{R}_{P \rightarrow S V}=-\frac{C_{L}}{C_{T}} \frac{C}{A_{0}},
$$

presumably because the $z$-component of displacement of the $S V$ wave on the boundary is in the negative z-direction.

## DISPLACBURNTS

Since we have calculated the reflection coefficients for this zerotraction surface it is useful to evaluate the displacements on the boundary al so. Since $v=0, a_{2}=0$ identically. The other displacements can be found from (4.32) and (4.33) to be

$$
\begin{equation*}
u_{1}(x, 0)=\frac{\partial \phi}{\partial x}(x, 0)-\frac{\partial A}{\partial z}(x, 0) \text {. } \tag{6.36}
\end{equation*}
$$

and

$$
\begin{equation*}
u_{3}(x, 0)=\frac{\partial \phi}{\partial z}(x, 0)+\frac{\partial A}{\partial x}(x, 0) . \tag{6.37}
\end{equation*}
$$

Substituting (6.4) and (6.5) into (6.36) and (6.37) we get

$$
\begin{equation*}
u_{1}(x, 0)=i A_{0}\left[k_{x}^{L, i}\left(1+\frac{B}{A_{0}}\right)-k_{z}^{T, r} \frac{C}{A_{0}}\right] e^{i k_{x}^{L, i} x} \tag{6.38}
\end{equation*}
$$

and

$$
\begin{equation*}
u_{3}(x, 0)=i A_{0}\left[k_{z}^{L, i}\left(1-\frac{B}{A_{0}}\right)-k_{x}^{T, r} \frac{C}{A_{0}}\right] e^{i k_{z}^{L, i} x} \tag{6.39}
\end{equation*}
$$

which can be further evaluated using (6.29), (6.31), (6.32) or (6.33).

## $\xrightarrow{\text { ZRROES OR RP }} \rightarrow \mathbf{S V}$

There is no conversion of $P$ to $S V$ waves when $R_{P \rightarrow S V}$ vanishes. From (6.32) or (6.33) this can be seem to occur for three cases. The first is when $p=0$ or $\theta_{L i}=0$, i.e. for normal incidence. The second is when $p=1 / c_{L}$ or $\theta_{L_{i}}=\pi / 2$, i.e. for grazing incidence, and the third is when $p=2^{-1 / 2} C_{T}^{-1}$ or $\boldsymbol{\theta}_{\mathrm{Tr}_{r}}=\pi / 4$. In each case energy is conserved since $\mathrm{R}_{\mathrm{P} \rightarrow \mathrm{P}}$ equals $-1,-1$, and +1 for the respective cases.

### 3.7 SV-WAVB INCIDENCE ON A FRBE SURFACE

The additional coupled case to the problem in Sec. 6 is that of SV-wave incidence on a flat, free surface. The scattered field has both $P$ and SV components as illustrated in Fig. 3.5.


Fig. 3.5

The potentials $\phi$ and $A$ satisfy the same boundary conditions as (6.2) and (6.3) but are here given by

$$
\begin{equation*}
\phi(x, z)=B \exp \left[i\left(k_{z}^{L}, r_{z}+k_{z}^{L}, r_{z}\right)\right] \text {. } \tag{7.1}
\end{equation*}
$$

which contains only a scattered field, and

$$
\begin{align*}
A(x, z)= & A_{0} \exp \left[i \left(k_{z}^{T, i_{x}}-k_{z}^{\left.\left.T, i_{z}\right)\right]}\right.\right. \\
& +C \exp \left[i \left(k_{z}^{T, r}+k_{z}^{\left.\left.T, r_{z}\right)\right]}\right.\right. \tag{7.2}
\end{align*}
$$

containing both incident ( $A_{0}$ ) and scattered (C) fields. Substituting these results into (6.2) and (6.3) and again using the translational invariance in $x$ given by (6.14) we get the set of equations for $R_{1}=B / A_{0}=R_{S V} \rightarrow P$ and $R_{2}=C / A_{0}=R_{S V} \rightarrow S V$ given by

$$
\begin{align*}
& a_{11} R_{2}-a_{12} R_{1}=-a_{11},  \tag{7.3}\\
& a_{21} R_{2}+a_{22} R_{1}=a_{21}, \tag{7.4}
\end{align*}
$$

Where the $a_{i j}$ are given by (6.10)-(6.13). Note the form of the equations (7.3) and (7.4) is the same as (6.18) and (6.19) except for the new interpretations of $R_{1}$ and $R_{2}$ and the fact that the right hand sides of the two equations are different. The latter occurs because the incident wave has changed from $P$ in Sec. 6 to $S V$ here.

The equations have the solutions

$$
\begin{equation*}
R_{S V \rightarrow S V}=\left(a_{12} a_{21}-a_{112} a_{22}\right) / \Delta \tag{7.5}
\end{equation*}
$$

and

$$
\begin{equation*}
R_{S V \rightarrow P}=2 a_{21} a_{21} / \Delta \tag{7.6}
\end{equation*}
$$

where $\Delta$ is the same denominator as in Sec. 6, given by (6.22). Osing the properties of the $a_{i j}$ given in (6.25)-(6.28) we get

$$
\begin{equation*}
R_{S V \rightarrow S V}=\frac{4 p^{2} c_{T}^{3}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-p^{2} c_{T}^{2}\right)^{1 / 2}-c_{L}\left(1-2 p^{2} c_{T}^{2}\right)^{2}}{4 p^{2} c_{T}^{3}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-p^{2} c_{T}^{2}\right)^{1 / 2}+c_{L}\left(1-p^{2} c_{T}^{2}\right)^{2}} \tag{7.7}
\end{equation*}
$$

which appears similar to the result in (6.29) but here note that

$$
\begin{equation*}
\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}=\cos \theta_{L r} . \tag{7.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(1-p^{2} c_{T}^{2}\right)^{1 / 2}=\cos \theta_{T i} \tag{7.9}
\end{equation*}
$$

Osing these results we can also write (7.7) as

$$
\begin{equation*}
R_{S V \rightarrow S V}=\frac{4 p^{2}\left(\cos \theta_{L r} / c_{L}\right)\left(\cos \theta_{T i} / c_{T}\right)-\left(c_{T}^{-2}-2 p^{2}\right)^{2}}{4 p^{2}\left(\cos \theta_{L r} / c_{L}\right)\left(\cos \theta_{T i} / c_{T}\right)+\left(c_{T}^{-2}-2 p^{2}\right)^{2}} \tag{7.10}
\end{equation*}
$$

We can also wite the other reflection coefficient as

$$
\begin{equation*}
R_{S V \rightarrow P}=\frac{4 p c_{L} c_{T}\left(1-p^{2} c_{T}^{2}\right)^{1 / 2}\left(1-2 p^{2} c_{T}^{2}\right)}{4 p^{2} c_{T}^{3}\left(1-p^{2} c_{L}^{2}\right)^{1 / 2}\left(1-p^{2} c_{T}^{2}\right)^{1 / 2}+c_{L}\left(1-2 p^{2} c_{T}^{2}\right)^{2}}, \tag{7.11}
\end{equation*}
$$

or as

$$
\begin{equation*}
R_{S V \rightarrow P}=\frac{4 p\left(\cos \theta_{T i} / c_{T}\right)\left(c_{T}^{-2}-2 p^{2}\right)}{4 p^{2}\left(\cos \theta_{L r} / c_{L}\right)\left(\cos \theta_{T i} / c_{T}\right)+\left(c_{T}^{-2}-2 p^{2}\right)^{2}} . \tag{7.12}
\end{equation*}
$$

Note that $\mathbf{R}_{\mathrm{SV} \rightarrow \mathrm{P}}$ vanishes when
(a) $p=0 \quad\left(\theta_{T i}=0\right.$, normal incidence)
(b) $p=c_{T}^{-1}\left(\theta_{T i}=\pi / 2\right.$, grazing incidence $)$
and
(c) $p=2^{-x / 2} c_{T}^{-1}\left(\theta_{T i}=\pi / 4\right)$,
and in these cases we get no conversion of $S V$ to $P$ waves. Again al so note, as in Sec. 6., the se are reflection coefficients for the potentials.

In addition we can al so compute the values of displacement on the $z=0$ surface from (6.35) and (6.36). These are given by

$$
\begin{equation*}
u_{1}(x, 0)=i A_{0}\left[k_{x}^{L, r} \frac{B}{A_{0}}+k_{z}^{T, i}\left(1-\frac{C}{A_{0}}\right)\right] e^{i k_{x}^{x}} . \tag{7.13}
\end{equation*}
$$

and

$$
\begin{equation*}
u_{3}(x, 0)=i A_{0}\left[\frac{k_{z}, r}{} \frac{B}{A_{0}}+k_{x}^{T, i}\left(1+\frac{C}{A_{0}}\right)\right] e^{i k_{x} x} . \tag{7.14}
\end{equation*}
$$

where $k_{x}$ stands for any of the $x$ components of wave number.

### 3.8 SA-WAVE INCIDENCR

We consider the three possible cases of SH-wave incidence on a flat boundary. As we have already noted for this type of problem, the $S H$ waves decouple from the $P$ - $S V$ waves. The first two cases are illustrated in Fig. 3.6 below.


Fig. 3.6

Here the 1 ines indicate the direction of propagation of the waves, and the circles with dots (the tips of arrows pointing out from the paper) indicate the direction of the displacements of the waves.

## CASB (a): Free Flat Surface

For this case the displacement is
written in terms of incident (amplitude A) and reflected (amplitude B) waves. For both cases we have that

$$
\begin{equation*}
\left(k_{x}^{T, i}\right)^{2}+\left(k_{z}^{T, i}\right)^{2}=\left(k_{x}^{T, r}\right)^{2}+\left(k_{z}^{T, r}\right)^{2}=k_{T}^{2} . \tag{8.2}
\end{equation*}
$$

The boundary condition at $z=0$ is

$$
\begin{equation*}
\tau_{z 2}=\mu \frac{\partial v}{\partial z}=0 \text {. } \tag{8.3}
\end{equation*}
$$

Since this must remain invariant for any value of $x$

$$
\begin{equation*}
k_{x}^{T, i}=k^{T} \sin \theta_{i}=k_{x}^{T, r}=k^{T} \sin \theta_{r}, \tag{8.4}
\end{equation*}
$$

so that the angle of reflection equals the angle of incidence. Further to satisfy the boundary condition, we require $B / A_{0}=1$.

## CASB (b): Rigid F1at Surface

For this case the displacement is again given by (8.1), and (8.2) is satisfied. The boundary condition is now

$$
\begin{equation*}
v(x, 0)=0 \tag{8.5}
\end{equation*}
$$

Again we get $\theta_{i}=\theta_{r}$ but now $B / A_{0}=-1$.

## CASB (c): Blastic Interface

For this case we have a flat surface separating two elastic media, (1), with elastic parameters $\mu_{1}, \lambda_{1}, k_{\mathrm{T}_{1}}, \mathrm{k}_{\mathrm{L} 1}, \rho_{1}$, and (2), with parameters $\mu_{2}, \lambda_{2}$, ${ }^{k_{T} 2}$, $k_{L_{2}}$, and $\rho_{2}$. Part of this case corresponding to incidence in region
(1) is illustrated in Fig. 3.7 below.


Fig. 3.7

The angles $\boldsymbol{\theta}_{\mathrm{ij}}$ correspond to scattering from region $i$ to region $j$. The displacements for both regions can be written as

$$
\begin{equation*}
v_{1}(x, z)=A_{0} \exp \left[i\left(k_{x}^{T, i} i_{z}-k_{z}^{T, i} z\right)\right]+B \exp \left[i \left(k_{x}^{T, r} x+k_{z}^{\left.\left.T, r_{z}\right)\right] .}\right.\right. \tag{8.6}
\end{equation*}
$$

and

$$
\begin{equation*}
v_{2}(x, z)=c \exp \left[i\left(k_{x}^{T, t 2} x-k_{z}^{T, t 2 z}\right)\right] . \tag{8.7}
\end{equation*}
$$

in terms of incident (i) and reflected (r) fields in region (1), and transmitted (tr) fieldin region (2). We have the dispersion relations

$$
\begin{equation*}
\left(k_{x}^{T, i}\right)^{2}+\left(k_{z}^{T, i}\right)^{2}=\left(k_{x}^{T, r}\right)^{2}+\left(k_{z}^{T, r}\right)^{2}=\left(k_{T 1}\right)^{2} . \tag{8.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(k_{x}^{T, t 2}\right)^{2}+\left(k_{z}^{T, t 2}\right)^{2}=\left(k_{T_{2}}\right)^{2} \tag{8.9}
\end{equation*}
$$

The continuity conditions at $z=0$ in terms of stress and displacement are

$$
\begin{equation*}
\tau_{z^{2}}^{(1)}(x, 0)=\tau_{z^{2}}^{(2)}(x, 0) \tag{8.10}
\end{equation*}
$$

and

$$
\begin{equation*}
u_{2}^{(1)}(x, 0)=u_{2}^{(2)}(x, 0) \tag{8.11}
\end{equation*}
$$

Written in terms of $v$ these become respectively

$$
\begin{equation*}
\mu_{1} \frac{\partial v_{1}}{\partial z}(x, 0)=\mu_{2} \frac{\partial v_{2}}{\partial z}(x, 0) \tag{8.12}
\end{equation*}
$$

and

$$
\begin{equation*}
v_{1}(x, 0)=v_{2}(x, 0) \tag{8.13}
\end{equation*}
$$

Substituting (8.6) and (8.7) in (8.12), and using the translational invariance in given by

$$
\begin{equation*}
k_{x}^{T, i}=k_{x}^{T, I}=k_{x}^{T, t^{2}} \tag{8.14}
\end{equation*}
$$

which imply

$$
\begin{equation*}
\theta_{i}=\theta_{11}, k_{T 1} \sin \theta_{11}=k_{T_{2}} \sin \theta_{12}, \tag{8.15}
\end{equation*}
$$

(angle of incidence equals angle of reflection, and Snell's Law) yields the result

$$
\begin{equation*}
A-B=a C \tag{8.16}
\end{equation*}
$$

where

$$
\begin{equation*}
a=\mu_{2} k_{z}^{T, t^{2} / \mu_{1}} k_{z}^{T, i} \tag{8.17}
\end{equation*}
$$

with $k_{z}^{T, i}=k_{z}^{T, r}$ from (8.14). Osing the results

$$
\begin{equation*}
k_{z}^{T, t_{2}}=k_{T_{2}} \cos \theta_{12} ; k_{z}^{T, i}=k_{T 1} \cos \theta_{11} \tag{8.18}
\end{equation*}
$$

and the definitions

$$
\begin{equation*}
\mu=\mu_{2} / \mu_{1} \quad, \quad K=k_{t 2} / k_{t 1} \tag{8.19}
\end{equation*}
$$

we can write $\alpha$, using Sne11's Law as

$$
\begin{equation*}
a=\mu\left(K^{2}-\sin ^{2} \theta_{11}\right)^{1 / 2} / \cos \theta_{11} \tag{8.20}
\end{equation*}
$$

Substituting (8.6) and (8.7) in the second equation, (8.13), yields the result

$$
\begin{equation*}
\mathbf{A}+\mathbf{B}=\mathbf{C} \tag{8.21}
\end{equation*}
$$

where we have used (8.14). Simultaneous solution of (8.16) and (8.21) yields two components of our scattering matrix $S_{i j}, S_{11}=B / A_{0}$, the refiection coefficient from region (1) to region (1'), and $S_{12}=C / A_{0}$, the scattering (transmission) coefficient from region (1) to region (2). They are

$$
\begin{equation*}
S_{11}=\frac{\cos \theta_{11}-\mu\left(K^{2}-\sin ^{2} \theta_{11}\right)^{1 / 2}}{\cos \theta_{11}+\mu\left(K^{2}-\sin ^{2} \theta_{11}\right)^{1 / 2}}, \tag{8.22}
\end{equation*}
$$

and

$$
\begin{equation*}
S_{12}=\frac{2 \cos \theta_{11}}{\cos \theta_{11}+\mu\left(K^{2}-\sin ^{2} \theta_{11}\right)^{1 / 2}} . \tag{8.23}
\end{equation*}
$$

Osing Snel1's Law these can al so be written as

$$
\begin{equation*}
S_{11}=\left(\cos \theta_{11}-\mu K \cos \theta_{12}\right)\left(\cos \theta_{11}+\mu K \cos \theta_{12}\right)^{-1} \tag{8.24}
\end{equation*}
$$

and

$$
\begin{equation*}
S_{12}=2 \cos \theta_{11}\left(\cos \theta_{11}+\mu K \cos \theta_{12}\right)^{-1} . \tag{8.25}
\end{equation*}
$$

or, expressing everything in terms of $\theta_{12}$ as

$$
\begin{equation*}
S_{11}=\frac{\left(1-K^{2} \sin ^{2} \theta_{12}\right)^{1 / 2}-\mu K \cos \theta_{12}}{\left(1-K^{2} \sin ^{2} \theta_{12}\right)^{1 / 2}+\mu K \cos \theta_{12}}, \tag{8.26}
\end{equation*}
$$

and

$$
\begin{equation*}
S_{12}=\frac{2\left(1-K^{2} \sin ^{2} \theta_{12}\right)^{1 / 2}}{\left(1-K^{2} \sin ^{2} \theta_{12}\right)^{1 / 2}+\mu K \cos \theta_{12}} . \tag{8.27}
\end{equation*}
$$

The other two components of the scattering matrix $S_{22}$ and $S_{21}$ correspond to incidence from region (2) as illustrated in Fig. 3.8 below.


Fig. 3.8

We can find the se components from (8.26) and (8.27) using the following interchange

$$
\begin{equation*}
\mu_{1} \leftrightarrow \mu_{2}, \quad \mathbf{k}_{1} \leftrightarrow \mathbf{k}_{2}, \quad \theta_{12} \rightarrow \theta_{21}, \tag{8.28}
\end{equation*}
$$

or equivalently

$$
\begin{equation*}
\mu \rightarrow 1 / \mu, K \rightarrow 1 / K, \theta_{12} \rightarrow \theta_{21} \tag{8.29}
\end{equation*}
$$

Using (8.29) in (8.26) we find $S_{22}$ given by

$$
\begin{equation*}
S_{22}=\frac{\mu\left(K^{2}-\sin ^{2} \theta_{21}\right)^{1 / 2}-\cos \theta_{21}}{\mu\left(K^{2}-\sin ^{2} \theta_{21}\right)^{1 / 2}+\cos \theta_{21}}, \tag{8.30}
\end{equation*}
$$

and using (8.29) in (8.27) we find

$$
\begin{equation*}
S_{21}=\frac{2 \mu\left(K^{2}-\sin ^{2} \theta_{21}\right)^{1 / 2}}{\mu\left(K^{2}-\sin ^{2} \theta_{21}\right)^{1 / 2}+\cos \theta_{21}} . \tag{8.31}
\end{equation*}
$$

By Sne11's Law, we have the same set of angles in Figs. (3.7) and (3.8), so $\theta_{21}=\theta_{11}$. Comparing (8.30) with (8.22) we see that

$$
\begin{equation*}
S_{22}=-S_{11} \tag{8.32}
\end{equation*}
$$

Other representations for $S_{22}$ and $S_{21}$ can be found by doing this $2 \leftrightarrow 1$ interchange in the pairs of equations (8.24) and (8.25), and in (8.22) and (8.23).

### 3.9 ELASTIC INTBRFACE (RIGID CONTACT)

In Sec. 8, Case (c), we treated the case of an SH-wave incident on a flat surface separating two elastic media with different elastic parameters and in rigid contact. Here we treat the same geometry for the remaining two cases, those of $P$-wave incidence and SV-wave incidence.

## CASB (a): P-wave Incidence

The geometry for this case is illustrated in Fig. 3.9 below.


Fig. 3.9

We have a p-wave incident at angle $\theta_{i}$, reflected and transmitted $P$-waves at angles $\theta_{1}$ and $\theta_{2}$ respectively, and reflected and transmitted SV-waves at angles $\boldsymbol{\theta}_{\mathrm{T} 1}$ and $\boldsymbol{\theta}_{\mathrm{T} 2}$ respectively. Our potentials can be written in Region 1 as:

$$
\begin{equation*}
\phi^{(1)}(x, z)=\phi^{\text {in }}(x, z)+\phi^{S C}(x, z), \tag{9.1}
\end{equation*}
$$

where

$$
\begin{equation*}
\phi^{i n}(x, z)=A_{0}^{P} \exp \left[i\left(k_{x}^{L, i} i_{z}^{L}, i_{z}\right)\right] \tag{9.2}
\end{equation*}
$$

and

$$
\begin{equation*}
\phi_{\phi}^{S C}(x, z)=B \exp \left[i\left(k_{z}^{L}, r_{x}+k_{z}^{L}, r_{z}\right)\right] \tag{9.3}
\end{equation*}
$$

and for the $A$ potential

$$
\begin{equation*}
A^{(1)}(x, z)=C \exp \left[i\left(k_{x}^{T x} x+k_{z}^{T i} z\right)\right] \tag{9.4}
\end{equation*}
$$

Also in Region 1 the displacement and stress components are written on the $z=0$ surface as

$$
\begin{align*}
& u_{1}^{(1)}(x, 0)=\frac{\partial \phi^{S C}}{\partial x}(x, 0)-\frac{\partial A^{(x)}}{\partial z}(x, 0)+\frac{\partial \phi^{\text {in }}}{\partial x}(x, 0),  \tag{9.5}\\
& u_{3}^{(1)}(x, 0)=\frac{\partial \phi^{S C}}{\partial z}(x, 0)+\frac{\partial A^{(1)}}{\partial x}(x, 0)+\frac{\partial \phi^{\text {in }}}{\partial z}(x, 0) \tag{9.6}
\end{align*}
$$

$$
\tau_{z 1}^{(1)}(x, 0)=\mu_{1}\left[\frac{\partial^{2} A^{(1)}}{\partial x^{2}}(x, 0)-\frac{\partial^{2} A^{(1)}}{\partial z^{2}}(x, 0)+2 \frac{\partial^{2} \phi C}{\partial x \partial z}(x, 0)\right]
$$

$$
\begin{equation*}
+2 \mu_{1} \frac{\partial^{2} \phi^{i n}}{\partial x \partial z}(x, 0) \tag{9.7}
\end{equation*}
$$

and

$$
\begin{align*}
\tau_{z^{3}}^{(1)}(x, 0)=2 \mu_{1} & {\left[\frac{\partial^{2} \phi S C}{\partial z^{2}}(x, 0)+\frac{\partial^{2} A^{(1)}}{\partial x \partial z}(x, 0)\right]-\lambda_{1} k_{L 1}^{2} \phi^{S C}(x, 0) } \\
& +2 \mu_{1} \frac{\partial^{2} \phi^{i n}}{\partial z^{2}}(x, 0)-\lambda_{1} k_{L 1}^{2} \phi^{i n}(x, 0) \tag{9.8}
\end{align*}
$$

In Region 2 the scalar potentials for the transmitted $P$ and $S V$ waves are

$$
\begin{equation*}
\phi^{(z)}(x, z)=D \exp \left[i\left(k_{x}^{L z} x-k_{z}^{L / z}\right)\right] . \tag{9.9}
\end{equation*}
$$

and

$$
\begin{equation*}
A^{(2)}(x, z)=E \exp \left[i\left(k_{x}^{T} x-k_{z}^{T 2} z\right)\right] \tag{9.10}
\end{equation*}
$$

Simarly the displacement and stress components expressed in terms of these potentials are given by (on the $z=0$ surface)

$$
\begin{align*}
u_{1}^{(2)}(x, 0)= & \frac{\partial}{\partial x} \phi^{(2)}(x, 0)-\frac{\partial}{\partial z} A^{(2)}(x, 0),  \tag{9.11}\\
u_{z}^{(2)}(x, 0)= & \frac{\partial}{\partial z} \phi^{(2)}(x, 0)+\frac{\partial}{\partial x} A^{(2)}(x, 0),  \tag{9.12}\\
\tau_{z x}^{(2)}(x, 0)= & \mu_{2}\left[\frac{\partial^{2}}{\partial x^{2}} A^{(2)}(x, 0)-\frac{\partial^{2}}{\partial z^{2}} A^{(2)}(x, 0)\right] \\
& +2 \mu_{2} \frac{\partial^{2}}{\partial x \partial z} \phi^{(2)}(x, 0) . \tag{9.13}
\end{align*}
$$

and

$$
\left.\begin{array}{rl}
\tau_{z_{3}}^{(2)}(x, 0)= & 2 \mu_{2}
\end{array} \frac{\partial^{2}}{\partial z^{2}} \phi^{(2)}(x, 0)+\frac{\partial^{2}}{\partial x \partial z} A^{(2)}(x, 0)\right] .
$$

The boundary conditions for two elastic media in rigid contact are the continuity of displacements and stresses at the $z=0$ interface. The conditions are given by

$$
\begin{align*}
& u_{1}^{(1)}(x, 0)=u_{1}^{(2)}(x, 0)  \tag{9.15}\\
& u_{3}^{(1)}(x, 0)=u_{3}^{(2)}(x, 0),  \tag{9.16}\\
& \tau_{z 1}^{(1)}(x, 0)=\tau_{21}^{(2)}(x, 0) . \tag{9.17}
\end{align*}
$$

and

$$
\begin{equation*}
\tau_{z^{3}}^{(1)}(x, 0)=\tau_{z^{3}}^{(2)}(x, 0) \tag{9.18}
\end{equation*}
$$

Substituting (9.5)-(9.8) and (9.11)-(9.14) into (9.15)-(9.18) and writing the results in terms of the incident field on the right hand side we get (all fields are evaluated at ( $x, 0$ ) )

$$
\begin{align*}
& \frac{\partial \phi^{S}}{\partial x}-\frac{\partial A}{\partial z}^{(1)}-\frac{\partial \phi}{\partial x}^{(2)}+{\frac{\partial A^{z}}{\partial z}}^{(2)}=-{\frac{\partial \phi^{\prime}}{\partial x}}^{\text {in }}  \tag{9.19}\\
& \frac{\partial \phi^{2}}{\partial C}+\frac{\partial A}{\partial x}^{(1)}-\frac{\partial \phi}{\partial z}^{(2)}-\frac{\partial A^{(2)}}{\partial x}=-{\frac{\partial \phi^{z}}{\partial z}}^{\text {in }} . \tag{9.20}
\end{align*}
$$

$$
2 \mu_{1} \frac{\partial^{2} \phi^{S C}}{\partial x \partial z}+\mu_{1}\left[\frac{\partial^{2} A^{(1)}}{\partial x^{2}}-\frac{\partial^{2} A^{(1)}}{\partial z^{2}}\right]
$$

$$
\begin{equation*}
-2 \mu_{2} \frac{\partial^{2} \phi}{\partial x \partial z}{ }^{(2)}-\mu_{2}\left[{\frac{\partial^{2} A}{\partial x^{2}}}^{(2)}-\frac{\partial^{2} A^{(2)}}{\partial z^{2}}\right]=-2 \mu_{1} \frac{\partial^{2} \phi^{i n}}{\partial x^{\partial} z} \tag{9.21}
\end{equation*}
$$

and

$$
\begin{align*}
& 2 \mu_{1} \frac{\partial^{2} \phi}{\partial z^{2}}-\lambda_{1} k_{L 1}^{2} \phi^{S C}+2 \mu_{1} \frac{\partial^{2} A}{\partial x \partial z} \\
& -2 \mu_{2}{\frac{\partial^{2} \phi}{\partial z^{2}}}^{(2)}+\lambda_{2} k_{L_{2}}^{2} \phi^{(2)}-2 \mu_{2} \frac{\partial^{2} A}{\partial x \partial z} \\
&  \tag{9.22}\\
& =-2 \mu_{1} \frac{\partial^{2} \phi^{i n}}{\partial z^{2}}+\lambda_{1} k_{L_{1}}^{2} \phi^{i n}
\end{align*}
$$

All the $z$-components of the phases in each of the waves vanish since we evaluate at $z=0$. By translational invariance all the components of the wave numbers are equal. Expressing these in terms of the angles in Fig. 9.1 we have

$$
\begin{gather*}
k_{x}^{L, i}=k_{L_{1}} \sin \theta_{i}, \quad k_{x}^{L, r}=k_{L_{1}} \sin \theta_{1}, \quad k_{I}^{T 1}=k_{T_{1}} \sin \theta_{T 1} \\
k_{X_{1}}^{L_{2}}=k_{L_{2}} \sin \theta_{2}, k_{x}^{T 2}=k_{T_{2}} \sin \theta_{T 2} . \tag{9.23}
\end{gather*}
$$

Equating these we get the 1 aws of reflection $\quad\left(\theta_{1}=\theta_{i}\right)$ for $P$-waves and $k_{L^{1}} \sin \theta_{i}=k_{T 1} \sin \theta_{T 1}$ for SV-waves, and the 1 aws of refraction for $P$-waves $\left(k_{L_{1}} \sin \theta_{i}=k_{L_{2}} \sin \theta_{2}\right)$ and $S V$-waves $\left(k_{T_{1}} \sin \theta_{T_{2}}=k_{T_{2}} \sin \theta_{T_{2}}\right)$. In addition we could express all these equal $k_{x}$-components in terms of the ray parameter, i.e. $k_{x}=\omega p$. Further the z-components of wave number can be writtenas

$$
\begin{align*}
& \mathbf{k}_{\mathrm{z}}^{\mathrm{L}, \mathbf{i}}=\mathbf{k}_{\mathrm{z}}^{\mathrm{L}, \mathbf{r}}=\mathbf{K}_{\mathrm{L}_{1}}=\mathbf{k}_{\mathrm{L}_{1}} \cos \theta_{i} .  \tag{9.24}\\
& \mathbf{k}_{\mathrm{Z}} \mathrm{~T}_{1}=\mathrm{K}_{\mathrm{T}_{1}}=\mathbf{k}_{\mathrm{T}_{1}} \cos \theta_{\mathrm{T}_{1}} \text {. }  \tag{9.25}\\
& \mathbf{k}_{\mathbf{L}}^{L_{2}}=\mathbf{K}_{L_{2}}=\mathbf{k}_{L_{2}} \cos \theta_{2} \text {, } \tag{9.26}
\end{align*}
$$

and

$$
\begin{equation*}
\mathbf{k}_{\mathrm{Z}}^{\mathbf{T}_{2}}=\mathbf{K}_{\mathrm{T}_{2}}=\mathbf{k}_{\mathrm{T}_{2}} \cos \theta_{\mathrm{T}_{2}} \tag{9.27}
\end{equation*}
$$

Substituting the wave forms and the above results into (9.19)-(9.22) we get the 1 inear equations (Zoeppritz Equations)

$$
\begin{align*}
& m_{11} S_{1}+m_{12} S_{2}+m_{13} T_{1}+m_{14} T_{2}=P_{1},  \tag{9.28}\\
& m_{21} S_{1}+m_{22} S_{2}+m_{23} T_{1}+m_{24} T_{2}=P_{2},  \tag{9.29}\\
& m_{31} S_{1}+m_{32} S_{2}+m_{33} T_{1}+m_{34} T_{2}=P_{3},  \tag{9.30}\\
& m_{41} S_{1}+m_{42} S_{2}+m_{43} T_{1}+m_{44} T_{2}=P_{4}, \tag{9.31}
\end{align*}
$$

in terms of the two components of the scattering and transmission matrices

$$
\begin{align*}
& \mathbf{S}_{\mathbf{1}}=\mathbf{B} / \mathbf{A}_{0}^{\mathbf{P}}, \quad \mathbf{S}_{2}=\mathbf{C} / \mathbf{A}_{0}^{\mathbf{P}}  \tag{9.32}\\
& \mathbf{T}_{1}=\mathbf{D} / \mathbf{A}_{0}^{\mathbf{P}}, \quad \mathbf{T}_{2}=\mathbf{E} / \mathbf{A}_{0}^{\mathbf{P}} \tag{9.33}
\end{align*}
$$

and the four tems $P_{j}$ related to the incident P-wave field occurring on the right hand sides of (9.19)-(9.22). The matrix elements $m_{i j}$ are given by

$$
\left[\begin{array}{llll}
i k_{x} & -i K_{T 1} & -i k_{x} & -i K_{T 2} \\
i K_{L 1} & i k_{x} & i K_{L_{2}} & -i k_{x} \\
-2 \mu_{1} k_{I} K_{L 1} & \mu_{1}\left(K_{T 1}^{2}-k_{x}^{2}\right) & -2 \mu_{2} k_{X} K_{L 2} & \mu_{2}\left(k_{x}^{2}-K_{T 2}^{2}\right) \\
-2 \mu_{1} K_{L 1}^{2}-\lambda_{1} k_{L 1}^{2} & -2 \mu_{1} k_{Y} K_{T 1} & 2 \mu_{2} K_{L 2}^{2}+\lambda_{2} k_{L 2}^{2} & -2 \mu_{2} k_{X} K_{T 2}
\end{array}\right]
$$

and the elements for the incident field are

$$
\left[\begin{array}{l}
P_{1}  \tag{9.35}\\
P_{2} \\
P_{3} \\
P_{4}
\end{array}\right]=\left[\begin{array}{l}
-m_{11} \\
m_{21} \\
m_{31} \\
-m_{41}
\end{array}\right]=\left[\begin{array}{l}
-i k_{x} \\
i K_{L 1} \\
-2 \mu_{1} k_{x} K_{L 1} \\
2 \mu_{1} K_{L 1}^{2}+\lambda_{1} k_{L 1}^{2}
\end{array}\right] \text {. }
$$

Note how the incident field elements are related to the first col umn of the m-matrix, i.e. that colun corresponding to the $P$-waves.

The solution is found once the inverse of the matrix $m$ is known, i.e.

$$
\left[\begin{array}{l}
S_{1}  \tag{9.36}\\
S_{2} \\
T_{1} \\
T_{2}
\end{array}\right]=m^{-1}\left[\begin{array}{l}
P_{1} \\
P_{2} \\
P_{3} \\
P_{4}
\end{array}\right]
$$

The matrix elements of the inverse can be writen as

$$
\begin{equation*}
\left[m^{-1}\right]_{i j}=D_{j i} / \Delta \tag{9.37}
\end{equation*}
$$

where the $D_{j}$ are the cofactors of the matrix $m$ and $\Delta$ is its determinant. For example, the reflection coefficients can be explicitly witten as

$$
\begin{equation*}
S_{1}=\Delta^{-1} \sum_{j=1}^{4} D_{j^{1}} P_{j} \tag{9.38}
\end{equation*}
$$

and

$$
\begin{equation*}
S_{2}=\Delta^{-1} \sum_{j=1}^{4} D_{j 2} P_{j} \tag{9.39}
\end{equation*}
$$

There are thas a total of eight cofactors to evaluate, and the determinant can be written in terms of these as

$$
\begin{equation*}
\Delta=m_{11} D_{11}+m_{21} D_{21}+m_{31} D_{31}+m_{41} D_{41}, \tag{9.40}
\end{equation*}
$$

as an expansion in the first colum of $m$. The specific form of the matrix elements in (9.34) enable us to write

$$
m=\left[\begin{array}{llll}
m_{11} & m_{12} & m_{11} & m_{14}  \tag{9.41}\\
m_{21} & m_{11} & m_{23} & -m_{11} \\
m_{31} & m_{32} & m_{33} & m_{34} \\
m_{41} & m_{42} & m_{43} & m_{44}
\end{array}\right]
$$

The cofactors are

$$
\begin{align*}
& D_{11}=m_{11}\left[m_{33} m_{44}+m_{33} m_{42}-m_{43} m_{32}-m_{43} m_{34}\right] \\
& +m_{23}\left[m_{42} m_{34}-m_{44} m_{32}\right] \quad .  \tag{9.42}\\
& D_{21}=m_{12}\left[m_{34} m_{43}-m_{33} m_{44}\right]+m_{11}\left[m_{34} m_{42}-m_{32} m_{44}\right] \\
& +m_{14}\left[m_{33} m_{42}-m_{32} m_{43}\right] \text {. }  \tag{9.43}\\
& D_{31}=m_{11}^{2}\left[m_{42}+m_{44}\right]+m_{11}\left[m_{43} m_{14}+m_{42} m_{12}\right]+m_{23}\left[m_{12} m_{44}-m_{14} m_{42}\right]  \tag{9.44}\\
& D_{41}=-m_{11}^{2}\left[m_{32}+m_{34}\right]-m_{11}\left[m_{14}+m_{12}\right] m_{33}+m_{23}\left[m_{14} m_{32}-m_{22} m_{34}\right],  \tag{9.45}\\
& D_{12}=m_{11}\left[m_{31} m_{43}-m_{33} m_{41}\right]+m_{21}\left[m_{34} m_{43}-m_{33} m_{44}\right] \\
& +m_{23}\left[m_{44} m_{31}-m_{41} m_{34}\right] \quad .  \tag{9.46}\\
& D_{22}=m_{11}\left[m_{33} m_{44}+m_{31} m_{44}-m_{34} m_{43}-m_{34} m_{41}\right] \\
& +m_{14}\left[m_{31} m_{43}-m_{33} m_{41}\right] \text {. }  \tag{9.47}\\
& D_{32}=m_{14}\left[m_{23} m_{41}-m_{21} m_{43}\right] \\
& -m_{11}\left[m_{23} m_{44}+m_{11} m_{43}+m_{21} m_{44}+m_{11} m_{41}\right] \quad, \tag{9.48}
\end{align*}
$$

and

$$
\begin{align*}
D_{42}=m_{14} & {\left[m_{21} m_{33}-m_{23} m_{31}\right] } \\
& +m_{11}\left[m_{23} m_{34}+m_{21} m_{34}+m_{11} m_{33}+m_{11} m_{31}\right] \quad . \tag{9.49}
\end{align*}
$$

and these can be explicitly evaluated using (9.34).

CASB (b): SV-wave incidence
The geometry for this case is illustrated in Fig. 3.10.


Fig. 3.10

We have an $S V$-wave incident at angle $\theta_{i}$, reflected and transmitted p-waves at angles $\theta_{1}$ and $\theta_{2}$ respectively, and reflected and transmitted SV-waves at angles $\theta^{\boldsymbol{T} 1}$ and $\theta_{\mathrm{T}_{2}}$ respectively. We retain much of the notation of CASE (a). The potentials in Region 1 are

$$
\begin{equation*}
g^{(1)}(x, z)=B \exp \left[i\left(k_{x}^{L, r} x+k_{z}^{L}, r_{z}\right)\right] \text {, } \tag{9.50}
\end{equation*}
$$

which is the same as the scattered field (9.3). The B amplitude will now represent a different reflection coefficient but it is convenient to keep the same notation. The $A$-potential now has incident and scattered parts

$$
\begin{equation*}
A^{(1)}(x, z)=A^{i n}(x, z)+A^{S C}(x, z) \tag{9.51}
\end{equation*}
$$

where

$$
\begin{equation*}
A^{i n}(x, z)=A_{0}^{V} \exp \left[i \left(k_{z}^{T, i} x-k_{z}^{\left.\left.T, i_{z}\right)\right], ~}\right.\right. \tag{9.52}
\end{equation*}
$$

and

$$
\begin{equation*}
A^{s c}(x, z)=C \exp \left[i\left(k_{x}^{T, i_{x}}+k_{z}^{T, i} z\right)\right] \tag{9.53}
\end{equation*}
$$

the latter of which agrees with (9.4) with the same proviso on the reflection coefficient $C$.

The displacement and stress components on the $z=0$ surface are given by

$$
\begin{align*}
& u_{1}^{(1)}(x, 0)=\frac{\partial}{\partial x} p^{(1)}(x, 0)-\frac{\partial}{\partial z} A^{s c}(x, 0)-\frac{\partial}{\partial z} A^{i n}(x, 0),  \tag{9.54}\\
& u_{3}^{(1)}(x, 0)=\frac{\partial}{\partial z} \phi^{(1)}(x, 0)+\frac{\partial}{\partial x} A^{s c}(x, 0)+\frac{\partial}{\partial x} A^{i n}(x, 0), \tag{9.55}
\end{align*}
$$

$\tau_{z 1}^{(1)}(x, 0)=\mu_{1}\left[\frac{\partial^{2}}{\partial x^{2}} A^{s c}(x, 0)-\frac{\partial^{2}}{\partial z^{2}} A^{s c}(x, 0)+2 \frac{\partial^{2}}{\partial x \partial z} \phi^{(1)}(x, 0)\right]$

$$
\begin{equation*}
+\mu_{1}\left[\frac{\partial^{2}}{\partial x^{2}} A^{i n}(x, 0)-\frac{\partial^{2}}{\partial z^{2}} A^{i n}(x, 0)\right] \tag{9.56}
\end{equation*}
$$

and

$$
\begin{align*}
& \tau_{z 3}^{(1)}(x, 0)= 2 \mu_{1}\left[\frac{\partial^{2}}{\partial z^{2} \phi^{(1)}(x, 0)}+\frac{\partial^{2}}{\partial z \partial x} A^{s c}(x, 0)\right]-\lambda_{1} k_{L 1}^{2} \phi(1) \\
&(x, 0)
\end{aligned} \quad \begin{aligned}
& \quad+2 \mu_{1} \frac{\partial^{2}}{\partial z \partial x} A^{i n}(x, 0) \tag{9.57}
\end{align*}
$$

In Region 2 the potentials for the transmitted $P$ - and $S V$-waves are analogous to (9.9) and (9.10)

$$
\begin{equation*}
q^{(2)}(x, z)=D \exp \left[i\left(k_{x}^{L^{2}} x-k_{z}^{L^{2}} z\right)\right] \tag{9.58}
\end{equation*}
$$

and

$$
\begin{equation*}
A^{(2)}(x, z)=E \exp \left[i\left(k_{x}^{T 2}-k_{z}^{T z}\right)\right] \tag{9.59}
\end{equation*}
$$

and the displacement and stress components at $z=0$ given by

$$
\begin{align*}
u_{i}^{(2)}(x, 0)= & \frac{\partial}{\partial x} \phi^{(2)}(x, 0)-\frac{\partial}{\partial z} A^{(2)}(x, 0)  \tag{9.60}\\
u_{3}^{(2)}(x, 0)= & \frac{\partial}{\partial z} \phi^{(2)}(x, 0)+\frac{\partial}{\partial z} A^{(2)}(x, 0)  \tag{9.61}\\
\tau_{z^{1}}^{(2)}(x, 0)= & \mu_{2}\left[\frac{\partial^{2}}{\partial x^{2}} A^{(2)}(x, 0)-\frac{\partial^{2}}{\partial z^{2}} A^{(2)}(x, 0)\right] \\
& +2 \mu_{2} \frac{\partial^{2}}{\partial x^{2} z} \phi \tag{9.62}
\end{align*}
$$

and

$$
\begin{gather*}
\tau_{z^{3}}^{(2)}(x, 0)=2 \mu_{2}\left[\frac{\partial^{2}}{\left.\partial z^{2} p^{(2)}(x, 0)+\frac{\partial^{2}}{\partial x \partial z} A^{(2)}(x, 0)\right]} \begin{array}{rl} 
& -\lambda_{2} k_{L 2}^{2} \phi^{(2)}(x, 0)
\end{array},\right.
\end{gather*}
$$

analogous to (9.11)-(9.14). The boundary conditions of continuity of di splacements and stresses are given by (9.15)-(9.18). Substituting (9.54)(9.57) and (9.60)-(9.63) into the se equations and witing the incident field on the rhs as in (9.19)-(9.22) we get (all fields evaluated at (x,0))

$$
\begin{align*}
& \frac{\partial \phi^{(1)}}{\partial x}-\frac{\partial A^{s c}}{\partial z}-{\frac{\partial \phi^{(2)}}{\partial x}}^{(2)}+\frac{\partial A^{(2)}}{\partial z}={\frac{\partial A^{i n}}{\partial z}}^{i n} .  \tag{9.64}\\
& \frac{\partial \phi}{\partial z}^{(1)}+{\frac{\partial A^{s c}}{\partial x}}^{s c}{\frac{\partial \phi^{(2)}}{\partial z}}^{\left(2 A^{(2)}\right.}{ }^{\partial x}=-\frac{\partial A^{i n}}{\partial x} . \tag{9.65}
\end{align*}
$$

$$
\begin{align*}
& 2 \mu_{1} \frac{\partial^{2} \phi^{(1)}}{\partial x \partial z}+\mu_{1}\left[{\left.\frac{\partial^{2} A^{s c}}{\partial x^{2}}-\frac{\partial^{2} A^{s c}}{\partial z^{2}}\right]}_{-2 \mu_{2} \frac{\partial^{2} \phi}{\partial x \partial z}-\mu_{2}\left[{\frac{\partial^{2} A^{(2)}}{\partial x^{2}}}^{(2)}-{\frac{\partial^{2} A^{2}}{\partial z^{2}}}^{(2)}\right]=-\mu_{1}\left[{\frac{\partial^{2} A^{i n}}{\partial x^{2}}}^{i n}-\frac{\partial^{2} A^{i n}}{\partial z^{2}}\right],} .\right.
\end{align*}
$$

and

$$
\begin{align*}
& 2 \mu_{1}{\frac{\partial^{2} \phi}{\partial z^{2}}}^{(1)}-\lambda_{1} k_{L 1}^{2} \phi^{(1)}+2 \mu_{1} \frac{\partial^{2} A^{s c}}{\partial x \partial z} \\
& -2 \mu_{2}{\frac{\partial^{2} \phi}{\partial z^{2}}}^{(2)}+\lambda_{2} k_{L 2}^{2} \phi^{(2)}-2 \mu_{2} \frac{\partial^{2} A^{(2)}}{\partial x \partial z}=-2 \mu_{1} \frac{\partial^{2} A^{i n}}{\partial z \partial z} . \tag{9.67}
\end{align*}
$$

Again all the $z$-components of the phases of the waveforms in these equations vanish, and, by translational invariance all the x-components are equal. The only difference in the forms from (9.23) is that the incident $P$-wave $k_{x}^{L}, i$ component is replaced by

$$
\begin{equation*}
k_{x}^{T, i}=k_{T_{1}} \sin \theta_{i} \tag{9.68}
\end{equation*}
$$

Equating this to the appropriate remaining components in (9.23) we get the usual laws of reflection and refraction. Substituting the wave forms into (9.64) to (9.67) we get the Zoeppritz equations for this problem

$$
\begin{align*}
& m_{11} S_{3}+m_{22} S_{4}+m_{13} T_{3}+m_{14} T_{4}=V_{1},  \tag{9.69}\\
& m_{21} S_{3}+m_{22} S_{4}+m_{23} T_{3}+m_{24} T_{4}=V_{2},  \tag{9.70}\\
& m_{31} S_{3}+m_{32} S_{4}+m_{33} T_{3}+m_{34} T_{4}=V_{3},  \tag{9.71}\\
& m_{41} S_{3}+m_{42} S_{4}+m_{43} T_{3}+m_{44} T_{4}=V_{4}, \tag{9.72}
\end{align*}
$$

in terms of the two components of the scattering and transmission matrices
for the SV-incidence problem

$$
\begin{align*}
& \mathbf{S}_{3}=\mathbf{B} / \mathbf{A}_{0}^{\mathbf{v}}, \quad \mathbf{S}_{4}=\mathbf{C} / \mathbf{A}_{0}^{\mathbf{v}},  \tag{9.73}\\
& \mathbf{T}_{3}=\mathbf{D} / \mathbf{A}_{0}^{\mathbf{v}}, \quad, \quad T_{4}=E / A_{0}^{\mathbf{v}}, \tag{9.74}
\end{align*}
$$

and the four terms $V_{j}$ related to the incident $S V$-wave field. The matrix components $m_{i j}$ are the same as those for the prave incidence problem and are defined in (9.34). The $V_{j}$ terms replace the $P_{j}$ tems for the $P$-wave incidence and are given by

$$
\left[\begin{array}{l}
V_{1}  \tag{9.75}\\
V_{2} \\
V_{3} \\
V_{4}
\end{array}\right]=\left[\begin{array}{ll}
-i & k_{T_{1}} \\
-i & k_{x} \\
\mu_{1}\left(X_{T 1}^{2}-\right. & \left.k_{z}^{2}\right) \\
-2 \mu_{1} k_{X} X_{T 1}
\end{array}\right]=\left[\begin{array}{l}
m_{12} \\
-m_{22} \\
m_{32} \\
m_{42}
\end{array}\right]
$$

which are related to the second column of the m-matrix in (9.34) corresponding to SV-waves.

The set of equations (9.69)-(9.72) is the same as (9.18)-(9.31) and the solution is in terms of $m^{-1}$ which was previously treated. It is, for the reflection coefficients given by

$$
\begin{equation*}
S_{3}=\Delta^{-1} \sum_{j=1}^{4} D_{j i} V_{j} \tag{9.76}
\end{equation*}
$$

and

$$
\begin{equation*}
S_{4}=\Delta^{-1} \sum_{j=1}^{4} D_{j 2} V_{j} \tag{9.77}
\end{equation*}
$$

in analogy with (9.38) and (9.39). In general we have

$$
\left[\begin{array}{l}
s_{3}  \tag{9.78}\\
S_{4} \\
T_{3} \\
T_{4}
\end{array}\right]=m^{-1}\left[\begin{array}{l}
v_{1} \\
v_{2} \\
v_{3} \\
v_{4}
\end{array}\right] .
$$

The cofactors are defined in (9.42)-(9.49).
Note finally that all the reflection and transmission coefficients for both $P$ - and SV-wave incidence have the same denominator $\Delta$ defined in (9.40).

## SURPACB DISPLACRMRNTS AND STRESSES

We can easily compute the surface values of the displacements and stresses for $p$ and $S V$-wave problems using results in this section, and for SH-incidence using the results in Sec. 8.

## P-VAVES

From (9.5)-(9.8) we get (dropping the superscript because of the continuity conditions)

$$
\begin{align*}
& u_{1}(x, 0)=i\left[k_{x}\left(S_{1}+1\right)-K_{T_{1}} S_{2}\right] A_{0}^{P} \exp \left(i k_{x} x\right)  \tag{9.79}\\
& u_{3}(x, 0)=i\left[\mathbb{K}_{L_{1}}\left(S_{1}-1\right)+k_{x} S_{2}\right] A_{0}^{P} \exp \left(i k_{x} x\right)  \tag{9.80}\\
& \tau_{z^{1}}(x, 0)=\mu_{1}\left[2 k_{x} K_{L_{1}}\left(S_{1}-1\right)+\left(\mathbb{K}_{T_{1}}^{2}-k_{k_{x}}^{2}\right) S_{2}\right] A_{0}^{P} \exp \left(i k_{x} x\right) \tag{9.81}
\end{align*}
$$

and

$$
\begin{equation*}
\tau_{z 3}(x, 0)=-\left[\left(2 \mu_{1} K_{L x}^{2}+\lambda_{1} k_{L 1}^{2}\right)\left(S_{1}+1\right)+2 \mu_{1} k_{x} K_{T 1} S_{2}\right]_{0}^{P} \exp \left(i k_{x} x\right) \tag{9.82}
\end{equation*}
$$

## SV-WAVES

From (9.54)-(9.57) we get

$$
\begin{align*}
& u_{1}(x, 0)=i\left[k_{x} S_{3}+K_{T 1}\left(1-S_{4}\right)\right] A_{0}^{V} \exp \left(i k_{x} x\right),  \tag{9.83}\\
& u_{3}(x, 0)=i\left[K_{L x} S_{3}+k_{x}\left(1+S_{4}\right)\right] A_{0}^{V} \exp \left(i k_{x}\right),  \tag{9.84}\\
& \tau_{z 1}(x, 0)=\mu_{1}\left[-2 k_{x} K_{L 1} S_{3}+\left(K_{T 1}^{2}-k_{x}^{2}\right)\left(S_{4}+1\right)\right] A_{0}^{V} \exp \left(i k_{x}\right), \tag{9.85}
\end{align*}
$$

and

$$
\tau_{z 3}(x, 0)=-\left[\left(2 \mu_{1} K_{L 1}^{2}+\lambda_{1} k_{L 1}^{2}\right) S_{3}+2 \mu_{2} k_{x} K_{T 1}\left(S_{4}-1\right)\right] A_{0}^{v} \exp \left(i k_{x}^{x}\right)
$$

## SH-DAVBS

From the results in (8.6) and (8.22) (or (8.24) or (8.26))

$$
\begin{equation*}
u_{2}(x, 0)=v^{(1)}(x, 0)=\left(1+S_{11}\right) A_{0}^{H} \exp \left(i k_{x} x\right) \tag{9.87}
\end{equation*}
$$

and

$$
\tau_{z^{2}}(x, 0)=\mu_{1} \frac{\partial v^{(1)}}{\partial z}(x, 0)=-i K_{T 1}\left(1-S_{11}\right) A_{0}^{H} \exp \left(i k_{x} x\right) \cdot(9.88)
$$

The result is that each of the displacements and stresses evaluated on the surface can be related to its incident plane wave field on the surface times factors involving the five plane wave reflection coefficients $S_{j}$, $j=1, \ldots, 4$, and $S_{11}$ from (9.32) or (9.38) and (9.39), (9.73) or (9.76) and (9.77), and (8.22) or (8.24) or (8.26) respectively.

In this section we briefly discuss the problem of P－wave incidence on a flat elastic layer，a three region problem．Each of the regions is assumed to have different elastic properties．The geometry is illustrated in Fig． 3.11.


Fig． 3.11

Region（1）has incident and reflected $P$－waves（with the angles of incidence and reflection equal，and a reflected $S V$－wave at a relected angle found via （6．17）．Region（2）has both up－and down－going $P$－and SV－waves（or equivalently standing waves）．Region（3）has down going $P$－and $S V$－waves． No coupling to $S H$－waves occurs becanse of the planar P－wave incidence and the flat geometries．

We thus have eight wave coefficents to solve for，all expressed as a ratio with the amplitude of the incident field．They are the reflection coefficients：

(10.3)
where the + and - symbols refer to up- and down-going waves. There are a total of eight continuity conditions. For $i=1,3$ they are

$$
\begin{array}{ll}
u_{i}^{(1)}=u_{i}^{(2)} & \left(z=z_{1}\right), \\
\tau_{z i}^{(2)}=\tau_{z i}^{(2)} & \left(z=z_{1}\right), \\
u_{i}^{(2)}=u_{i}^{(3)} & \left(z=z_{2}\right), \tag{10.6}
\end{array}
$$

and

$$
\begin{equation*}
\tau_{z i}^{(2)}=\tau_{z i}^{(3)} \quad\left(z=z_{2}\right) \tag{10.7}
\end{equation*}
$$

In each region, we have the representations

$$
\begin{align*}
& \mathbf{u}_{1}=\frac{\partial \phi}{\partial x}-\frac{\partial A}{\partial z},  \tag{10.8}\\
& \mathbf{u}_{3}=\frac{\partial \phi}{\partial z}+\frac{\partial A}{\partial x},  \tag{10.9}\\
& \tau_{z 1}=\mu\left[\frac{\partial^{2} A}{\partial x^{2}}-\frac{\partial^{2} A}{\partial z^{2}}\right]+2 \mu \frac{\partial^{2} \phi}{\partial x^{2} z}, \tag{10.10}
\end{align*}
$$

and

$$
\begin{equation*}
\tau_{z 3}=2 \mu\left[\frac{\partial^{2} \phi}{\partial z^{2}}+\frac{\partial^{2} A}{\partial x \partial z}\right]-\lambda k_{L}^{2} \phi \tag{10.11}
\end{equation*}
$$

Again, we are able to retain the potential representation for each of the se
quantities since we have a planar problem in 2 dimensions, $x$ and $z$. Applying the boundary conditions we are also able to retain the overall coservation of horizontal wave number $k_{x}$. The translational invariance is maintained. We are thas able to write the potential wave shapes in each region as

REGION (1):

$$
\begin{equation*}
\varphi^{(1)}(x, z)=A_{0}^{P} \exp \left[i\left(k_{x} x-K_{L^{1}} z\right)\right]+B \exp \left[i\left(k_{x^{x}}+K_{L^{1}} z\right)\right] \tag{10.12}
\end{equation*}
$$

and

$$
\begin{equation*}
A^{(1)}(x, z)=C \exp \left[i\left(k_{x} x+K_{T i} z\right)\right] \tag{10.13}
\end{equation*}
$$

and to apply the boundary conditions we need to find

$$
u_{1}^{(1)}\left(x, z_{1}\right), u_{3}^{(1)}\left(x, z_{1}\right), \quad \tau_{z_{1}}\left(x, z_{1}\right) \text { and } \tau_{z^{3}}\left(x, z_{1}\right)
$$

## REGION (2):

$$
\begin{equation*}
q^{(2)}(x, z)=\phi_{-} \exp \left[i\left(k_{x} x-k_{L i} z\right)\right]+p_{+} \exp \left[i\left(k_{x} x+K_{L z} z\right)\right] \tag{10.14}
\end{equation*}
$$

and

$$
\begin{equation*}
A^{(2)}(x, z)=V_{-} \exp \left[i\left(k_{z} x-K_{T 2^{2}}\right)\right]+V_{+} e x p\left[i\left(k_{x} x+K_{T_{2}} z\right)\right] \tag{10.15}
\end{equation*}
$$

and to apply the boundary conditions we need to find $u_{1}{ }^{(2)}\left(x_{1}, z_{1}\right)$, $u_{3}^{(2)}\left(x, z_{1}\right), \tau_{z_{1}}^{(2)}\left(x, z_{1}\right)$ and $\tau_{z^{3}}^{(2)}\left(x, z_{1}\right)$ at the upper boundary, and


## REGION (3):

$$
\begin{equation*}
\phi^{(3)}(x, z)=\phi_{3} \exp \left[i\left(k_{x} x-K_{L_{3}} z\right)\right] \tag{10.16}
\end{equation*}
$$

and

$$
\begin{equation*}
A^{(3)}(x, z)=A_{3} \exp \left[i\left(k_{x^{\prime}}-K_{T^{3}} z\right)\right] \tag{10.17}
\end{equation*}
$$

and to use the boundary conditions we need

$$
u_{1}^{(3)}\left(x, z_{2}\right), u_{3}^{(3)}\left(x, z_{2}\right), \tau_{z^{1}}^{(3)}\left(x, z_{2}\right) \text { and } \tau_{z^{3}}^{(3)}\left(x, z_{2}\right)
$$

Note that in general we will be left with exponentials involving both $z_{1}$ and $z_{2}$. We can set one interface at zero, but not both.

### 3.11 MOLTLLAYERS

In this section me briefly describe the systematics of notation, etc., involved in setting up the scattering problem for a planer multilayer structure with no variability in the y-direction. We thus maintain the decoupling between $P$-SV-waves and $S H$-waves. The displacements and stresses are written as usual for any layer in terms of $\phi$ and A potentials as (i= 1,3)

$$
\begin{equation*}
u_{i}(x, z)=\delta_{i 1}\left[\frac{\partial \phi}{\partial x}-\frac{\partial A}{\partial z}\right]+\delta_{i 3}\left[\frac{\partial \phi}{\partial z}+\frac{\partial A}{\partial x}\right] \tag{11.1}
\end{equation*}
$$

and

$$
\begin{align*}
\tau_{z i}(x, z)= & \mu \delta_{i x}\left[\frac{\partial^{2} A}{\partial x^{2}}-\frac{\partial^{2} A}{\partial z^{2}}+2 \frac{\partial^{2} \phi}{\partial x \partial z}\right] \\
& +\delta_{i 3}\left[2 \mu \frac{\partial^{2} \phi}{\partial z^{2}}-\lambda k_{L}^{2} \phi+2 \mu \frac{\partial^{2} A}{\partial x \partial z}\right] \tag{11.2}
\end{align*}
$$

At the $m^{\text {th }}$ interface, $z=z_{m}$ between 1 ayers $m$ and $m+1$, we have the continuity conditions

$$
\begin{equation*}
u_{i}^{(m+1)}\left(x, z_{m}\right)=u_{i}^{(m)}\left(x, z_{m}\right) \tag{11.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\tau_{z i}^{(m+1)}\left(x, z_{m}\right)=\tau_{z i}^{(m)}\left(x, z_{m}\right) \tag{11.4}
\end{equation*}
$$

Also at each interface we have up- and down- going waves from each 1 ayer. The geometry is illustrated in Fig. 3.12. The potentials can be decomposed

in terms of down- and up- going waves in each layer. In the (m) th layer we
can write

$$
\begin{equation*}
\phi^{(m)}(x, z)=d^{(m)} \phi_{d}^{(m)}(x, z)+u^{(m)_{\phi}^{(m)}(x, z)} \tag{11.5}
\end{equation*}
$$

where $\phi_{d}^{(m)}$ is a downard traveling plane wave with amplitude $d^{(m)}$ and $\phi_{u}^{(m)}$ upward traveling with amplitude $u^{(m)}$. For the A-potential we have an analogous expression

$$
\begin{equation*}
A^{(m)}(x, z)=D^{(m)} A_{D}^{(m)}(x, z)+v^{(m)} A_{0}^{(m)}(x, z) \tag{11.6}
\end{equation*}
$$

The wave shapes have the general form

$$
\begin{equation*}
\phi_{\mathrm{d}}^{(m)}(x, z)=\exp \left[i\left(k_{x} x \mp \mathbb{k}_{L}^{(m)} z\right)\right] \tag{11.7}
\end{equation*}
$$

and

$$
\begin{equation*}
A_{d}^{(m)}(x, z)=\exp \left[i\left(k_{x} x \mp K_{T}^{(m)} z\right)\right] \tag{11.8}
\end{equation*}
$$

Osing the se up- and down going wave representations we can wite the displacements and stresses in each region in terms of an up- and down- going wave decomposition. In the $\mathrm{m}^{\text {th }}$ region we have

$$
\begin{equation*}
\mathbf{u}_{i}^{(m)}(x, z)=\mathbf{u}_{i}^{(m, n)}(x, z)+u_{i}^{(m, d)}(x, z) \tag{11.9}
\end{equation*}
$$

and

$$
\begin{equation*}
\tau_{z i}^{(m)}(x, z)=\tau_{z i}^{(m, u)}(x, z)+\tau_{z i}^{(m, d)}(x, z) \tag{11.10}
\end{equation*}
$$

At $z=z_{m}$ the displacement continuity condition is

$$
\begin{equation*}
u_{i}^{(m, u)}\left(x, z_{m}\right)+u_{i}^{(m, d)}\left(x, z_{m}\right)=u_{i}^{(m+1, u)}\left(x, z_{m}\right)+u_{i}^{(m+1, d)}\left(x, z_{m}\right) \tag{11.11}
\end{equation*}
$$

and the stress continuity is

$$
\begin{equation*}
\tau_{z i}^{(m, u)}\left(x, z_{m}\right)+\tau_{z i}^{(m, d)}\left(x, z_{m}\right)=\tau_{z i}^{(m+1, u)}\left(x, z_{m}\right)+\tau_{z i}^{(m+1, d)}\left(x, z_{m}\right) \tag{11.12}
\end{equation*}
$$

There are various way to rewrite the above equations. One possible way is to treat them in analogy with the incident and scattered wave interpretation in Secs. 6-9. Here the "incident' waves are those whose propagation is directed towards the interface, and the "scattered" waves those directed away from it. For example, the displacements in (11.11) would be written as

$$
\begin{equation*}
u_{i}^{(m, u)}\left(x, z_{m}\right)-u_{i}^{(m+1, d)}\left(x, z_{m}\right)=u_{i}^{(m+1, u)}\left(x, z_{m}\right)-u_{i}^{(m, d)}\left(x, z_{m}\right), \tag{11.13}
\end{equation*}
$$

where the "incident" waves are on the rhs, and the "scattered" waves on the 1hs. An analogous equation can be written for stresses, and from the resulting four equations we can wite a matrix equation for the manown (here the "scattered") coefficients in tems of the known ("incident") coefficients as

$$
M^{(m)}\left[\begin{array}{l}
u^{(m)}  \tag{11.14}\\
0^{(m)} \\
d^{(m+1)} \\
D^{(m+1)}
\end{array}\right]=N^{(m)}\left[\begin{array}{l}
u^{(m+1)} \\
0^{(m+1)} \\
d^{(m)} \\
D^{(m)}
\end{array}\right]
$$

in terms of matrices $M$ and $N$ for each interface. These involve the wave shapes evaluated on the interface. The x-variability cancels due to translational invariance.

Alternatively the coefficients of the m region may be known, and it is desired to find the coefficients in the next layer. The se can be determined from (11.11) and (11.12) in terms of a propagation matrix $p(m)$ which "propagates" the coefficients from one 1 ayer to the next, viz.

$$
\left[\begin{array}{l}
u^{(m+1)}  \tag{11.15}\\
0^{(m+1)} \\
d^{(m+1)} \\
d^{(m+1)}
\end{array}\right]=p^{(m)} \quad\left[\begin{array}{l}
u^{(m)} \\
0^{(m)} \\
d^{(m)} \\
d^{(m)}
\end{array}\right]
$$

In this section we give a brief description of the types of waves which can arise at a free surface or a fluid-elastic interface and which in general propagate along the interface and decay away from it. The results are related to the free surface problems in Secs. 6 and 7. We consider only two-dimensional problems.

## (a) FRRE SURPACB

We search for zeroes of the denominator of any of the reflection coefficients in Secs. 6 or 7 , for example the denominator appearing in (6.28). It is convenient to multiply this denominator by $\left(c_{L} c_{T}\right)^{-1}$ and to parameterize the ray parameter in terms of a wave speed as $p=-1 / c$. This Rayleigh denominator then be comes

$$
\begin{equation*}
D_{R}=4\left[\frac{c_{T}}{c}\right]^{2}\left[1-\left[\frac{c_{T}}{c}\right]^{2}\right]^{1 / 2}\left[\left[\frac{c_{T}}{c_{L}}\right]^{2}-\left[\frac{c_{T}}{c}\right]^{2}\right]^{1 / 2}+\left[1-2\left[\frac{c_{T}}{c}\right]^{2}\right]^{2} \tag{12.1}
\end{equation*}
$$

and we want solutions of the secular equation

$$
\begin{equation*}
D_{R}=0 \tag{12.2}
\end{equation*}
$$

for the unknow wave speed $c$ (or the ray parameter). It turns out that there is a real root $c=c_{\mathbf{R}}$ where

$$
\begin{equation*}
0<c_{R}<c_{T}<c_{L} \tag{12.3}
\end{equation*}
$$

At this value of $c$, the vertical components of wave number $K_{L}$ and $K_{T}$ become

$$
\begin{align*}
& \mathbf{K}_{\mathbf{L}}=\mathbf{i} \gamma_{\mathrm{L}},  \tag{12.4}\\
& \mathbf{K}_{\mathbf{T}}=\mathbf{i} \gamma_{\mathrm{T}},
\end{align*}
$$

where

$$
\begin{equation*}
\gamma_{L}^{2}=\omega^{2}\left[c_{R}^{-2}-c_{L}^{-2}\right] \tag{12.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\gamma_{T}^{2}=\omega^{2}\left[c_{R}^{-2}-c_{T}^{-2}\right] \tag{12.6}
\end{equation*}
$$

so that the full time-dependent wave shapes for the potentials in the elastic material are

$$
\begin{equation*}
\phi(x, z)=\hat{p}_{0} \exp \left[-\gamma_{L} z\right] \exp \left[i\left[k_{x} x-\omega t\right]\right] \tag{12.7}
\end{equation*}
$$

and

$$
\begin{equation*}
A(x, z)=A_{0} e \operatorname{xp}\left[{ }^{-\gamma} T^{z}\right] e \operatorname{xp}\left[i\left[k_{x} x-\omega t\right]\right] \tag{12.8}
\end{equation*}
$$

These waves are:
(a) non-dispersive (independent of frequency) since (12.2) is inde pendent of frequency
(b) undampedin the direction of propagation ( $x$ )
(c) damped normal to the bo undary (z)
(d) a coupled compressional-shear system

The wave is called a Rayleigh wave. It is schematically illustrated in Fig. 3.13. The solid 1 ines represent the decay of intensity away from the boundary in the positive z-direction.


Fig. 3.13

## (b) LIQUID-SOLID INTBRFACB

1. GENERALIZED RAYLEIGB PAVES

For this case we al so have a Rayleigh wave, but in addition energy leaks into the 1 iquid, so the Rayleigh wave becomes damped in the direction of motion, and is called a generalized Rayleigh wave. The secular equation is given by

$$
\begin{equation*}
D_{R}=-\left[\frac{\rho_{L}}{\rho_{s}}\right]\left[\frac{\left[c_{T} / / c_{L}\right]^{2}-\left[c_{T} / c\right]^{2}}{\left[c_{T} / c_{L}\right]^{2}-\left[c_{T} / c\right]^{2 / 2}}\right]^{2} \tag{12.9}
\end{equation*}
$$

where $\rho_{L}$ and $\rho_{s}$ are the densities of the 1 iquid and solid respectively, and $c_{L}$ is the sound speed in the 1 iquid. Note that as $\left(\rho_{L} / \rho_{s}\right) \rightarrow 0$, the root $c \rightarrow \mathbf{c}_{R^{\prime}}$ the Rayleigh root. The root of (12.9) is complex and the wave decays as it travels along the surface. If $c_{R}$, is the real part of the velocity then

$$
\begin{equation*}
c_{R}<c_{R}^{\prime}<c_{T}<c_{L} \tag{12.10}
\end{equation*}
$$

Its decay into the 1 iquid categorizes it as a leaky wave. It al so turns out that most of the energy in the wave is in the solid, and it is launched at a specific angle given by $\sin \theta_{R^{\prime}}=c_{L} / c_{R^{\prime}}$.
2. STONELEY MAVE

There al so exists a real root of (12.9) given by

$$
\begin{equation*}
c=c_{S}<c_{L} \tag{12.11}
\end{equation*}
$$

called a Stoneley wave. It propagates paralled to the boundary without
attenuation along $x$. It is exponentially damped in both directions away from the surface, and doesn't always exist at a solid-solid interface. For $\rho_{L} / \rho_{s} \ll 1$, most of the energy is in the 1 iquid.

### 3.13 FRER FLAT SURFACE DSING INTEGRAL BQDATIONS

In Secs. 6-8 we studied the scattering from a free flat elastic surface using conventional methods involving potentials and plane waves. In particular we noted that the $P$ - and SV-waves decoupled from the SH-waves. We also calculated in Sec. 3 the coupled integral equations for the displacement components for scattering from a free surface which however was not flat, but arbitrarily rough. We compute in Sec. 14 , asing perturbation theory, that this roughness induces a coupling between the $P-$, $S V-$, and SHwaves, i.e. a polarization change occurs in the scattering from a rough surface. Our perturbation theory will be about the flat surface 1 imit of the equations in Sec. 3 and we must show that this flat surface limit yields the same results for the total displacement on the surface as those found conventionally in Secs. 6-8. We begin by solving the convolution equation (3.74) and projecting the results on the $x-z$ plane in order to compare with the conventional results (Ref. 3.8).

Equation (3.74) for the (flat) surface values of displacement on a free surface is a convolution equation and can be solved using Fourier transforms. Introduce the two-dimensional transform

$$
\begin{equation*}
u_{j}\left(x_{t}^{\prime}\right)=(2 \pi)^{-2} \iint \exp \left(i k_{t}^{\prime} \cdot{\underset{\sim}{x}}_{\prime}^{\prime}\right) \tilde{u}_{j}\left({\underset{\sim}{k}}_{t}^{\prime}\right) d_{\sim_{t}^{\prime}}^{\prime} \tag{13.1}
\end{equation*}
$$

Using (13.1) and (3.68) in (3.74) we get the result

$$
\begin{equation*}
1 / 2 \tilde{u}_{j}\left({\underset{\sim}{k}}^{\prime}\right)={\underset{u}{j}}_{i n^{n}}^{\left({\underset{\sim}{k}}_{t}^{\prime}\right)}-M_{j i}^{(0)}\left(\underset{\sim}{k_{t}^{\prime}}\right) \tilde{u}_{i}\left({\underset{\sim}{k}}_{t}^{\prime}\right) \tag{13.2}
\end{equation*}
$$

which can al so be written as

$$
\begin{equation*}
\left[\delta_{j i}+2 M_{j i}^{(0)}\left(k_{t}^{\prime}\right)\right] \tilde{u}_{i}\left(k_{t}^{\prime}\right)=2 \tilde{u}_{j}^{i n}\left(k_{t}^{\prime}\right) \tag{13.3}
\end{equation*}
$$

Note that the matrix on the 1 hs of (13.3) has components

$$
I+2 M^{(0)}=\left[\begin{array}{lll}
1 & 0 & 2 M_{13}^{(0)}  \tag{13.4}\\
0 & 1 & 2 M_{23}^{(0)} \\
2 M_{31}^{(0)} & 2 M_{32}^{(0)} & 1
\end{array}\right]
$$

Define a matrix $L$ as the following inverse

$$
\begin{equation*}
L_{m j}\left({\underset{\sim}{t}}_{\prime}^{\prime}\right)\left[\delta_{j i}+2 M_{j i}^{(0)}\left({\underset{\sim}{k}}_{t}^{\prime}\right)\right]=\delta_{m i} \tag{13.5}
\end{equation*}
$$

which is explicity given by

$$
L=\Delta^{-1}\left[\begin{array}{lll}
1-4 M_{32}^{(0)} M_{23}^{(0)} & 4 M_{13}^{(0)} M_{32}^{(0)} & -2 M_{13}^{(0)}  \tag{13.6}\\
4 M_{31}^{(0)} M_{23}^{(0)} & 1-4 M_{13}^{(0)} M_{31}^{(0)} & -2 M_{23}^{(0)} \\
-2 M_{31}^{(0)} & -2 M_{32}^{(0)} & 1
\end{array}\right]
$$

where the denominator determinant is given by

$$
\begin{equation*}
\Delta=1-4 M_{13}^{(0)} M_{31}^{(0)}-4 M_{32}^{(0)} M_{23}^{(0)} \tag{13.7}
\end{equation*}
$$

Multiply (13.3) by $L_{\text {mj }}$ and the result is the solution

$$
\begin{equation*}
\tilde{\mathbf{u}}_{m}\left(\underset{\sim}{k_{t}^{\prime}}\right)=2 L_{m j}\left(\underset{\sim}{k_{t}^{\prime}}\right) \tilde{u}_{j}^{i n}\left(\underset{\sim}{k_{t}^{\prime}}\right) \tag{13.8}
\end{equation*}
$$

for the Fourier transform of the total surface field values in terms of the

Fourier transform of the incident field. The spatial value is found by Fourier transforming (13.8) to yield

$$
\begin{equation*}
u_{m}\left({\underset{x}{t}}_{\prime}\right)=2 \iint \tilde{L}_{m j}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right) u_{j}^{i n}\left({\underset{\sim}{x}}_{t}\right) d{\underset{\sim}{x}}_{t}, \tag{13.9}
\end{equation*}
$$

where L is the Fourier transform of (13.6) given by

$$
\begin{equation*}
\tilde{L}_{m j}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right)=(2 \pi)^{-2} \iint e^{i k_{t}^{\prime} \cdot\left(x_{t}^{\prime}-{\underset{\sim}{x}}_{t}^{\prime}\right)} L_{m j}\left(k_{t}^{\prime}\right) d k_{t}^{\prime} \tag{13.10}
\end{equation*}
$$

with $L_{m j}$ given by (13.6). Since no components of $L$ vanish, (13.9) illustrates that for a flat surface in two dimensions. all the displacement components are related to all the incident field components. For example, an SH-incident component $u_{2}{ }^{i n}$ couples to the $P$ - $S V$ components $u_{1}$ and $u_{3}$ on the surface. Equation (13.9) is the 1 imit of arbitrary incidence (a threedimensional plane-wave for example) so all displacement components couple. This is not yet the cases we discussed in Secs. 6-8 since we have not projected our results onto the $x-z$ plane. We do this now.

Assume the incident displacement field on the surface is independent of y. i.e. let

$$
\begin{equation*}
u_{j}^{i n}\left({\underset{\sim}{x}}_{t}\right)=u_{j}^{i n}(x) \tag{13.11}
\end{equation*}
$$

This now corresponds to the incident fields evaluated on the surface $z=0$ in Secs. 6-8. Here what it means is that we can carry out the y-integration in (13.9). That is using ( 13.10 ) we get

$$
\begin{equation*}
\int d y \tilde{L}_{m j}\left(x_{t}^{\prime}-x_{t}\right)=(2 \pi)^{-1} \int d k_{x}^{\prime} L_{m j}\left(k_{x}^{\prime}, 0\right) e^{i k_{x}^{\prime}\left(x^{\prime}-x\right)}, \tag{13.12}
\end{equation*}
$$

which is independent of $y^{\prime}$. Further we note from (3.73) that for $k_{y}^{\prime}=0$ we
have that

$$
\begin{equation*}
M_{23}^{(0)}\left(k_{x}^{\prime}, 0\right)=M_{32}^{(0)}\left(k_{x}^{\prime}, 0\right)=0 \tag{13.13}
\end{equation*}
$$

$$
\begin{equation*}
M_{13}^{(0)}\left(k_{X}^{\prime}, 0\right)=-k_{X}^{\prime}\left(k_{T}^{2}-2 K_{T}^{\prime} K_{L}^{\prime}-2 k_{X}^{\prime 2}\right) / 2 k_{T}^{2} K_{L}^{\prime} \tag{13.14}
\end{equation*}
$$

and

$$
M_{31}^{(0)}\left(k_{X}^{\prime}, 0\right)=-M_{13}^{(0)}\left(k_{x}^{\prime}, 0\right)\left(K_{L}^{\prime} / K_{T}^{\prime}\right)
$$

where here $K_{L}^{\prime}=\left(k_{L}^{2}-k_{X}^{\prime 2}\right)^{1 / 2}$ and $K_{T}^{\prime}=\left(k_{T}^{2}-k_{X}^{\prime 2}\right)^{1 / 2}$, and from (13.7)

$$
\begin{equation*}
\Delta=1-4 M_{13}^{(0)}\left(k_{x}^{\prime}, 0\right) M_{31}^{(0)}\left(k_{x}^{\prime}, 0\right) \tag{13.16}
\end{equation*}
$$

The resulting matrix $L$ is given from (13.6) and (13.13)

$$
L\left(k_{x}^{\prime}, 0\right)=\Delta^{-1}\left[\begin{array}{lll}
1 & 0 & -2 M_{13}^{(0)}\left(k_{I}^{\prime}, 0\right)  \tag{13.17}\\
0 & \Delta & 0 \\
-2 M_{31}^{(0)}\left(k_{x}^{\prime}, 0\right) & 0 & 1
\end{array}\right]
$$

with $\Delta$ from (13.16). The resulting surface field values can be written using (13.9) and (13.12) as

$$
\begin{equation*}
u_{m}\left(x^{\prime}, 0\right)=\pi^{-1} \int d k_{x}^{\prime} L_{m j}\left(k_{x}^{\prime}, 0\right) e^{i k_{x}^{\prime} x^{\prime}} \sim_{u_{j} n}\left(k_{x}^{\prime}, 0\right), \tag{13.18}
\end{equation*}
$$

in terms of the Fourier transform of the incident displacement on the surface. Now because of the form of the projected value of $L$ in (13.17). (13.18) illustrates the fact that the 1- and 3- components of total displacement $P$ and $S V$ only couple to the 1- and 3- components of incident displacement, and both decouple from the 2-component (SH) of displacement. This corresponds to our examples in Secs. 6-8.

Note that the cross-coupling between $P-S V$ and $S H$ waves was destroyed by projecting onto the $y=0$ plane ( $x, z$ plane). This was accomplished by choosing the incident field to be a function of only $x$ and $z$, with the corresponding $v a l u e$ of the incident field on the $z=0$ surface to be only a function of $x$. We must still show however that the results in (13.18) reduce to our results in Secs. 6-8. We do two cases, the first for SH waves which is easy, and the second for $P$-wave incidence which is more involved.

## CASE 1 - SE-WAVE INCIDENC:

From Sec. 8 we know that on the surface $z=0$ the incident displacement is

$$
\begin{equation*}
u_{2}^{i n}(x, 0)=A_{0}^{H} \exp \left(i k_{x}^{T, i} x\right) \tag{13.19}
\end{equation*}
$$

Its Fourier transform is

$$
\begin{equation*}
\tilde{u}_{2}^{i n}\left(k_{x}^{\prime}, 0\right)=2 \pi A_{0}^{H} \delta\left(k_{x}^{\prime}-k_{x}^{T, i}\right) \tag{13.20}
\end{equation*}
$$

and the total displacement on the surface is with $B=A$

$$
\begin{equation*}
u_{2}\left(x^{\prime}, 0\right)=v\left(x^{\prime}, 0\right)=2 A_{0}^{H} \exp \left(i k_{x}^{\left.T, i_{x^{\prime}}\right)}\right. \tag{13.21}
\end{equation*}
$$

From the integral equation (13.18) we get that

$$
\begin{equation*}
u_{2}\left(x^{\prime}, 0\right)=\pi^{-1} \int d k_{x}^{\prime} \exp \left(i k_{x}^{\prime} x^{\prime}\right){\underset{u}{j}}_{i n}^{\left(k_{x}^{\prime}, 0\right)} \tag{13.22}
\end{equation*}
$$

since $L_{22}=1$ from (13.17). Substituting (13.20) in (13.22) we again recover (13.21). Thus for SH-wave incidence on a free flat surface (in one dimension), the integral equation produces the result found in Sec. 8 .

## CASE 2 - P-WAVE INCIDENCE

From (6.37) the first component of displacement on the surface is given by

$$
\begin{equation*}
u_{1}(x, 0)=i A_{0}\left[k_{x}\left(1+B / A_{0}\right)-k_{T}\left(C / A_{0}\right)\right] e^{i k x^{x}} \tag{13.23}
\end{equation*}
$$

where all $k_{x}$ components are equal and where $\mathrm{K}_{\mathrm{Z}}, \mathrm{r}_{\mathrm{r}}=\mathrm{K}_{\mathrm{T}}$. From (6.30) and (6.32) it is possible to write

$$
\begin{equation*}
\mathbf{B} / \mathbf{A}_{0}=(L-R) /(L+\mathbf{R}): \mathbf{C} / \mathbf{A}_{0}=\mathbf{Q} /(L+\mathbf{R}) \tag{13.24}
\end{equation*}
$$

where

$$
\begin{align*}
& L=4 p^{2}\left(\cos \theta_{L_{i}} / c_{L}\right)\left(\cos \theta_{T_{\mathbf{r}}} / c_{T}\right)  \tag{13.25}\\
& \mathbf{R}=\left(c_{T}^{-2}-2 p^{2}\right)^{2} \tag{13.26}
\end{align*}
$$

and

$$
\begin{equation*}
Q=-4 p\left(\cos \theta_{L i} / c_{L}\right)\left(c_{T}^{-2}-2 p^{2}\right) \tag{13.27}
\end{equation*}
$$

Note the additional symmetry restriction

$$
\begin{equation*}
Q / 2 R=(2 L / Q)\left(K_{L} / \mathbb{K}_{T}\right) . \tag{13.28}
\end{equation*}
$$

From (13.23) we can thas define

$$
\begin{equation*}
a=k_{X}\left(1+B / A_{0}\right)-\mathbb{E}_{T}\left(C / A_{0}\right) \tag{13.29}
\end{equation*}
$$

Osing ( 13.24 ) it becomes

$$
\begin{equation*}
a=\left(2 R k_{x}-Q \mathbb{R}_{T}\right) /(L+R) \tag{13.30}
\end{equation*}
$$

or
$a=(L+R)^{-1}\left[K_{X} 8 p^{2}\left(\cos \theta_{L_{i}} / c_{L}\right)\left(\cos \theta_{T_{T}} / c_{T}\right)+4 p \mathbb{K}_{T}\left(\cos \theta_{L i} / c_{L}\right)\left(c_{T}^{-2}-2{ }^{2}\right)\right]$.

Similarly from (6.38) the third component of displacement on the surface is given by

$$
\begin{equation*}
u_{3}(x, 0)=-i A_{0}\left[\underline{R}_{L}\left(1-B / A_{0}\right)-k_{x} c / A_{0}\right]^{i k_{x}}, \tag{13.32}
\end{equation*}
$$

and if we define

$$
\begin{equation*}
\beta=\mathbb{K}_{L}\left(1-B / A_{0}\right)-E_{X} C / A_{0}, \tag{13.33}
\end{equation*}
$$

we can write it using (13.24) as

$$
\begin{equation*}
\beta=\left(2 R \mathbf{K}_{\mathbf{L}}-\mathbf{Q} \mathbf{K}_{\mathbf{Z}}\right) /(\mathbf{L}+\mathbf{R}) \tag{13.34}
\end{equation*}
$$

Similarly we can write

$$
\begin{equation*}
\beta-2 \mathbf{K}_{\mathbf{L}}=\left(-2 \mathbf{L} \mathbf{K}_{\mathbf{L}}-\mathbf{Q} \mathbf{k}_{\mathbf{x}}\right) /(\mathbf{L}+\mathbf{R}) \tag{13.35}
\end{equation*}
$$

or substituting $f$ or the numerator from (13.25) and (13.27)

$$
\begin{align*}
& \beta-2 \mathbb{K}_{L}=(L+R)^{-1}\left[-K_{L} 8 p^{2}\left(\cos \theta_{L i} / c_{1}\right)\left(\cos \theta_{T r} / c_{T}\right)\right. \\
&\left.+k_{X} 4 p\left(\cos \theta_{L i} / c_{L}\right)\left(c_{T}^{-2}-2 p^{2}\right)\right] . \tag{13.36}
\end{align*}
$$

Dividing (13.36) by (13.31) and asing

$$
\begin{equation*}
\mathbf{k}_{\mathrm{X}}=\omega \sin \theta_{\mathrm{L}_{\mathrm{i}}} / c_{\mathrm{L}}, \mathbf{K}_{\mathrm{L}}=\omega \cos \theta_{\mathrm{L}_{\mathrm{i}}} / c_{\mathrm{L}}, \mathbb{E}_{\mathrm{T}}=\omega \cos \theta_{\mathrm{T}_{\mathrm{r}}} / c_{\mathrm{T}} \tag{13.37}
\end{equation*}
$$

we get

$$
\begin{equation*}
\frac{\left(\beta-2 \mathbf{K}_{L}\right)}{a}=\frac{4 p\left(c_{T}^{-2}-2 p^{2}\right)\left(\sin \theta_{L_{i}} / c_{L}\right)-8 p^{2}\left(\cos \theta_{T_{r}} / c_{T}\right)\left(\cos \theta_{L_{i}} / c_{L}\right)}{4 p\left(c_{T}^{-2}-2 p^{2}\right)\left(\cos \theta_{T r} / c_{T}\right)+8 p^{2}\left(\cos \theta_{T_{r}} / c_{T}\right)\left(\sin \theta_{L_{i}} / c_{L}\right)} . \tag{13.38}
\end{equation*}
$$

We will use the se results to simplify the matrix elements from the integral equation approach below.

We now compate the values $a_{1}(x, 0)$ and $a_{1}(x, 0)$ from the integral equation approach. For p-wave incidence we have that on tho surface the incident displacement is

$$
\begin{equation*}
\mathfrak{n}_{j}^{i n}(x, 0)=i A_{0}\left(k_{x} \delta_{j 2}-\mathbb{K}_{L} \delta_{j 3}\right) e^{i k_{z} x} \tag{13.39}
\end{equation*}
$$

and its Fourier transform is

$$
\begin{equation*}
\tilde{u}_{j}^{i n}\left(k_{x}^{\prime}, 0\right)=2 \pi i A_{0}\left(k_{z} \delta_{j 2}-K_{L} \delta_{j z}\right) \delta\left(k_{z}^{\prime}-k_{z}\right) . \tag{13,40}
\end{equation*}
$$

Substituting this in (13.18) wo have that

$$
\begin{equation*}
\left.a_{m}(x, 0)=2 i A_{0}\left[k_{x_{m 1}} L_{m^{2}}\left(k_{x}, 0\right)-K_{L_{m 3}} L_{x^{1}}, 0\right)\right] e^{i k_{x} z} . \tag{13.41}
\end{equation*}
$$

Evaluating the matrix elements from (13.17) and defining

$$
\begin{equation*}
0=M_{13}^{(0)}\left(\mathbf{k}_{z^{2}}, 0\right), \tag{13.42}
\end{equation*}
$$

so that from (13.15)

$$
\begin{equation*}
M_{31}^{(0)}\left(\underline{k}_{z}, 0\right)=-0 K_{L} / \mathbb{R}_{T} . \tag{13,43}
\end{equation*}
$$

we get for $m=1$

$$
\begin{equation*}
u_{1}(x, 0)=2 i A_{0}\left(k_{x}+3 K_{L} 0\right)\left(1+40^{2} K_{L} / \mathbb{K}_{T}\right)^{-1} e^{i k_{x} x} . \tag{13.44}
\end{equation*}
$$

and for $m=3$

$$
\begin{equation*}
u_{3}(x, 0)=-2 i A_{0} \mathbb{R}_{L}\left(1-2 k_{z} 0 / \mathbb{R}_{T}\right)\left(1+40^{2} \mathbb{E}_{L} / \mathbb{R}_{T}\right)^{-1} e^{i k_{x} x} . \tag{13.45}
\end{equation*}
$$

Comparing these results to (13.23) with (13.29) and (13.22) with (13.33) we get agreement of the displacement components provided

$$
\begin{equation*}
2\left(k_{X}+2 \mathbb{K}_{L} 0\right)\left(1+40^{2} \mathbb{K}_{L} / \mathbb{K}_{T}\right)^{-1}=a \text {. } \tag{13.46}
\end{equation*}
$$

and

$$
\begin{equation*}
2 \mathbb{R}_{L}\left(1-2 k_{Z} 0 / \mathbb{K}_{\mathrm{T}}\right)\left(1+40^{2} \mathbb{K}_{\mathrm{L}} / \mathbb{R}_{\mathrm{T}}\right)^{-1}=\beta \tag{13.47}
\end{equation*}
$$

Dividing ( 13.46 ) and ( 13.47 ) and solving the resalt for $V$ we get

$$
\begin{equation*}
0=-\frac{\mathbf{K}_{T}}{2 \mathbf{K}_{L}} \frac{\beta \mathbf{Z}_{\mathbf{I}}-\alpha \mathbf{K}_{L}}{\alpha K_{\mathbf{I}}+\beta \mathbf{K}_{\mathbf{T}}} . \tag{13.48}
\end{equation*}
$$

We next use the definition of 0 to wite it in another form involving a and B, and show the comparison of the result with (13.48) is an identity.

From (13.42) and the definition (13.14) we have

$$
\begin{equation*}
0=-\left(k_{X} / 2 \mathbf{K}_{L}\right)\left(1-2 k_{X}^{2} \mathbf{k}_{T}^{-2}-2 \mathbf{K}_{T} \mathbf{K}_{L} \mathbf{k}^{-2}\right) \tag{13.49}
\end{equation*}
$$

Osing the definition of the ray parameter we can write $k_{X^{2}} \mathbf{T}^{-2}=p^{2} c_{T}{ }^{2}$ and using (13.37) 0 becomes

$$
\begin{equation*}
0=-\frac{\mathbf{k}_{X}}{2 \bar{K}_{L}} \frac{\left(c_{T}^{-2}-2 p^{2}\right)-2\left(\cos \theta_{T_{T}} / c_{T}\right)\left(\cos \theta_{L i} / c_{L}\right)}{\left(c_{T}^{-3}-2 p^{2}\right)+2 p^{2}} \tag{13.50}
\end{equation*}
$$

where wo have divided by $c_{T}^{2}$ and added and subtracted a term in the denominator. Multiplying numerator and denominator using the identity

$$
\begin{equation*}
\frac{4 p^{2}}{4 p^{2}} \frac{\cos \theta_{T_{r}} / c_{T}}{\cos \theta_{T r} / c_{T}}=\frac{\cos \theta_{T r}}{p^{c} T} \frac{4 p^{2}}{4 p \cos \theta_{T r} / c_{T}} \tag{13.51}
\end{equation*}
$$

we get

$$
0=-\frac{k_{X}}{2 Z_{L}} \frac{\cos \theta_{T r}}{p c_{T}}\left[\frac{4 p^{2}\left(c_{T}^{-2}-2 p^{2}\right)-8 p^{2} \cos \theta_{T r} \cos \theta_{L i} / c_{T} c_{L}}{4 p\left(c_{T}^{-2}-2 p^{2}\right)\left(\cos \theta_{T r} / c_{T}\right)+8 p^{3}\left(\cos \theta_{T r} / c_{T}\right)}\right]
$$

Dsing $p=\sin \theta_{L i} / c_{L}$ in one power of $p$ in the left hand term in the numerator of the bracket term, we see that the bracket term is just (13.38). Substituting for the remainder of the coefficients in $V$ we get finally

$$
\begin{equation*}
0=-\frac{\mathbf{K}_{\mathbf{I}}}{2 \mathbf{E}_{L}} \frac{\left(\beta-2 \mathbf{k}_{L}\right)}{\alpha} \tag{13.53}
\end{equation*}
$$

This is the expression for 0 directly from the matrix element definition. Comparing it to (13.48) from the integral equation we must prove that

$$
\begin{equation*}
\left(\beta \mathbf{k}_{\mathrm{I}}-\alpha \mathbf{K}_{L}\right)\left(\alpha \mathbf{k}_{\mathrm{I}}+\beta \mathbf{K}_{\mathrm{T}}\right)^{-1}=\left(\beta-2 \mathbf{K}_{L}\right) \alpha^{-1} \tag{13.54}
\end{equation*}
$$

is true. Cross maltipication of (13.54) yields the relation

$$
K_{L} \alpha^{2}+K_{T} \beta^{2}=3 K_{L}\left(k_{x} \alpha+K_{T} \beta\right)
$$

This is easily proved to be an identity by substituting (13.30) for a and (13.34) for $\beta$.

We have this shown that the value of the matrix olement 0 necessary for the integral equation to agree with the standard result (13.48) is the same as its definition (13.53). The integral equation results are thas the same as the standard results for P-wave incidence.

### 3.14

In Sec. 3 we presented the surface integral equations and field representations for displacements $\mathbf{u}_{\mathbf{j}}$ for the case of roagh surface $h$. These were oxact but formal expressions for the coupling of the displacement components due to the horizontal variability induced by $h$. In particular it was at least fomally illustrated how the incident displacement field polarization (i.e. its particalar vector value) coupled to all the total surface displacement field values, and correspondingly to all the displacement values of the surface. We also compated the flat surface 1 imit in Sec. 3 and showed in Sec. 13 that this flat surface 1 imit agreed with our standard results using potentials in Secs. 6-8.

In this section we are more explicit. We show how the polarization changes for the case of a free surface (Sec. 3. Ex. 2) for a deterministic surface $h$ which is small, i.e. in the sense that the problem can be treated in a perturbation theory expansion in powers of $h$. From (3.39) we had the integral equation for displacements on the surface given by

$$
\begin{equation*}
Q_{j i}\left(x_{t}^{\prime}\right) u_{i}\left(x_{s}^{\prime}\right)=n_{j}^{i n}\left(x_{s}^{\prime}\right)+\iint X_{j i}\left({\underset{\sim}{x}}_{s}^{\prime} \cdot x_{s}\right) n_{i}\left(x_{s}\right) d \underline{z}_{t} . \tag{14.1}
\end{equation*}
$$

Where, from (3.23), we had

$$
\begin{equation*}
Q_{j i}\left(x^{\prime}\right)=\frac{1}{2} \delta_{j i}+Q_{j i}^{(1)}\left({\underset{z}{t}}_{\prime}^{\prime}\right) \tag{14.2}
\end{equation*}
$$

where

$$
\begin{equation*}
a_{j i}^{(1)}\left(x_{t}^{\prime}\right)=\frac{1}{2}\left[\delta_{i,} \partial_{j t}^{\prime} h\left({\underset{\sim}{t}}_{t}^{\prime}\right)+\Lambda_{j z} \partial_{i t} h\left(x_{t}^{\prime}\right)\right] \tag{14.3}
\end{equation*}
$$

is first order in the height h. Further, from (A.32) we had

$$
\begin{equation*}
X_{j i}\left({\underset{\sim}{x}}_{s}^{\prime},{\underset{\sim}{x}}_{s}\right)=n_{p}\left(x_{i}\right) X_{p i j}\left({\underset{\sim}{x}}_{s}^{\prime}-x_{s}\right) \tag{14.4}
\end{equation*}
$$

where

$$
\begin{equation*}
n_{p}({\underset{\sim}{x}})=\delta_{p 3^{3}}-\partial_{p t} h\left({\underset{\sim}{x}}_{t}\right), \tag{14.5}
\end{equation*}
$$

and, from (A.32),

$$
\begin{equation*}
X_{p i j}\left({\underset{\sim}{x}}_{s}^{\prime}-x_{s}\right)=(2 \pi)^{-3} \iiint e \operatorname{xp}\left[i k \cdot\left({\underset{\sim}{x}}_{s}^{\prime}-z_{s}\right)\right] M_{p i j}(\underset{\sim}{k}) d \underline{z} \tag{14.6}
\end{equation*}
$$

with ${ }_{\text {pij }}$ defined using (A.29), (A.23) and (A.7) as

$$
\begin{aligned}
M_{p i j}(k)= & k_{T}^{-2}\left[\tilde{G}^{T}(k) p_{p i j}^{T}(k)-\tilde{G}^{L}(k) p_{p i j}^{L}(\underline{\sim})\right] \\
& -\frac{1}{2}\left[8_{i j} \tilde{G}^{T}(k) P_{p}^{T}(\underline{\sim})+\delta_{p j} \tilde{G}^{T}(k){\underset{p}{i}}_{T}^{(k)}+\delta_{i p} \wedge \tilde{G}^{L}(k) p_{j}^{L}(k)\right]
\end{aligned}
$$

Here the $P_{p i j}^{T}$ terms are defined in (A.21) with $P_{p i j}^{L}$ defined by (A.21) if $\mathbf{k}_{\mathrm{T}}$ is replaced by $k_{L}$, and the $P_{j}$ terms defined by (5.18) in Ch. 1 with the appropriate wavenumber $\mathbf{k}_{\mathbf{T}}$ or $\mathbf{k}_{\mathbf{L}}$ 。

We next expand each of the terms in (14.1) to first order in h. We have

$$
\begin{align*}
& u_{i}\left(x_{s}^{\prime}\right)=u_{i}^{(0)}\left(x_{t}^{\prime}\right)+h\left({\underset{\sim}{x}}_{t}^{\prime}\right) u_{i}^{(1)}\left({\underset{\sim}{x}}_{t}^{\prime}\right) \quad,  \tag{14.8}\\
& u_{j}^{i n}\left(x_{j}^{\prime}\right)=u_{j}^{i n}\left(x_{\mathfrak{t}}^{\prime}\right)+h\left(z_{\mathfrak{t}}^{\prime}\right) v_{j}^{(1)}\left(z_{\mathfrak{t}}^{\prime}\right) \quad, \tag{14.9}
\end{align*}
$$

and to first order

$$
\begin{align*}
X_{j i}\left(x_{s}^{\prime}, x_{s}\right)= & X_{j i}^{0}\left(x_{t}^{\prime}-x_{t}\right)+\left[h\left(x_{t}^{\prime}\right)-h\left({\underset{\sim}{x}}_{t}\right)\right] B_{j i}\left({\underset{\sim}{x}}^{\prime}-x_{t}\right) \\
& -\partial_{p t} h\left({\underset{\sim}{x}}_{t}\right) K_{p i j}\left({\underset{\sim}{x}}^{\prime}-{\underset{\sim}{x}}_{t}\right) \tag{14.10}
\end{align*}
$$

where $\mathrm{K}_{\mathrm{j} i}^{0}$ is defined in (3.59) and

$$
\begin{equation*}
B_{j i}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right)=(2 \pi)^{-3} \iiint \exp \left[i{\underset{\sim}{t}}_{t} \cdot\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right)\right] \quad i k_{z} M_{3 i j}(k) d k \tag{14.11}
\end{equation*}
$$

The terms zeroth-order in $h$ on the right and left hand sides of (14.1) form the flat surface 1 imit discussed in Secs. 3 and 13. Equating the first order terme, we can write a convolntion equation for the quantity

$$
\begin{equation*}
w_{j}\left({\underset{\sim}{x}}_{t}\right)=h\left({\underset{\sim}{x}}_{t}\right) \mathfrak{u}_{j}^{(1)}\left({\underset{\sim}{x}}_{t}\right) \tag{14.12}
\end{equation*}
$$

which is given by

$$
\begin{equation*}
\frac{1}{2} m_{j}\left(x_{t}^{\prime}\right)=b_{j}\left(x_{t}^{\prime}\right)+\iint \mathbb{E}_{j i}^{0}\left(x_{t}^{\prime}-x_{t}\right) m_{i}\left(x_{t}\right) d x_{t} \tag{14.13}
\end{equation*}
$$

where the Born term is known and given by

$$
\begin{align*}
& b_{j}\left({\underset{\sim}{x}}_{t}^{\prime}\right)=h\left(x_{t}^{\prime}\right) \nabla_{j}^{(1)}\left({\underset{\sim}{x}}_{t}^{\prime}\right)-Q_{j i}^{(i)}\left(x_{t}^{\prime}\right) u_{i}^{i n}\left(x_{t}^{\prime}\right) \\
& +\iint\left[h\left({\underset{\sim}{x}}_{t}^{\prime}\right)-h\left({\underset{\sim}{x}}_{t}\right)\right] B_{j i}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right) u_{i}^{i n}\left({\underset{\sim}{N}}_{t}\right) d \underline{x}_{t} \\
& -\iint \partial_{\rho t} h\left({\underset{\sim}{x}}_{t}\right) K_{p i j}\left({\underset{\sim}{x}}_{t}^{\prime}-x_{t}\right) u_{i}^{i n}\left(x_{t}\right) d x_{t} . \tag{14.14}
\end{align*}
$$

Equation (14.13) can be solved by Fonrier transform techniques. In fact we have already solved the equation since it is the same as (3.74) if we replace the incident displacement on the surface by $b_{j}$. The solution of (14.13) is given in analogy with (13.9) as

$$
\begin{equation*}
W_{i}\left(x_{t}^{\prime}\right)=2 \iint \tilde{L}_{i j}\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right) b_{j}\left({\underset{\sim}{x}}_{t}\right) d{\underset{\sim}{x}}_{t}, \tag{14.15}
\end{equation*}
$$

where the matrix $L$ is defined in (13.6). To first order the fall solution of the surface displacement is, using (14.8) and (14.12)

$$
\begin{equation*}
\mathbf{u}_{i}\left({\underset{\sim}{x}}_{s}^{\prime}\right)=u_{i}^{(0)}\left({\underset{\sim}{x}}_{t}^{\prime}\right)+\mathbb{T}_{i}\left({\underset{\sim}{x}}_{t}^{\prime}\right) \tag{14.16}
\end{equation*}
$$

where $u_{i}^{(0)}$ is given by (13.9), i.e. the total displacement field at a flat surface where we have as yet not projected the incident field onto the $x-z$ plane. We explicitly showed how this projection reduced $u_{i}^{(0)}$ to the standard results for $S$-wave incidence in (13.22) and for P-wave incidence from (13.44) and (13.45) (with the subsequent proof of an identity.) As we remarked, the coupling of all $P-S V-$, and $S H-w a v e s$ for this case was due solely to the non-planar nature of the incident displacement. The conpling in the $W_{i}$ term however arises both from this non-planar incident displacement as well as the height and slope variability of the surface. Combining (13.9) and (14.15) we can write to first order approximation in surface height that the total displacement field as the surface is

$$
\begin{equation*}
u_{i}\left(x_{i}^{\prime}\right)=2 \iint \tilde{L}_{i j}\left({\underset{\sim}{x}}_{t}^{\prime}-z_{i}\right)\left[u_{j}^{i n}\left({\underset{\sim}{x}}_{t}\right)+b_{j}\left({\underset{\sim}{x}}_{t}\right)\right] d{\underset{\sim}{x}}_{t} . \tag{14.17}
\end{equation*}
$$

The displacement components off the surface (i.e. at a field point) can be evaluated using (14.17) in (3.38).

### 3.15 KIRCBROFR APPROXIHATION

For the scalar case, Ch. 2 Sec. 6, we found approximate expressions for the surface values of the total (velocity potential) field and its normal derivative $N$ in terms of the reflection coefficient $R$ and the incident field $q^{\text {in }}$. These were

$$
\begin{equation*}
\phi\left(\underline{x}_{s}\right)=(1+R) \phi^{i n}\left(\underline{x}_{s}\right) \tag{15.1}
\end{equation*}
$$

and

$$
\begin{equation*}
N\left({\underset{\sim}{x}}_{s}\right)=(1-R) n_{m}\left({\underset{\sim}{x}}_{t}\right) \partial_{m} q^{i n}({\underset{x}{s}}) \tag{15.2}
\end{equation*}
$$

These were three-dimensional approximations in the sense that the evaluation was on a surface $h(\underset{\sim}{x})$ which was a function of two variables. The incident field in general had a component ont of the $x-z$ plane. Surface slope terms appeared in the normal $n_{m}$. Also, of course, the reflection coefficient $R$ was given for the full transmission problem into another medimm with different parameters. The results al so rednced to the cases where we had perfect reflection, $B=1$ (Nemann boundary condition) and $R=-1$ (Dirichlet boundary condition). Only a single scalar field was incident on the surface, and the terms involving the reflection coefficients (1 $\mathrm{I}_{\mathrm{R}}$ ) were coordinate-independent.

For the elastic case we have threc possible incident fields. $P$, $S V$, and SB, so the factoring out of the incident field becomes a problem. We al so must use the reflection coefficients for the full transmission problem as in Secs. 8 and 9. It is not enough to choose the reflection coefficients for the free or perfectly rigid surface since these do not include energy loss in the lower medinm. In addition there are a total of five reflection coofficients to consider, $S_{j}, j=1, \ldots, 4$, from Sec. 9 and $S_{11}$ from Sec. 8 . Me computed the flat surface values of total displacement and stress at the
end of Sec. 9.
Before discussing the Kirchhoff approximation we consider some remarks:

1. We have expressed the displacements and stresses in terms of potentials $\rho$, $A$, and $v$. All were assumed independent of $y$, and this is why the $P$ and $S V$ potentials decoupled from the $S H$ potential. Indeed, this is why we were able to make the Cauchy-Riemann argument in Sec. 4 wich led to only a single component for the vector potential A. The development of these potential arguments was fundamentally based not only on flat surface but also on two-dimensional ( $x$ and $z$ ) bohavior. Nevertheless, we argue in this section that wo can maintain the specific potential forms to find displacements and stresses on the boundary.
2. F1at surface arguments al so obscure the resulting wave shapes evaluated on the surface. If the $z$-dependent parts of the incident and scattered fields are exp(-iKz) and exp(iKz) respectively, setting $z=0$ makes them both equal one. However setting $z=h$ makes them complex conjugate pairs (for real K ), which are different wave shapes. This is actually true even in the scalar Kirchhoff approximation and is ignored. Note that this doesn't affect the wave shape of the scattered field since it's an integral over the surface values and is an ontgoing wave. But its surface approzimation arises from an incident (incoming) wave shape.
3. Factorization of any one of the three incident wave shapes al so skews the remaining wave shapes, and in general leads to overall coordinate dependent reflection coefficients as we
show.
4. Lowest order (i.e. two-dimensional flat surface) coupling occurs betweon $P$ - $S V$ - and SH-waves if the incident field is assumed to have a component out of the $x-z$ plane, or, equivalently, if it is a function of $y$. To discussed This in Sec. 13.

Here we develop the Kirchhoff approximation by retaining the potential formalism. The displacements and stresses are defined in terms of the potentials. $A$, and $v$ as

$$
\begin{align*}
& \mathbf{u}_{1}=\frac{\partial \phi}{\partial x}-\frac{\partial A}{\partial z},  \tag{15.3}\\
& \mathbf{u}_{2}=\nabla,  \tag{15.4}\\
& \mathbf{u}_{3}=\frac{\partial \phi}{\partial z}+\frac{\partial A}{\partial x},  \tag{15.5}\\
& \tau_{z_{1}}=\mu\left[\frac{\partial^{2} A}{\partial z^{2}}-\frac{\partial^{2} A}{\partial z^{2}}\right]+2 \mu \frac{\partial^{2} \phi}{\partial x^{2} z}  \tag{15.6}\\
& \tau_{z_{2}}=\mu \frac{\partial v}{\partial z}, \tag{15.7}
\end{align*}
$$

and

$$
\begin{equation*}
\tau_{z^{3}}=2 \mu\left[\frac{\partial^{2} \phi}{\partial z^{2}}+\frac{\partial^{2} A}{\partial z_{z}}\right]-\lambda{k_{L}^{2}}^{2} \cdot \tag{15.8}
\end{equation*}
$$

These follow from Sec. 4 .

## P-TAFB:

For $P$-wave incidence the $p$ and $A$ potentials aro found from (9.1)-(9.4). Substitute these forms into (15.3), (15.5), and (15.8), ovaluate the results on the surface, and factor out the incident wave. For example for $u_{2}$ we have that

$$
\begin{equation*}
u_{1}\left({\underset{\sim}{x}}_{s}\right)=i k_{x} \phi^{i n}\left(x_{s}\right)+i k_{x} \phi^{s c}\left(x_{s}\right)-i \mathbb{E}_{T} A\left(x_{s}\right) \text {, } \tag{15.9}
\end{equation*}
$$

which can be written as

$$
\begin{align*}
u_{1}\left({\underset{\sim}{x}}_{s}\right) & =i\left[\underline{k}_{x}+\mathbf{k}_{x} \phi^{s c}\left({\underset{\sim}{x}}_{s}\right) / \phi^{i n}\left({\underset{\sim}{x}}_{s}\right)-\underline{k}_{T} A\left({\underset{\sim}{x}}_{s}\right) / \phi^{i n}\left({\underset{\sim}{x}}_{s}\right)\right] \phi^{i n}\left(x_{s}\right) \\
& =i\left[\underline{k}_{x}\left(1+S_{1} u\right)-K_{T} S_{2}\right] \phi^{i n}\left(x_{s}\right) \tag{15.10}
\end{align*}
$$

Where

$$
\begin{equation*}
u=\exp \left(2 i \mathbb{K}_{1} h\right), \quad=\exp \left[i\left(\mathbb{K}_{T}+\mathbb{K}_{L}\right) h\right] \text {, } \tag{15.11}
\end{equation*}
$$

and $S_{1}$ and $S_{2}$ are the reflection coefficients defined in (9.32). If we further write the incident $P$-wave on the surface as

$$
\begin{equation*}
I_{1}=\phi^{i n}({\underset{I}{s}})=A_{0}^{P} \exp \left[i\left(k_{x} x-k_{L} h\right)\right], \tag{15.12}
\end{equation*}
$$

we canwrite (15.10) as

$$
\begin{equation*}
\mathfrak{a}_{1}\left(x_{\sim}\right)=R_{11} I_{1}, \tag{15.13}
\end{equation*}
$$

where

$$
\begin{equation*}
R_{11}=i\left[\mathbf{k}_{\mathbf{x}}\left(1+S_{1} n\right)-\mathbb{E}_{T^{n}} S_{2}\right] \tag{15.14}
\end{equation*}
$$

is the overall spatially dependent reflection coefficient from a pave incident field to the first component of displacement. We al so get that

$$
\begin{equation*}
u_{3}(\underset{\sim}{x})=R_{31} I_{1} . \tag{15.15}
\end{equation*}
$$

where

$$
\begin{equation*}
R_{31}=-i\left[\mathbb{X}_{L}\left(1-S_{1} u\right)-k_{2} w S_{2}\right] \text {. } \tag{15.16}
\end{equation*}
$$

Both displacement terms (15.13) and (15.15) are the first and third components of total displacement resulting from an incident $P$-wave.

The stress components from (15.6) and (15.8) evaluated on the sarface can be written as

$$
\begin{equation*}
T_{1}\left(x_{s}\right)=\tau_{z 1}\left(x_{\sim}\right)=T_{11} I_{1} . \tag{15.17}
\end{equation*}
$$

and

$$
\begin{equation*}
T_{3}\left({\underset{\sim}{x}}_{s}\right)=\tau_{z_{3}}\left({\underset{\sim}{x}}_{s}\right)=T_{31} I_{1}, \tag{15.18}
\end{equation*}
$$

where

$$
\begin{equation*}
T_{11}=\mu\left(\mathbf{K}_{T}^{2}-\mathbf{k}_{\mathbf{Z}}^{2}\right) S_{2} w+2 \mu \mathbf{k}_{\mathbf{X}} \mathbf{K}_{L}\left(1-S_{1} u\right) \tag{15.19}
\end{equation*}
$$

and

$$
\begin{equation*}
T_{32}=-\left(2 \mu \mathbf{E}_{L}^{2}+\lambda k_{L}^{2}\right)\left(1+S_{1} u\right)-2 \mu k_{z} \mathbb{K}_{T_{2}} S_{2} w \tag{15.20}
\end{equation*}
$$

Here $T_{11}$ and $T_{31}$ are the overall traction reflection coefficients which take us from an incident p-wave field to the first and third components of traction on the surface which are due to the $P$-wave incidence.

SV-WAVE8
For SV-wave incidence the $p$ and $A$ potentials are fond from (9.50)(9.53). If we define the incident field on the surface as

$$
\begin{equation*}
I_{3}=A^{i n}\left({\underset{\sim}{x}}_{s}\right)=A_{0}^{v} \exp \left[i\left(k_{x^{I}}-K_{T^{h}}\right)\right] . \tag{15.21}
\end{equation*}
$$

and

$$
\begin{equation*}
t=\exp \left(2 i \mathbf{K}_{\mathrm{T}} \mathrm{~h}\right) \tag{15.22}
\end{equation*}
$$

we can write the results asing (15.3), (15.5), (15.6) and (15.8) as

$$
\begin{align*}
& u_{1}(\underset{\sim}{x})=R_{13} I_{3},  \tag{15.23}\\
& u_{3}(\underset{\sim s}{x})=R_{33} I_{3},  \tag{15.24}\\
& T_{1}(\underset{\sim}{x})=T_{13} I_{3}, \tag{15.25}
\end{align*}
$$

and

$$
\begin{equation*}
T_{3}({\underset{\sim}{x}})=T_{33} I_{3}, \tag{15.26}
\end{equation*}
$$

where

$$
\begin{align*}
& R_{13}=i\left[X_{T}\left(1-S_{4} t\right)+E_{X^{W}} S_{3}\right] \text {, }  \tag{15.27}\\
& R_{33}=i\left[K_{X}\left(1+S_{4} t\right)+\mathbb{K}_{1} S_{3}\right] \text {. }  \tag{15.28}\\
& T_{13}=\mu\left(K_{T}^{2}-\mathbf{K}_{X}^{2}\right)\left(1+S_{4} t\right)-2 \mu K_{X} K_{L} w S_{3}, \tag{15.29}
\end{align*}
$$

and

$$
\begin{equation*}
T_{33}=2 \mu \mathbb{K}_{\mathbf{Z}} \mathbb{R}_{T}\left(1-S_{4} t\right)-\left(2 \mu \mathbb{R}_{L}^{2}+\lambda \mathbb{K}_{L}^{2}\right) \nabla S_{3}, \tag{15.30}
\end{equation*}
$$

in terms of the reflection coefficients $S_{3}$ and $S_{4}$ defined in (9.73).

## S日-TAVBS

For SH-incidence we define the incident SH-wave evaluated on the surface from (8.6) as

$$
\begin{equation*}
I_{2}=A_{0}^{H} e \operatorname{xp}\left[i\left(k_{x} x-k_{T} h\right)\right] . \tag{15.31}
\end{equation*}
$$

so that the components of displacement and stress are from (15.4) and (15.7) witten as

$$
\begin{equation*}
\mathrm{a}_{2}\left(\mathrm{x}_{s}\right)=\mathrm{R}_{22} \mathrm{I}_{2}, \tag{15.32}
\end{equation*}
$$

and

$$
\begin{equation*}
T_{2}\left(X_{\sim}\right)=T_{22} I_{2} . \tag{15.33}
\end{equation*}
$$

where

$$
\begin{equation*}
R_{22}=1+S_{12} t \tag{15.34}
\end{equation*}
$$

and

$$
\begin{equation*}
T_{22}=-\mu \mathbf{R}_{T}\left(1-S_{12} t\right) \tag{15.35}
\end{equation*}
$$

written in terms of the reflection coefficient $S_{11}$ defined by (8.22). All these results can be combined in the matrix representations as the Kirchhoff approximation for displacements

$$
\left[\begin{array}{l}
u_{1}\left(x_{\sim}\right)  \tag{15.36}\\
u_{2}\left(x_{3}\right) \\
n_{3}\left(x_{3}\right)
\end{array}\right]=\left[\begin{array}{lll}
R_{12} & 0 & R_{13} \\
0 & R_{22} & 0 \\
R_{31} & 0 & R_{33}
\end{array}\right]\left[\begin{array}{l}
I_{1} \\
I_{2} \\
I_{3}
\end{array}\right],
$$

and the Kirchhoff approximation for stresses or tractions

$$
\left[\begin{array}{l}
T_{1}\left({\underset{\sim}{x}}_{s}\right) \\
T_{2}\left({\underset{\sim}{x}}_{s}\right) \\
T_{3}\left({\underset{\sim}{x}}_{s}\right)
\end{array}\right]=\left[\begin{array}{lll}
T_{11} & 0 & T_{13} \\
0 & T_{22} & 0 \\
T_{31} & 0 & T_{33}
\end{array}\right]\left[\begin{array}{l}
I_{1} \\
I_{2} \\
I_{3}
\end{array}\right]
$$

An additional version more closely corresponding to the scalar case would have $\mathrm{w}=\mathrm{w}=\mathrm{t}=1$, in which case the $\mathrm{R}_{\mathrm{ij}}$ and $\mathrm{T}_{\mathrm{ij}}$ are spatially independent. Note al so that this version of the Eirchhoff approximation doesn't couple P- and SV-waves to SH-wave on the surface. The field values, however, are coupled through the integral relations.

APPRNDIX 3A. REGOLARIZATION OR TG ${ }^{0}$

$$
\mathrm{TG}^{0} \text { is defined by }
$$

$$
\begin{gather*}
{\left[T G^{0}(\underset{\sim}{x}-\underline{x})\right]_{p i j}=\mu \delta_{p m} \partial_{m} G_{i j}^{0}\left({\underset{x}{x}}^{\prime}-x\right)+\mu \delta_{p m} \partial_{i} G_{m j}^{0}\left(x_{\sim}^{\prime}-x\right)} \\
+\lambda \delta_{i p} \partial_{m} G_{m j}^{0}\left(x^{\prime}-\underline{z}\right) \tag{A.1}
\end{gather*}
$$

$\mathbf{G}^{0}$ is defined by

$$
\begin{equation*}
G_{m j}^{0}\left(x^{\prime}-\underline{\sim}\right)=\mu^{-1} \delta_{m j} G^{T}\left(x^{\prime}-\underline{\sim}\right)+\mathbb{K}^{-2} \partial_{m} \partial_{j}\left[G^{T}\left({\underset{\sim}{x}}^{\prime}-\underline{\sim}\right)-G^{L}\left({\underset{\sim}{x}}^{\prime}-\underline{\sim}\right)\right], \tag{A.2}
\end{equation*}
$$

Where GT, $\mathrm{L}_{\text {are }}$ the scalar free space Green's functions defined by (we choose the retarded Green's functions)

$$
\begin{equation*}
G^{T, L}\left(\underset{\sim}{x^{\prime}-\underset{\sim}{x}}\right)=\exp \left[i k_{T, L}\left|x^{\prime}-\underline{z}\right|\right] / 4 \pi\left|x^{\prime}-\underline{\sim}\right| . \tag{A.3}
\end{equation*}
$$

From (A.2) we can compate the third term in (A.1)

$$
\begin{align*}
\partial_{m} G_{m j}^{0}(\underset{\sim}{x}-\underline{x}) & =\mu^{-1} \partial_{j} G^{T}(\underset{\sim}{x}-\underset{\sim}{x})+\mathbb{K}^{-2} \partial_{j}\left[-k^{2} G^{T}(\underset{\sim}{x}-\underline{\sim})+k_{L}^{2} G^{L}\left({\underset{\sim}{x}}^{\prime}-\underline{x}\right)\right] \\
& =(\lambda+2 \mu)^{-1} \partial_{j} G^{L}\left(x_{\sim}^{\prime}-\underset{\sim}{x}\right) \tag{A.4}
\end{align*}
$$

where we have used the differential equation for the free-space Green's fanctions and the identity $k_{L}^{2}=\mathbb{K}^{2} /(\lambda+2 \mu) . \quad 0 \operatorname{sing}(A .4)$ and (A.2) in (A.1) and the identity $L_{T}^{2}=K^{2} / \mu$ we get

$$
\begin{align*}
{\left[T G^{0}\left(x^{\prime}-\underset{\sim}{x}\right)\right]_{p i j}=} & {\left[\delta_{i j} \partial_{p}+\delta_{p j} \partial_{i}\right] G^{T}\left(x_{\sim}^{\prime}-x\right) } \\
& +[\lambda /(\lambda+2 \mu)]^{\delta}{ }_{i p^{\prime}} \partial_{j} G^{L}\left(x^{\prime}-x\right)  \tag{A.S}\\
& +\left(2 / k_{t}^{2}\right) \partial_{p} \partial_{i} \partial_{j}\left[G^{T}\left(x^{\prime}-x\right)-G^{L}\left(x^{\prime}-\underset{\sim}{x}\right)\right] .
\end{align*}
$$

We have already regularized the first derivative of the free-space retarded Green's function in Ch. 1, Sec. 5. Here we are differentiating on the source coordinate and we can write for example

$$
\begin{equation*}
\partial_{j} G^{T}(\underset{\sim}{x}-x)=-\frac{1}{2} R_{j}^{T}\left(z_{\sim}^{\prime}-\underset{\sim}{x}\right)+\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right) \delta_{j 3} \delta\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right), \tag{A.6}
\end{equation*}
$$

where
and

$$
\begin{equation*}
P_{j}^{T}(\underset{\sim}{k})=2 i\left[k_{j t}+\delta_{j 3} P\left[\frac{\mathbb{R}_{T}^{2}}{\mathbf{k}_{z}}\right]\right], \tag{A.8}
\end{equation*}
$$

where $\mathbb{K}_{T}^{2}=k_{T}^{2}-k_{t}^{2}$ and $\tilde{G}^{T}(k)=\left[k^{2}-k_{T}^{2}\right]^{-1}$. The result for the longitudinal Green's function $G^{L}$ requires the replacement $\mathbb{k}_{T} \rightarrow k_{L}$. We can thin write (A.S) as

$$
\begin{align*}
& {\left[T G^{0}\left(\underline{x}^{\prime}-\underline{x}\right)\right]_{p i j}=-\frac{1}{2}\left[\delta_{i j} R_{p}^{T}\left(x^{\prime}-\underline{x}\right)+\delta_{p j} R_{j}^{T}(\underset{\sim}{x}-\underline{x})+\delta_{i p}(\lambda /(\lambda+2 \mu)) R_{j}^{L}\left(x^{\prime}-x\right)\right]} \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-x_{t}\right) \\
& {\left[\delta_{i j} \delta_{p 3}+\delta_{p j} \delta_{i 3}+\left[(\lambda /(\lambda+2 \mu)] \delta_{i p} \delta_{j} 3\right]\right.} \\
& +\frac{2}{k_{T}^{2}} \partial_{p} \partial_{i} \partial_{j}\left[G^{T}\left(x^{\prime}-\underline{x}\right)-G^{L}\left(x_{\sim}^{\prime}-x\right)\right] . \tag{A.9}
\end{align*}
$$

Thus (A.S) is partially regalarized in (A.9). It remains to regularize the triple derivative terms. One is sufficient, the other will follow with the replacement $\mathbf{k}_{\mathbf{T}} \leftrightarrow \mathbf{k}_{\mathrm{L}}$.

We begin with the Moyl representation

We take the triple derivative $\partial_{p} \partial_{i}{ }^{\partial} j_{j}$ of this. Break it up as

$$
\begin{align*}
\partial_{p} \partial_{i} \partial_{j}= & {\left[\partial_{p t}+\delta_{p 3} \partial_{z}\right]\left[\partial_{i t}+\delta_{i 3} \partial_{z}\right]\left[\partial_{j t}+\delta_{j 3} \partial_{z}\right]=} \\
= & \partial_{p t} \partial_{i t} \partial_{j t}+\left[\partial_{p t} \partial_{i t} \delta_{j 3}+\partial_{p t} \partial_{j t} \delta_{i 3}+\partial_{i t} \partial_{j t} \delta_{p 3}\right]_{z} \\
& +\left[\partial_{p t} \delta_{i z} \delta_{j 3}+\partial_{i t} \delta_{p 3} \delta_{j 3}+\partial_{j t} \delta_{p 3} \delta_{i 3}\right]_{z}^{\partial}+\delta_{i 3} \delta_{j 3} \delta_{p 3} \partial_{z}^{3} . \tag{A.11}
\end{align*}
$$

For the z-derivatives we have

$$
\begin{align*}
\partial_{z} e^{i K_{T}\left|z^{\prime}-z\right|} & =-i K_{T} s g n\left(z^{\prime}-z\right) e^{i K_{T}\left|z^{\prime}-z\right|}  \tag{A.12}\\
\partial_{z}^{z} e^{i K_{T}\left|z^{\prime}-z\right|} & =\left[2 i K_{T} \delta\left(z^{\prime}-z\right)-K_{T}^{2}\right] e^{i K_{T}\left|z^{\prime}-z\right|} \\
& =2 i K_{T} \delta\left(z^{\prime}-z\right)-K_{T}^{2} e^{i K_{T}\left|z^{\prime}-z\right|} \tag{A.13}
\end{align*}
$$

and

$$
\partial_{z}^{3} e^{i K_{T}\left|z^{\prime}-z\right|}=-2 i K_{X^{\prime}}^{\prime}\left(z^{\prime}-z\right\rangle+i \mathbb{Z}_{T}^{3} \operatorname{sgn}\left(z^{\prime}-z\right) e^{i K_{T}\left|z^{\prime}-z\right|} \text {. (A.14) }
$$

It is the z-derivative terms which detemine whether or not we have-a regular or singular representation. For example, (A.12) yields a singular result since it containg a single power of $K_{T}$. The first term in (A.13) and both terms in (A.14) also yield singular results.

We can thus write the full result as

$$
\begin{aligned}
& \partial_{p} \partial_{i}{ }_{j} G^{T}(\underset{\sim}{x}-x)
\end{aligned}
$$

$$
\begin{aligned}
& +\left[\partial_{p t}{ }^{\partial}{ }_{i t}{ }^{\delta}{ }_{j 3}+\partial_{p t} \partial_{j t} \delta_{i 3}+\partial_{i t}{ }^{\delta}{ }_{p z}\right]^{(-i) s g n\left(z^{\prime}-z\right)} . \\
& \cdot \frac{\pi i}{(2 \pi)^{3}} \iiint \underline{z}_{t} e^{i k_{t} \cdot\left[x_{t}^{\prime}-x_{t}\right]} e^{i \mathbb{R}_{T}\left|z^{\prime}-z\right|} \\
& +\left[\partial_{p t} \delta_{i 3}+\partial_{i t} \delta_{p^{3}} \delta_{j 3}+\partial_{j t^{\prime}} \delta_{i s^{3}} \delta_{p^{3}}\right]^{2 i} \delta\left(z^{\prime}-z\right) \cdot \\
& \cdot \frac{\pi i}{(2 \pi)^{3}} \iiint_{\sim_{t}} e^{i{\underset{\sim}{t}}_{t} \cdot\left[{\underset{x}{t}}_{\prime}-{\underset{\sim}{x}}_{t}\right]} \\
& +\left[{ }_{p t} \delta_{i 3} \delta_{j 3}+\partial_{i t} \delta_{p^{3}} \delta_{j 3}+\partial_{j t} \delta_{i 3} \delta_{p 3}\right]^{(-1)} . \\
& \cdot \frac{\pi i}{(2 \pi)^{3}} \iint d z_{t} \underline{k}_{T} e^{i z_{t} \cdot\left[z_{t}^{\prime}-z_{t}\right]} e^{i k_{T}\left|z^{\prime}-z\right|} \\
& +\delta_{i 3} \delta_{j}{ }^{3} \delta_{p^{3}}(-2 i) \delta^{\prime}\left(z^{\prime}-z\right) \cdot
\end{aligned}
$$

The first term in (A.15) is not singular and can be integrated up to a three-dimensional integral. The second term must be regularized. The third and fifth terms have the singularities directly. The fourth term is not singalar and the sixth term contains singularities which we recover.

$$
\begin{aligned}
& \partial_{p} \partial_{i} \partial_{j} G^{T}\left(x^{\prime}-x_{\sim}\right) \\
& =\frac{1}{[2 \pi]^{3}} \iiint d k e^{i k \cdot\left[x^{\prime}-\underset{\sim}{x}\right]} \tilde{G}^{T}(k)\left[i k_{p t} k_{i t} k_{j t}\right] \\
& +(-i) s g n\left(z^{\prime}-z\right)\left[\partial_{p t}{ }^{\partial}{ }_{i t}{ }^{\delta}{ }_{j}{ }^{3}+\partial_{p t}{ }^{\partial}{ }_{j t}{ }^{\delta}{ }_{i 3}+\partial_{i t}{ }^{\partial}{ }_{j t}{ }^{\delta}{ }_{p}{ }^{3}\right] \text {. }
\end{aligned}
$$

$$
\begin{align*}
& -\delta\left(z^{\prime}-z\right)\left[\delta_{i 3} \delta_{j} 3^{\partial}{ }_{p t}+\delta_{p^{3}} \delta_{j}{ }^{3} \partial_{i t}+\delta_{i 3} \delta_{p 3}{ }^{\partial}{ }_{j t}\right] \delta\left(x_{t}^{\prime}-x_{t}\right) \\
& +(-1)\left[\delta_{i 3} \delta_{j 3}{ }^{2}{ }_{p t}+\delta_{p^{3}} \delta_{j 3}{ }^{\partial}{ }_{i t}+\delta_{i 3^{3}} \delta_{p}{ }^{\partial}{ }_{j t}\right] . \\
& \frac{1}{(2 \pi)^{3}} \iiint d k e^{i k \cdot\left[z^{\prime}-\Sigma\right]} \tilde{G}^{T}(k) R_{T}^{2} \\
& +\delta_{i 3^{\prime}} \delta_{j 3^{3}} \delta^{3} \delta^{\prime}\left(z^{\prime}-z\right) \delta\left(z_{i}^{\prime}-z_{i t}\right) . \\
& +i \delta_{i 3} \delta_{j 3} \delta_{p^{3}} \operatorname{sgn}\left(z^{\prime}-z\right) . \\
& \text { - }\left[\frac{\pi i}{(2 \pi)^{3}} \iint \operatorname{dk}_{\sim} \mathbf{R}_{T}^{2} e^{i{\underset{\sim}{t}}_{t} \cdot\left[\underline{z}_{t}^{\prime}-z_{t}\right]}\left[e^{i K_{T}\left|z^{\prime}-z\right|}-1\right]\right. \\
& \left.+\frac{\pi i}{(2 \pi)^{3}} \iint d{\underset{\sim}{t}} E_{T}^{2} e^{i{\underset{k}{t}} \cdot\left[x_{t}^{\prime}-x_{t}\right]}\right], \tag{A.16}
\end{align*}
$$

where we have used the relation

$$
\begin{equation*}
K_{T} e^{i K_{T}\left|z^{\prime}-z\right|}=\frac{K_{T}^{2}}{\pi i} \int d k_{z} e^{i k_{z}\left(z^{\prime}-z\right)} \tilde{G}^{T}(\underline{k}) \tag{A.17}
\end{equation*}
$$

in the fourth term. Wo al so use

$$
\begin{equation*}
\operatorname{sgn}\left(z^{\prime}-z\right)\left[e^{i \mathbb{K}_{T}\left|z^{\prime}-z\right|_{-1}}\right]=\frac{1}{\pi i} \int d k_{z} e^{i k_{z}\left(z^{\prime}-z\right)} \tilde{G}^{T}(k) P\left[\frac{\mathbb{K}_{T}^{2}}{\mathbb{K}_{z}}\right] \tag{A.18}
\end{equation*}
$$

in the second and last terms in (A.16) to get

$$
\begin{aligned}
& \partial_{p} \partial_{i} \partial_{j} G^{T}\left(x^{\prime}-x\right) \\
& =\frac{1}{(2 \pi)^{3}} \iiint d{ }_{z} e^{i k \cdot\left[x^{\prime}-\underset{\sim}{z}\right]} \tilde{G}^{T}(k)\left[i k_{p t} k_{i t} k_{j t}\right] \\
& +(-i)\left[\delta_{j} \partial_{p t} \partial_{i t}+\delta_{i z} \partial_{p t} \partial_{j t}+\delta_{p z} \partial_{i t} \partial_{j t}\right] \text {. }
\end{aligned}
$$

$$
\begin{aligned}
& -\delta\left(z^{\prime}-z\right)\left[\delta_{i z} \delta_{j 3} \partial_{p t}+\delta_{p z} \delta_{j 3} \partial_{i t}+\delta_{i 3} \delta_{p z} \partial_{j t}\right] \delta\left({\underset{\sim}{i}}_{\prime}^{\prime}-{\underset{Z}{x}}\right) \\
& -\left[\delta_{i 3} \delta_{j 3} \partial_{p t}+\delta_{p^{2}} \delta_{j 3} \partial_{i t}+\delta_{i z} \delta_{p^{3}} \partial_{j t}\right] . \\
& \left.\cdot \frac{1}{(2 \pi)^{3}} \iiint d e^{i \underline{z} \cdot\left[z^{\prime}-x\right]}\right]_{G} T(k) R_{T}^{2} \\
& +\delta_{i 3} \delta_{j 3} \delta_{p} \delta^{\prime}\left(z^{\prime}-z\right) \delta\left(\underset{\sim}{x}-{\underset{\sim}{x}}_{t}\right)
\end{aligned}
$$

$$
\begin{align*}
& -\frac{1}{2} \delta_{i} \delta_{j} j_{j} \delta_{p} k_{T}^{2} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-x_{t}\right) \\
& -\frac{1}{2} \delta_{i} \delta_{j} \delta_{j} \delta_{p} \operatorname{sgn}\left(z^{\prime}-z\right) \partial_{t}^{2} \delta\left(x_{t}^{\prime}-x_{t}\right) \quad . \tag{A.19}
\end{align*}
$$

Carrying out the differentiations in the regular terms and combining the results we get

$$
\begin{aligned}
& \partial_{p} \partial_{i} \partial_{j} G^{T}\left(x^{\prime}-I_{i}\right) \\
& =\frac{1}{(2 \pi)^{3}} \iiint d k e^{i k \cdot\left[x^{\prime}-x\right]} \tilde{G}^{T}(k) . \\
& \text { - }\left[i k_{p t} k_{i t} k_{j t}+i\left[k_{p t} k_{i t} \delta_{j 3}+k_{p t} k_{j t} \delta_{i 3}+k_{i t} k_{j t} \delta_{p 3}\right] P\left[\frac{K_{T}^{2}}{k_{z}}\right]\right. \\
& +i\left[\delta_{i 3} \delta_{j 3} k_{p t}+\delta_{p^{3}} \delta_{j 3} k_{i t}+\delta_{i 3} \delta_{p^{3}} k_{j t}\right] \mathbb{R}_{T}^{2} \\
& \left.+i \delta_{i 3} \delta_{j 3} \delta_{p 3} K_{T}^{4} P\left[\frac{1}{k_{z}}\right]\right] \text {. }
\end{aligned}
$$

$$
\begin{aligned}
& -\delta\left(z^{\prime}-z\right)\left[\delta_{i 3} \delta_{j 3} \partial_{p t}+\delta_{p y^{3}} \delta_{j 3} \partial_{i t}+\delta_{i 3} \delta_{p ;} \partial_{i t}\right] \delta\left(x_{t}^{\prime}-x_{t}\right) \\
& +\delta_{i 3^{\prime}} \delta_{j 3^{\prime}} \delta_{p} \delta^{\prime}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-z_{t}\right) \\
& -\frac{1}{2} \delta_{i} \delta_{j} \delta_{j} \delta_{p} z^{z^{2}} T^{8 g n\left(z^{\prime}-z\right)} \delta\left(x_{t}^{\prime}-x_{t}\right)
\end{aligned}
$$

which can be revitten as

$$
\begin{aligned}
& \partial_{p} \partial_{i} \partial_{j} G^{T}\left(x^{\prime}-x\right) \\
& =\frac{1}{2} \frac{1}{(2 \pi)^{3}} \iiint d e^{i k \cdot\left[x^{\prime}-\underset{\sim}{x}\right]_{\tilde{G}} T}(k) P_{p i j}^{T}(k) \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right)\left[\delta_{j} \partial_{p t} \partial_{i t}+\delta_{i} \partial_{j t} \partial_{j t}+\delta_{p ;} \partial_{i t} \partial_{j t}\right] \delta\left(z_{t}^{\prime}-x_{t}\right) \\
& -\delta\left(z^{\prime}-z\right)\left[\delta_{i 3} \delta_{j 3} \partial_{p t}+\delta_{p 3} \delta_{j z} \partial_{i t}+\delta_{i 3} \delta_{p 3}{ }^{\partial}{ }_{j t}\right] \delta\left(x_{i}^{\prime}-x_{t}\right) \\
& +\delta_{i 3} \delta_{j}{ }^{\delta}{ }_{p}{ }^{3} \delta^{\prime}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-x_{t}\right)
\end{aligned}
$$

$$
\begin{align*}
& -\frac{1}{2} \delta_{j} \delta_{j z} \delta_{p z} \operatorname{sgn}\left(z^{\prime}-z\right) \partial_{t}^{2} \delta\left(x_{t}^{\prime}-x_{t}\right) \quad . \tag{A.20}
\end{align*}
$$

where the regular part is given by

$$
\begin{align*}
P_{p i j}^{T}(k)=2 i & {\left[k_{p t} k_{i t} k_{j t}\right.} \\
& +\left[\delta_{j 3} k_{p t} k_{i t}+\delta_{i z} k_{p t} k_{j t}+\delta_{p z} k_{i t} k_{j t}\right] p\left[\frac{\mathbb{K}_{T}^{2}}{k_{z}}\right] \\
& \left.+\left[\delta_{i z} \delta_{j 3} k_{p t}+\delta_{p 3} \delta_{j 3} k_{i t}+\delta_{i z} \delta_{p 3} k_{j t}\right]\right]_{T}^{2} \\
& \left.+\delta_{i z} \delta_{j z} \delta_{p z} \mathbb{E}_{T}^{4} p\left[\frac{1}{k_{z}}\right]\right] . \tag{A.21}
\end{align*}
$$

Alternatively we can write (A.20) as the sum of a regular part plus singular terms where we can combino soveral singular tems

$$
\begin{aligned}
& \partial_{p} \partial_{i} \partial_{j} G^{T}\left(x_{\sim}^{\prime}-x\right) \\
& =\frac{1}{2} R_{p i j}^{T}\left(x^{\prime}-x\right) \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right)\left[\delta_{j 3} \partial_{p t} \partial_{i t}+\delta_{i 3} \partial_{p t} \partial_{j t}+\delta_{p z} \partial_{i t}{ }^{\partial}{ }_{j t}-\delta_{i 3} \delta_{j 3} \delta_{p 3} \partial_{t}^{2}\right] \delta\left(z_{t}^{\prime}-z_{t}\right) \\
& -\left[\delta_{i 3} \delta_{j 3} \partial_{p t}+\delta_{p 3} \delta_{j 3} \partial_{i t}+\delta_{i 3} \delta_{p 3} \partial_{j t}+\delta_{i 3} \delta_{j 3} \delta_{p 3} \partial_{z}\right] \delta\left(z_{i}^{\prime}-z\right)
\end{aligned}
$$

where the regalar part is given by

As a check, we note that if we set $p=i$ in (A.22) and sum so that we have

$$
\begin{align*}
\partial_{p} \partial_{p} \partial_{j} G^{T}\left(x^{\prime}-\underline{\sim}\right)= & \frac{1}{2} \mathbb{R}_{p p j}^{T}\left({\underset{\sim}{x}}^{\prime}-\underline{x}\right)-\partial_{j} \delta\left({\underset{\sim}{x}}^{\prime}-\underset{\sim}{x}\right) \\
& -\frac{1}{2} k_{T}^{2} \delta_{j 3} \operatorname{sgn}\left(x^{\prime}-x\right) \delta\left({\underset{\sim}{x}}^{\prime}-\underline{x}_{t}\right), \tag{A.24}
\end{align*}
$$

and we use the result

$$
\begin{align*}
& =k^{2} P_{j}^{T}(\underline{i}) \quad . \tag{A.25}
\end{align*}
$$

(where $P_{j}^{T}(k)$ was defined in $(5.18)$ of Ch. 1 , so that

$$
\begin{equation*}
R_{p p j}^{T}\left(x^{\prime}-x\right)=k_{T_{j}}^{2} R_{j}^{T}\left(x^{\prime}-x\right), \tag{A.26}
\end{equation*}
$$

and we get the regularization of the first derivative of $G^{T}$ as defined in Ch. 1. For example, we could write, using the results in Ch. 1,

$$
\begin{aligned}
& \partial_{p} \partial_{p} \partial_{j} G^{T}\left(x^{\prime}-x\right)=\partial_{j} \partial_{p} \partial_{p} G^{T}\left(x^{\prime}-\underline{x}\right) \\
& =\partial_{j}\left[-k_{G^{2}}^{T}\left(x^{\prime}-x\right)-\delta\left(x^{\prime}-x\right)\right] \\
& =-E^{2}\left[-\frac{1}{2} R_{j}^{T}\left(x^{\prime}-x_{\sim}\right)+\frac{1}{2} \delta_{j s} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left({\underset{\sim}{x}}^{\prime}-x_{t}\right)\right] \\
& -\partial_{j} \delta\left(x^{\prime}-x\right) \\
& =\frac{1}{2} k_{i}^{2} R_{j}^{T}\left(x^{\prime}-\underline{\sim}\right)-\partial j \delta\left(x_{j}^{\prime}-x\right) \\
& -\frac{1}{2} \varepsilon_{T}^{2} \delta_{j 3} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left({\underset{\sim}{x}}_{t}^{\prime}-{\underset{\sim}{x}}_{t}\right) \quad .
\end{aligned}
$$

which is the same as (A.24) using (A.26).

Osing (A.22) we can thas write

$$
\begin{aligned}
& \partial_{p} \partial_{i} \partial_{j}\left[G^{T}\left(x^{\prime}-x\right)-G^{L}\left(x^{\prime}-\underset{\sim}{x}\right)\right] \\
& =\frac{1}{2}\left[\mathbb{R}_{p i j}^{T}\left(x^{\prime}-x\right)-R_{p i j}^{L}\left(x^{\prime}-x\right)\right]
\end{aligned}
$$

Where $\mathbb{R}_{\mathrm{pij}}^{\mathrm{L}}$ is analogous to ( A .23 ) with $\tilde{G}^{T}$ and $\mathrm{P}_{\mathrm{pij}}^{\mathrm{T}}$ replaced by $\tilde{G}^{L}$ and $\mathrm{P}_{\mathrm{p} i j}^{\mathrm{L}}$.

The latter follows from (A.21) by replacing $\mathbf{R}_{T}$ by $\mathbb{K}_{L}=\left(k_{L}^{2}-k_{t}^{2}\right)^{2 / 2}$. We thus have from (A.9) and (A.27) that

$$
\begin{align*}
& {\left[T G^{0}\left(x^{\prime}-x\right)\right]_{p i j}=K_{p i j}\left(x^{\prime}-z^{\prime}\right)} \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-z_{t}\right) \cdot \\
& \cdot\left[\delta_{i j} \delta_{p{ }^{3}}+\delta_{p j} \delta_{i 3}+[\lambda /(\lambda+\mu)] \delta_{i p} \delta_{j}{ }^{3}\right. \\
& \left.+2\left(E_{L}^{2}-E_{T}^{2}\right) L_{T}^{-2} \delta_{i 3} \delta_{j 3} \delta_{p 3}\right] \text {, } \tag{A.28}
\end{align*}
$$

where

$$
\begin{align*}
& K_{p i j}\left(x^{\prime}-x\right)=K_{T}^{-2}\left[R_{p i j}^{T}\left(x^{\prime}-x_{i}\right)-R_{p i j}^{L}\left(x^{\prime}-x\right)\right] \\
& -\frac{1}{2}\left[\delta_{i j} R_{p}^{T}\left(x^{\prime}-x\right)+\delta_{p j} R_{i}^{T}\left({\underset{\sim}{x}}^{\prime}-x\right)\right. \\
& +\delta_{i p}\left[(\lambda /(\lambda+\mu)] \mathbb{R}_{j}^{L}\left(x^{\prime}-x\right)\right] \quad . \tag{A.29}
\end{align*}
$$

Osing the definitions of $k$ and $k_{T}$ we can rewrite the singnlar part of (A.28) to yield

$$
\begin{align*}
{\left[T G^{0}\left(x^{\prime}-x^{\prime}\right)\right]_{p i j}=} & K_{p i j}\left(x^{\prime}-x_{\sim}\right) \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-x_{t}\right) \cdot \\
& \cdot\left[\delta_{i j} \delta_{p^{3}}+\delta_{p j} \delta_{i ;}+[\lambda / \lambda+2 \mu] \delta_{i p^{3}} \delta_{j}\right. \\
& \left.-2[(\lambda+\mu) /(\lambda+2 \mu)] \delta_{i 3} \delta_{j 3} \delta_{p}{ }^{3}\right] \tag{A.30}
\end{align*}
$$

In the integral equation it is (A.30) dotted into the normal vector $n_{p}\left({\underset{\sim}{x}}_{t}\right)=\delta_{p^{3}}-\partial_{p t} h\left({\underset{\sim}{x}}^{\prime}\right)$ which is important. The result from (A.30) is

$$
\begin{align*}
& n_{p}\left(x_{t}\right)\left[T G^{0}\left(x^{\prime}-x_{i}\right)\right]_{p i j} \\
& =K_{j i}\left({\underset{\sim}{n}}^{\prime}, \underset{\sim}{x}\right)+ \\
& +\frac{1}{2} \operatorname{sgn}\left(z^{\prime}-z\right) \delta\left(x_{t}^{\prime}-x_{t}\right) . \\
& \text { - }\left[\delta_{i j}-\delta_{i z} \partial_{j t} h\left(\underline{x}_{t}\right)-\left[\frac{\lambda}{\lambda+2 \mu}\right] \delta_{j 3} \delta_{i t} h\left({\underset{\sim}{x}}_{t}\right)\right] \text {, } \tag{A.31}
\end{align*}
$$

Where

$$
\begin{equation*}
\mathbf{K}_{j i}\left({\underset{\sim}{x}}^{\prime}, \underline{\sim}\right)=\underline{n}_{p}\left({\underset{\sim}{x}}_{t}\right) \mathbf{K}_{p i j}\left(x_{\sim}^{\prime}-\underline{\sim}\right), \tag{A.32}
\end{equation*}
$$

which is not a function of the difference of coordinates.

APPENDIX 3B. DERIVATION OR THR FRER-SPACE BLASTIC GRREN'S FONCTION
a. Coordinate-space derivation

The free-space elastic Green's function satisfies the differential equation

$$
\begin{equation*}
\left[\Delta G^{0}\left(x, x^{\prime}\right)\right]_{i j}+\mathbb{x}^{2} G_{i j}^{0}\left(x, x^{\prime}\right)=-\delta_{i j} \delta\left(x-x^{\prime}\right) \tag{B.1}
\end{equation*}
$$

or, explicitly writing the operator $\Delta^{( }$(from (1.17))

$$
\begin{equation*}
\mu \nabla^{2} G_{i j}^{0}+(\lambda+\mu) \partial_{i} \partial_{m} G_{m j}^{0}+K^{2} G_{i j}^{0}=-\delta_{i j} \delta\left(z-z^{\prime}\right) \tag{B.2}
\end{equation*}
$$

Where we have suppressed the coordinate dependence. The scalar free-space Green's functions satisfy the equations

$$
\begin{equation*}
\left(\vartheta^{2}+\underline{k}_{L}^{2}\right) G^{L}=-\delta\left(\underset{\sim}{x}-\underline{z}^{\prime}\right) \tag{B.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(\nabla^{2}+\underline{1}_{T}^{2}\right) G^{I}=-\delta\left(\underset{\sim}{x}-x^{\prime}\right), \tag{B.4}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathbf{k}_{\mathrm{L}}^{2}=\mathbf{R}^{2} /(\lambda+2 \mu), \quad \mathbf{k}_{\mathrm{T}}^{2}=\mathbf{R}^{2} / \mu \tag{B.5}
\end{equation*}
$$

If the divergence term in (B.2) were absent we could solve the equation as a scalar equation, so we first solve for it by taking the divergence $\partial_{i}$ of (B.2) to get

$$
\begin{equation*}
(\lambda+2 \mu) \forall^{2}\left(\partial_{i} G_{i j}^{0}\right)+\mathbf{K}^{2}\left(\partial_{i} G_{i j}^{0}\right)=-\partial_{j} \delta\left(\underset{\sim}{x}-\mathbb{Z}^{\prime}\right) \tag{B.6}
\end{equation*}
$$

which can be rewritten as

$$
\begin{equation*}
\forall^{2}\left(\partial_{i} G_{i j}^{0}\right)+\mathbf{k}_{L}^{2}\left(\partial_{i} G_{i j}^{0}\right)=-\frac{\mathbf{k}^{2}}{\mathbf{K}^{2}} \partial_{j} \delta\left(\underset{\sim}{x-x^{\prime}}\right) \tag{B.7}
\end{equation*}
$$

This has the solution

$$
\begin{equation*}
\partial_{i} G_{i j}^{0}=\frac{\mathbf{k}_{L}^{2}}{\mathbf{K}^{2}} \partial_{j} G^{L} \tag{B.8}
\end{equation*}
$$

which follows from (B.3). Hence (B.2) can be rewritten as

$$
\begin{equation*}
\forall^{2} G_{i j}^{0}+k_{T}^{2} G_{i j}^{0}=-\frac{1}{\mu} \delta_{i j} \delta\left(\underset{\sim}{x}-x^{\prime}\right)-\frac{(\lambda+\mu)}{\mu} \frac{k_{L}^{2}}{K^{2}} \partial_{i} \partial_{j} G^{L} \tag{B.9}
\end{equation*}
$$

From (B.4) we note that

$$
\begin{equation*}
\left(\psi^{2}+k_{T}^{2}\right)\left[\frac{1}{\mu} \delta_{i j} G^{T}\right]=-\frac{1}{\mu} \delta_{i j} \delta\left(x^{T}-x^{\prime}\right) \tag{B.10}
\end{equation*}
$$

Subtract (B.10) from (B.9) to get

$$
\begin{equation*}
\left(\nabla^{2}+k_{T}^{2}\right)\left[G_{i j}^{0}-\frac{1}{\mu} \delta_{i j} G^{T}\right]=-\frac{\lambda+\mu \underline{k}_{L}^{2}}{\mu} \frac{K^{2}}{Z_{i} \partial_{j} G^{L}} \tag{B.11}
\end{equation*}
$$

Next rewrite the rhs of (B.11) using

$$
\begin{align*}
\frac{\lambda+\mu}{\mu} & =\frac{\lambda+2 \mu-\mu}{\mu} \\
& =\left(k_{T}^{2} / k_{L}^{2}-1\right) \tag{B.12}
\end{align*}
$$

so that (B.11) becomes

$$
\begin{equation*}
\left(\nabla^{2}+k_{T}^{2}\right)\left[G_{i j}^{0}-\frac{1}{\mu} \delta_{i j} G^{T}\right]=-\frac{k_{T}^{2}}{k^{2}} \partial_{i} \partial_{j} G^{L}+\frac{k_{L}^{2}}{k^{2}} \partial_{i} \partial_{j} G^{L} \tag{B.13}
\end{equation*}
$$

For the second term on the rhs of (B.13) we substitute from (B.3)

$$
\begin{equation*}
\mathbf{k}_{L}^{2} G^{L}=-v^{2} G^{L}-\delta\left(\underset{\sim}{x-z^{\prime}}\right), \tag{B.14}
\end{equation*}
$$

so that (B.13) becomes

$$
\begin{equation*}
\left(\nabla^{2}+k^{2}\right)\left[G_{i j}^{0}-\frac{1}{\mu} \delta_{i j} G^{T}+\frac{1}{K^{2}} \partial_{i} \partial_{j} G^{L}\right]=-\frac{1}{K^{2}} \partial_{i} \partial_{j} \delta\left(\underset{\sim}{x-z^{\prime}}\right) \tag{B.15}
\end{equation*}
$$

whose solution is

$$
\begin{equation*}
G_{i j}^{0}-\frac{1}{\mu} \delta_{i j} G_{i j}^{T}+\frac{1}{K^{2}} \partial_{i} \partial_{j} G^{L}=\frac{1}{K^{2}} \partial_{i} \partial_{j} G^{T} \tag{B.16}
\end{equation*}
$$

which is just (2.2) quoted in the text.
b. Fourier transform-space derivation

The coefficients in (B. 2) are constant so we introdace the Fourier transform

$$
\begin{equation*}
\tilde{G}_{i j}^{0}(\underline{\sim})=\iiint G_{i j}^{0}\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right) \text { exp }\left[-i \underset{\sim}{a} \cdot\left(\underset{\sim}{x}-{\underset{\sim}{x}}^{\prime}\right)\right] d \underset{\sim}{x} \tag{B.17}
\end{equation*}
$$

where we one the homogeneity of the Green's functions. The transform of (B.2) is thus

$$
\begin{equation*}
\mu a^{2} \tilde{G}_{i j}^{0}+(\lambda+\mu) a_{i} \alpha_{m} \tilde{G}_{m j}^{0}-\mathbb{R}^{2} \tilde{G}_{i j}^{0}=\delta_{i j} . \tag{B.18}
\end{equation*}
$$

Again we must first solve for the divergence term so maltiply (B.18) by $a_{i}$ so that the solution for

$$
\begin{equation*}
g_{j}=a_{m} \tilde{G}_{m j}^{0} \tag{B.19}
\end{equation*}
$$

is given by

$$
\begin{equation*}
g_{j}=\left[(\lambda+2 \mu) a^{2}-K^{2}\right]^{-1} a_{j} \tag{B.20}
\end{equation*}
$$

Substituting (B.20) into the second term on the lhs of (B.18) yields

$$
\begin{equation*}
\left(\mu a^{2}-\mathbb{K}^{2}\right) \tilde{G}_{i j}^{0}=\delta_{i j}-(\lambda+\mu)\left[(\lambda+2 \mu) a^{2}-\mathbf{K}^{2}\right]^{-1} a_{i} a_{j} \tag{B.21}
\end{equation*}
$$

This can be written using (B.5) and partial fractions as

$$
\begin{equation*}
\tilde{G}_{i j}^{0}(a)=\delta_{i j} \mu^{-1} \tilde{\mathbf{G}}^{\mathrm{T}}(\alpha)+a_{i} a_{j} \mathbf{K}^{-2}\left[\tilde{G}^{L}(a)-\tilde{G}^{T}(\alpha)\right] \tag{B.22}
\end{equation*}
$$

where

$$
\begin{equation*}
\tilde{G}^{\mathrm{T}, \mathrm{~L}}(a)=\left(a^{2}-\mathrm{E}_{\mathrm{T}, \mathrm{~L}}^{2}\right)^{-1} \tag{B.23}
\end{equation*}
$$

It is easily seen that the Fourier invorse of ( B .22 ) is $\mathrm{just}(2.2)$ in the text.

## APPENDIX 3 C. DIFEERENTIAL RELATIONS FOR INBOMOGBNEOOS MEDIA

In Section 3 we derived differential relations for the Green's functions and surface displacements for homogeneons media. We specified the resulting integral equations to surfaces and this illastrated the coupling of displacement components due to surface variability. Here we derive these differential relations for inhomogeneors media and this illustrates the coupling of displacement components due to volume variability. We begin with the equations of motion of the stress tensor, (1.12), witten as

$$
\begin{equation*}
\partial_{k} \tau_{j k}(\underline{x})+\mathbb{R}_{u_{j}}(\underset{\sim}{x})=0_{j} \tag{C.1}
\end{equation*}
$$

The stress is given by

$$
\begin{equation*}
\tau_{j k}(\underline{\sim})=\mu\left(\partial_{j} u_{k}+\partial_{k} u_{j}\right)+\lambda \delta_{j k} \partial_{m}^{u} u_{m} \tag{C.2}
\end{equation*}
$$

Where now $\mu$ and $\lambda$ can be spatially dependent. We define the traction operator T as

$$
\begin{equation*}
{ }_{j k}=T_{j k p}{ }_{p}^{u} \tag{C.3}
\end{equation*}
$$

so that it is explicitly given by

$$
\begin{equation*}
T_{j k p}=\mu \delta_{j p} \partial_{k}+\mu \delta_{k p} \partial_{j}+\lambda \delta_{j k} \partial_{p} \tag{C.4}
\end{equation*}
$$

For the scattered displacement components ( $C .1$ ) can be witten using (C. 3 ) as

$$
\begin{equation*}
\partial_{k} T_{j k p} \mathbf{u}_{p}^{s c}(\underset{\sim}{x})+K_{\mathbf{u}_{j}}^{s c}(\underset{\sim}{x})=0 \tag{C.5}
\end{equation*}
$$

The Green's function $G$ for this equation is given by the solution of

$$
\begin{equation*}
\partial_{k} T_{j k p} G_{p r}\left(x, x^{\prime}\right)+\mathbb{K}^{2} G_{j r}\left(\underset{\sim}{x}, x^{\prime}\right)=-\delta_{j r} \delta(R) \tag{C.6}
\end{equation*}
$$

where

$$
\underset{\sim}{\mathbf{R}}=\underset{\sim}{\underline{I}}-\mathbf{I}^{\prime} ;{\underset{\sim}{x}}^{\prime}=\text { source position }
$$

and where we have the same operator $T$, $i$. . the same inhomogeneous $\lambda$ and $\mu$ as in (C.5).

Next, cross multiply the solutions, i.e. form the quantity

$$
\begin{equation*}
G_{i j}\left(\underset{\sim}{x},{\underset{\sim}{x}}^{\prime}\right)\left[\partial_{k} T_{i k p} n_{p}^{s c}(x)\right]-\left[\partial_{k} T_{i k p} G_{p j}\left(x, x^{\prime}\right)\right] n_{i}^{s c} \tag{C.7}
\end{equation*}
$$

Substituting in (C. 5) and (C.6) as appropriate we can derive the relation

$$
\begin{align*}
n_{j}^{s c}(\underset{\sim}{x}) \delta(\underset{\sim}{R})= & \partial_{k}\left[G_{i j} T_{i k p}{ }^{u_{p}^{s c}}-\left(T_{i k p} G_{p j}\right) n_{i}^{s c}\right] \\
& +\left(T_{i k p} G_{p j}\right) \partial_{k} n_{i}^{s c}-\left(\partial_{k} G_{i j}\right)\left(T_{i k p} n_{p}^{s c}\right) \tag{C.8}
\end{align*}
$$

Explicit evaluation of the latter two terms in (C.8) shows that they vanish. The result is that if the equations for $G$ and $a$ contain the same $\lambda$ and $\mu$ w get a pare divergence

$$
\begin{equation*}
n_{j}^{s c}(\underset{\sim}{x}) \delta(\underset{\sim}{R})=\partial_{k}\left[G_{i j} T_{i k p} n_{p}^{s c}-\left(T_{i k p} G_{p j}\right) n_{i}^{s c}\right] \tag{C.9}
\end{equation*}
$$

Which is our first differential relation.
To derive the second differential relation we start with the equation for the homogeneous Green's function (2.1) written in terms of the traction operator $T^{0}$

$$
\begin{equation*}
\partial_{k} I_{j k p}^{0} G_{p r}^{0}+\left(\mathbb{K}^{0}\right)^{2} G_{j r}^{0}\left(\underset{\sim}{x}-\underline{z}^{\prime}\right)=-\delta r^{\delta(R)}, \tag{C.10}
\end{equation*}
$$

where $T^{0}$ is given by

$$
\begin{equation*}
T_{j k p}^{0}=\mu^{0} \delta_{j p} \partial_{k}+\mu^{0} \delta_{k p} \partial_{j}+\lambda^{0} \delta_{j k} \partial_{p} \tag{C.11}
\end{equation*}
$$

and we explicitly note that we have a constant backgronndmedim, i.e. where $\mu^{0}$, $\lambda 0$, and $\rho^{0}$ are constant. Next, cross maltiply the solutions of (C.10) and (C.S) to form the quantity

$$
\begin{equation*}
G_{i j}^{0}\left(\partial_{k} T_{i k p}{ }_{p}^{s c}\right)-\partial_{k}\left[T_{i k p}^{0} G_{p j}^{0}\right] \mathbf{u}_{i}^{s c} \tag{C.12}
\end{equation*}
$$

Substituting the solution forms we get the result

$$
\begin{align*}
u_{j}^{s c}(\Sigma) \delta(R)=\partial_{k}[ & \left.G_{i j}^{0} T{ }_{i k p}{ }_{p}^{s c}-\left[T_{i k p}^{0} G_{p j}^{0}\right] n_{i}^{s c}\right] \\
& +\left[K^{2}-\left(K^{0}\right)^{2}\right] G_{i j}^{0} u_{i}^{s c} \\
& +\left(T_{i k p}^{0} G_{p j}^{0}\right) \partial_{k} u_{i}^{s c}-\left(\partial_{k} G_{i j}^{0}\right)\left(T_{i k p} u_{p}^{s c}\right) \tag{C.13}
\end{align*}
$$

Evaluation of the latter three terms in (C.13) yields the second differential relation

$$
\begin{align*}
& \mathrm{n}_{\mathrm{j}}^{\mathrm{sc}}(\mathrm{x}) \delta(\mathrm{R})=\partial_{k}\left[\mathrm{G}_{\mathrm{ij}}^{0} \mathrm{~T}_{\mathrm{ikp}} \mathrm{n}_{\mathrm{p}}^{\mathrm{sc}}-\left(\mathrm{I}_{\mathrm{ikp}}^{0} \mathrm{G}_{\mathrm{pj}}^{0}\right) \mathrm{n}_{\mathrm{i}}^{\mathrm{sc}}\right] \\
& +\omega^{2}\left(p-p^{0}\right) G_{i j}^{0}{ }^{\text {n }}{ }_{i}^{c c} \\
& -\left(\mu-\mu^{0}\right)\left(\partial_{k} G_{i j}^{0}\right)\left(\partial_{k} \mathbf{u}_{i}^{s c}+\partial_{i} \mathbf{u}_{k}^{s c}\right) \\
& -\left(\lambda-\lambda^{0}\right)\left(\lambda^{0}+2 \mu^{0}\right)^{-1}\left(\partial_{j} G^{L}\right)\left(\partial_{p} \mathbf{n}_{p}^{s c}\right) \quad . \tag{C.14}
\end{align*}
$$

In addition to a divergence term which, when integrated, yields surface
integrals, we al so get pare volume terms proportional to the differences
between the homogeneons (constant background) parameters and the
inhomogeneons ones.

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